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A METHOD FOR DETERMINING

THE HIGH ENERGY PHOTON SPECTRUM OF A

PULSED PLASMA SOURCE

THESIS

Charles W. Beason
Captain,   USAF

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PULSED PLASMA SOURCE

THESIS

Presented to the Faculty of the School of Engineering

of the Air Force Institute of Technology

Air University

In Partial Fulfullment of the

Requirements for the Degree of

Master of Science in Nuclear Engineering

Charles W. Beason, B.S.

Captain,   USAF

March 1984

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Acknowledgments

This thesis is dedicated to the memory of my grandfather,

William Erskine Beason

March 4 1900 – September 29 1983.

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Charles W. Beason

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Abstract

This investigation examined the feasibility of using an array of plastic scintillator photomultiplier tube radiation detectors to determine the high energy ($h\nu > 30$ keV) photon spectrum of SHIVA STAR. A method of determining the spectrum is outlined in which the detectors are filtered with different materials, and the spectrum is deconvoluted by an iterative technique on a computer. Data from SHIVA STAR was not available, so measurements of the radiation output of a plasma gun were analyzed.

List of Tables

I. Introduction

Background

The problem of deconvoluting an unknown spectrum from the responses of differently filtered identical detectors is not a new one. Early work by Ross (14:425; 15:433) involved the use of filters made from two elements of consecutive atomic number. The difference in the energies of the K-edges of the two filtering elements were used to determine the amount of photon energy lying between the K-edges. This technique was later expanded and greatly refined by Kirkpatrick (9:186; 10:223).

The method used by Ross and later by Kirkpatrick allowed the determination of the photon energy in only a narrow region of the electromagnetic spectrum. Silberstein (16:375) reported a method of determining a continuous spectrum, which has been called the Laplace Transform method. However, the drawback in using this method is that the maximum energy of the spectrum needs to be known (19:529).

The present technique used follows that used by Twidell(19:529-539), Tominaga (17:415-421), and exactly that of Degnan (3:264-269). The technique involves performing the deconvolution of the unknown spectrum by an iterative technique on a computer. The deconvolution method used in the computer program is described in the chapter on theory, Chapter II, and is listed in Appendix A.

The spectrum to be determined was to have been the high energy (h𝜈 > 20 keV) photon spectrum of SHIVA STAR. SHIVA STAR

is a large (10 MJ), fast (quarter cycle time of 5 μ sec), capacitor

bank used to implode deuterated formvar / aluminum cylindrical

liners in a z-pinch at the Simulators and Advanced Weapons

Concepts Branch of the Air Force Weapons Laboratory (AFWL/NTYP).

Predecessors to SHIVA STAR, such as SHIVA II produced plasmas

which radiated several hundred kilojoules of energy in the x-ray

region of the spectrum over a period of about 80 nsec. The

electron temperature of the plasmas produced were about 300-

400 eV in the hottest regions of the pinch (3:267). Fusion of the

deuterium in the liners produced on the order of 10 exp(8)

neutrons that were detected by silver activation counters and

time-of-flight (TOF) detectors (3:268). The TOF detectors were

filtered by 2 in. and 3/8 in. of lead. The difference in signal

from the two differently shielded TOF detectors can not be

accounted for by assuming that only neutrons are producing the

signal. From the signals seen, up to a few joules of high energy

photons (Hν > 20 keV) had to have been radiated.

The availability of SHIVA STAR for acquiring experimental

data could not be guaranteed, so measurements were taken of the

spectrum of the PUFF bank driven coaxial plasma gun at AFWL/NTYP.

The PUFF bank is a 500 kJ fast capacitor bank which had a coaxial

plasma gun mounted on it. The PUFF bank and the plasma gun are

described in the chapter on equipment, Chapter III.

Problem

The problem presented in this thesis is that of devising a method to determine the high energy photon spectrum (hν > 3 0 keV) of SHIVA STAR. The method is to be developed using data acquired from a coaxial plasma gun mounted on the PUFF capacitor bank at AFWL/NTYP. The radiation detectors to be used are eight NPM54X plastic scintillator/photomultiplier tube detectors manufactured by EG&G. Filters for the detectors are to be chosen based on estimates of the spectrum. The computer program to be used to deconvolute the spectrum will be a modification of the program XSPEC.

Scope

The scope of this thesis is to be limited to five areas. They are: filter selection, detector calibration, modification of the program XSPEC, spectrum deconvolution, and revised filter selection.

The first area considered is that of filter selection. A resonable estimate of the expected high energy photon output is first to be made. Filters are then to be chosen based upon what types of filters are available and what combinations of these filters produce a reasonable result from the spectrum deconvolution program.

The detectors are to be calibrated using various radioactive sources that are available. The calibration is to be relative with respect to the photon energies of the sources. An absolute calibration is to be made for each detector from the comparison

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of ^{60}Co calibration data provided by the manufacturer of the detectors.

The program XSPEC is to be modified so that it will accept data in the form of time versus detector voltages. The program is to output a spectrum giving photon group energies and energy amounts.

Data from the coaxial plasma gun is to be deconvoluted using the modified program. A spectrum and error estimates will be produced for each set of experimental data deconvoluted.

A revised set of filters based upon the deconvoluted spectrum will be determined.

Assumptions

Several assumptions have to be made concerning this thesis. They are:

1) The ^{60}Co calibration data for the NPM54X detectors provided by EG&G is correct to the errors specified;

2) The method of deconvolution used by the computer program XSPEC is valid.

Sequence Of Presentation

The sequence of presentation of material in this thesis will be as follows:

1) The theories behind the operation of a coaxial plasma gun, the operation of a z-pinch, the radiation detectors, and the method of spectrum deconvolution will be presented in Chapter II;

2) The equipment used will be listed and described in Chapter III;

3) The procedures followed in detector calibration, filter selection, data acquisition, spectrum deconvolution, and new filter selection will be presented in Chapter IV;

4) The experimental results will be presented and described in Chapter V;

5) Conclusions and a discussion of the experimental results and their implications will be presented in Chapter VI.

II. Theory

The Plasma Gun

The plasma gun presently used is similar in design and operation to that reported by Marshall (13:134-135). The plasma gun, depicted in Figure 1, is coaxial and is driven by a fast capacitor bank. The capacitor bank drives a high current through a gas that is puffed into the gun at the base of the gun. The current ionizes the puff gas, transforming it into a highly conductive plasma. A magnetic field, $\vec{B}$, produced by the current, $\vec{j}$, acts by the $\vec{j} \times \vec{B}$ force to move the current (and thus the plasma) up and out the barrel of the plasma gun at velocities that exceed $10 \exp(7)$ cm/sec (13:134; 1:212).

The $\vec{j} \times \vec{B}$ force acting upon the gas puff can be derived from either Maxwell's equations and the Lorentz force, or by **considering the energy density of the self magnetic field as a** pressure pushing on the plasma. Figure 2 shows the direction of the current in the plasma, which from the right-hand rule gives the direction of the magnetic field produced by the current in the plasma. Ampere's law gives the conditions requiring the existance of the magnetic field only in the electrode gap beneath the plasma.

As the plasma leaves the gun, two different phenomena could occur. They are the creation of either a focus, a z-pinch, or both. A focus is just a high current discharge, and would be characterized in this configuration by high current arcing through the plasmoid as the plasmoid is leaving the gun.
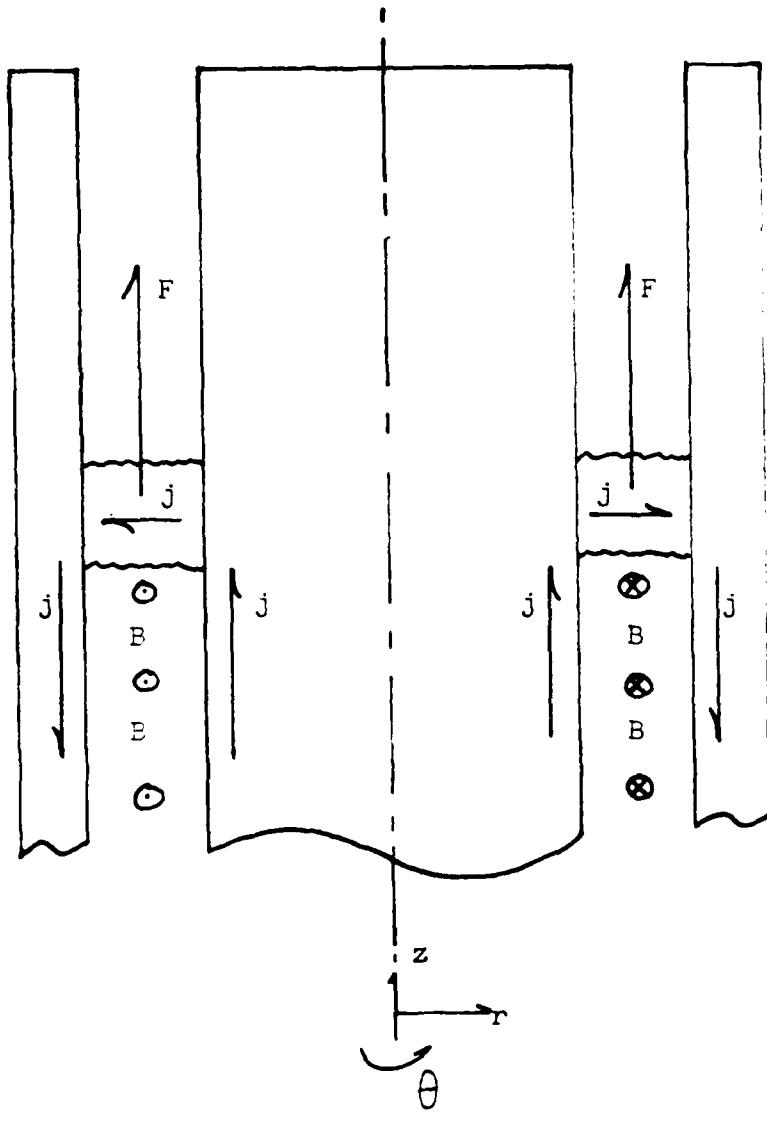
Figure 2. Plasma and current design the field.
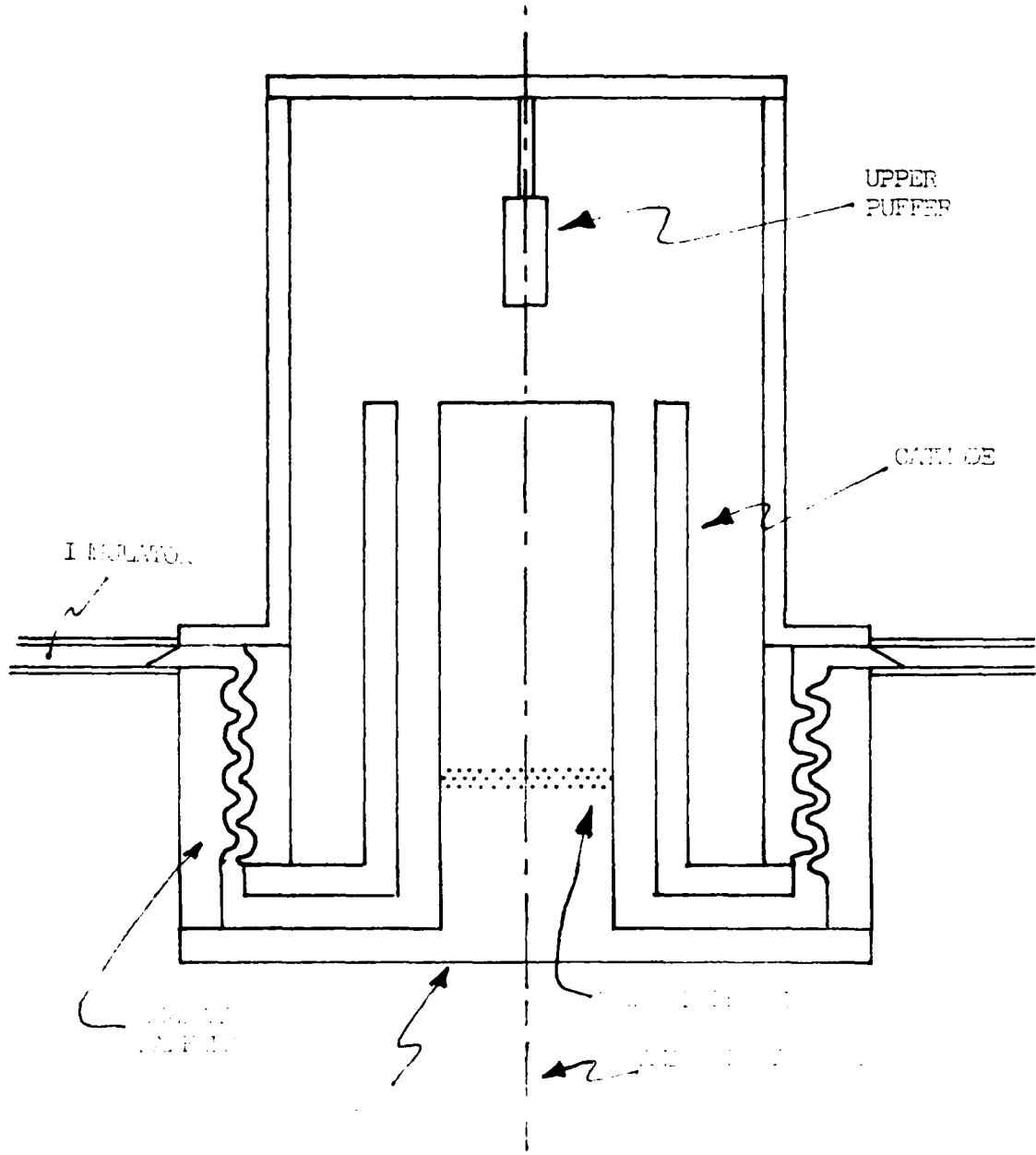
UPPER PUFFER

CATHODE

INSULATOR

Figure 1. [illegible caption text]

Any heating of puff gas from a focus would come only from
resistive heating, which is self-limiting. A focus is self-
limiting because T^4 radiative losses are much higher than the
electrical power that can be delivered to a focus. However, this
is not true for a z-pinch.

The z-pinch is a electromagneticly driven plasma implosion.
Chen (2:304) describes a pinch as follows: 'A plasma carrying a
current is confined by the magnetic field of the current itself.'
The confining (in this case imploding) force is the $\vec{j} \times \vec{B}$ force
(18:4936). The z-pinch is so named because the because the
direction of the current that is driving the implosion is along
the z-axis. SHIVA (18:4936) and SHIVA STAR experiments are
designed to produce z-pinches. In the SHIVA type geometry, as
seen in Figure 3, a high current is carried from anode to cathode
by a thin cylindrical plasma. The $\vec{j} \times \vec{B}$ force ($\vec{j}$ is in the z
direction, $\vec{B}$ in the theta direction) is directed inward (in the
negative r direction) towards the z axis. The direction of the
magnetic field can be determined from the right-hand rule. The
requirement that the magnetic field be zero in the interior of
the plasma may be determined from Ampere's law. At the z axis,
the plasma collapses on itself where it thermalizes and radiates
x-rays as a blackbody. The z-pinch is not self-limiting in the
temperature it can reach by electrical power delivery
constraints. It is limited only by how much electrical energy
can be transformed into kinetic energy of the plasma during the
implosion.

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CATHODE

FOIL INITIATION

ANODE

TIME

z

r

θ

10
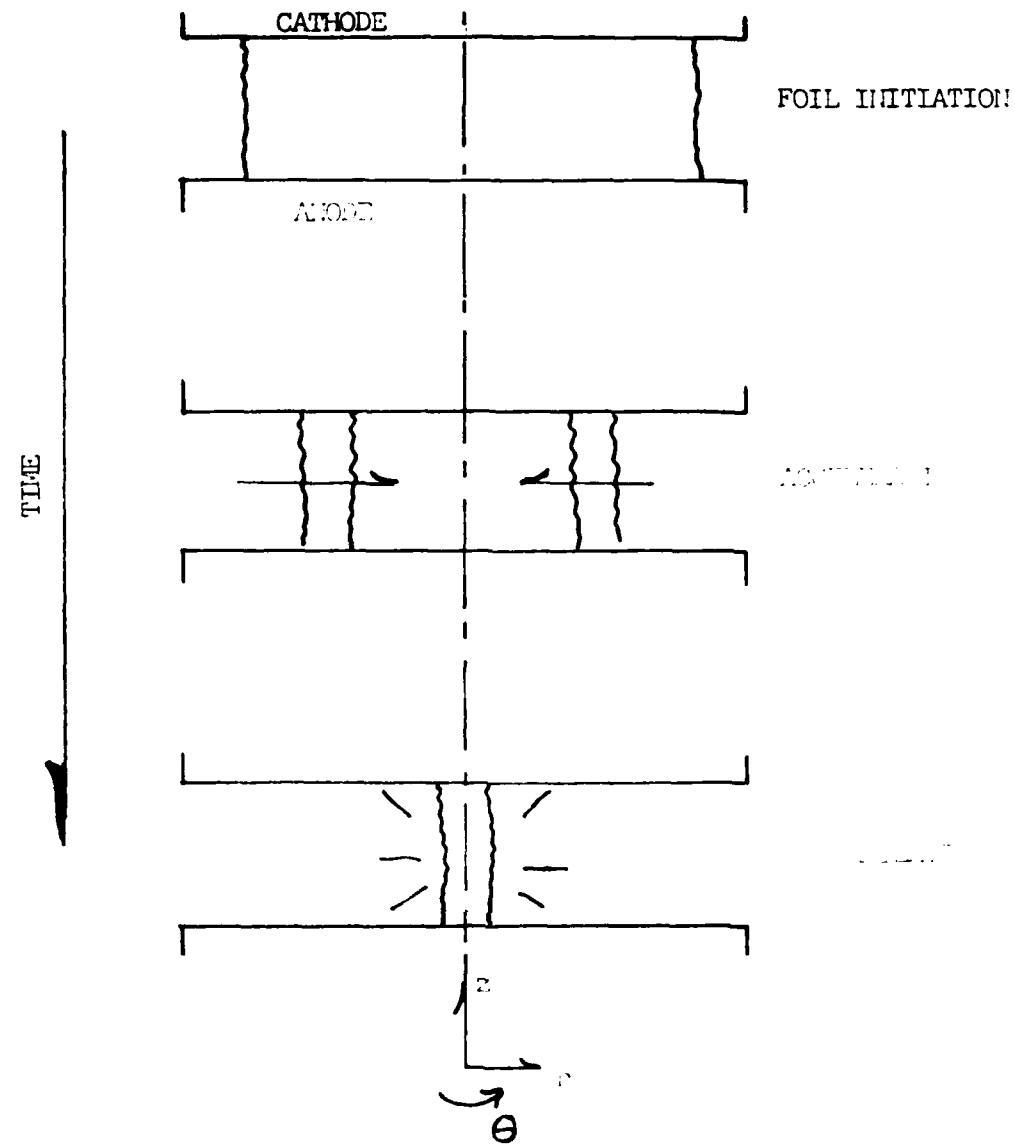
For the present experiment, a z-pinch would be most likely to occur (if in fact it does) for experiments using the double puff. Figure 4 depicts how a z-pinch would occur in this instance. The plasmoid produced by the plasma gun will act the same as the cathode in the Shiva-type geometry as in Figure 3. The differences in the two geometries are that the cathode will be moving at a very high velocity in the plasma gun driven z-pinch, and that the interior of the liner will be filled with puff gas from the second puffer. The hot plasma produced by a z-pinch would radiate its energy as blackbody radiation.

Blackbody Radiation

A blackbody is a body that absorbs all radiation incident upon it. Also, all blackbodies at the same temperature will emit radiation in the same spectrum (4:5). For a plasma to radiate as a blackbody, a very important condition must be met. Consider the optical path length of a photon of energy equal to the temperature of the plasma. This path length must be short when compared to the size of the plasma. If this condition is met, the plasma will emit radiation similar to that of a blackbody at the same temperature. This forces radiation emitted by the plasma to be in thermal equilibrium with the plasma through many scatters, absorptions, and emissions from when it is initially emitted until it escapes. For the present experiment, it is assumed that if a plasma is produced by either a z-pinch or a focus, the plasma will radiate as a blackbody.
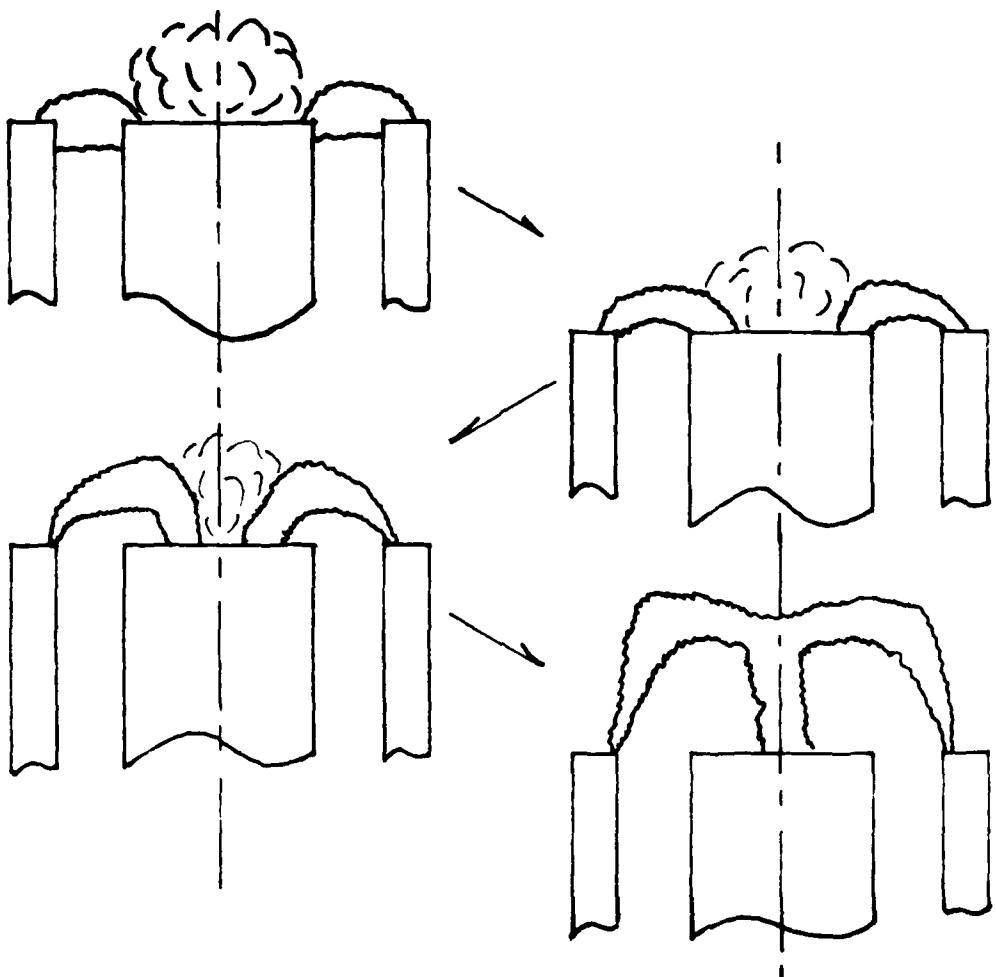
Figure ... (illegible caption text)

In order to calculate the temperature of any plasma produced by the experiment, it will be first assumed that all of the electrical energy stored in the capacitor bank is coupled into the plasma. The electrical energy stored in the capacitor bank when the bank is charged to ± 40kV is 230 kJ. The power, P, radiated by a blackbody is

$$P = \sigma AT^4 t \tag{1}$$

where T is the temperature of the plasma, A is the surface area of the plasma, and sigma is the Stefan-Boltzmann constant. Assuming the plasma radiates a constant power for a time t, the energy, E, radiated by the plasma is

$$E = Pt = \sigma AT^4 t \tag{2}$$

The plasma is taken to be a cylinder, where the surface area A is given by

$$A = 2\pi r^2 + 2\pi rh \tag{3}$$

where h is the height and r is the radius of the cylinder. Solving equation (2) for T, we find

$$T = \sqrt[4]{\frac{E}{\sigma At}} \tag{4}$$

This equation will return the blackbody temperature of the plasma given the energy radiated, the time the energy is radiated, and the surface area through which the energy is radiated.

If the plasma diameter is 0.5 cm, has a length of 2.0 cm, and a radiation time of 100 nsec, the plasma temperature would have to be 53.3 eV in order to radiate 230 kJ. If the diameter was 0.1 cm, the length was 1.0 cm, and the energy was radiated in 10 nsec, the temperature would be 161.8 eV. If the diameter was

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0.1 cm, the length was 0.1 cm, and the energy was radiated in 1 nsec, the temperature would be 468.0 eV. It must be remembered that these numbers are extremely unrealistic and err on the high side of any expected blackbody temperature. They are presented in this manner so that blackbody radiation can be ruled out as the production mechanism of the radiation signals seen.

The normalized plankian distribution $P(h\nu)$ is

$$P(h\nu) = \frac{15}{(\pi kT)^4} \frac{(h\nu)^3}{\exp(h\nu/kT) - 1} \qquad (5)$$

If kT is taken to be 470 eV and the total energy radiated is 230 kJ, then the plankian distribution numerically integrated from 30 keV to infinity gives $1.34 \times 10 \exp(-18)$ J as being radiated. The amount of this energy that can pass through the vacuum chamber is given by

$$E = E_o \exp(-\sigma\rho x) \qquad (6)$$

where sigma is the cross section (for a 30 keV photon passing through iron, sigma is 8.13 cm/gm), rho is the density (for iron rho is 7.874 gm/cm), x is the thickness of the vacuum chamber (x is 3/8 inch or 0.9525 cm), and E_o is the integral of equation (5) from kT to infinity times the total energy radiated. Evaluating equation (5) gives the amount of energy between 30 keV and infinity that passes through the vacuum chamber as being 6.08 $\times 10 \exp(-45)$ J or $3.79 \times 10 \exp(-26)$ eV. There is definitely not enough energy passing through to make up even one 30 keV photon. Integrating (5) for other energies to infinity and evaluating (6) to determine the corresponding attenuation also

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produces similar results. This demonstrates that it is very
unlikely that any radiation escaping from the chamber and being
detected by the radiation detectors is blackbody radiation.

Fusion Radiation

The gas that is puffed into the plasma gun is deuterium.
Deuterium, if raised to a high enough temperature, will undergo
nuclear fusion through the two nearly equally probable reactions
(5 : 15):

$$D + D \qquad {}^3He(0.82 \text{ MeV}) + n(2.45 \text{ MeV}) \qquad\qquad (7)$$

and

$$D + D \qquad T(1.01 \text{ MeV}) + p(3.02 \text{ MeV}) \qquad\qquad (8)$$
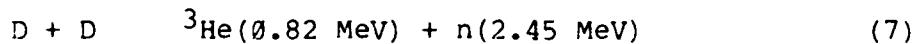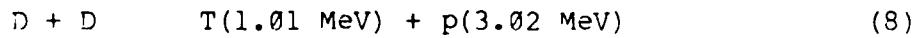
The ${}^3He$ produced in the first reaction can fuse with deuterium,
but the cross section for this reaction is 10 exp(-4) to 10 exp(-2)
that of the D-D reaction, so any ${}^3He$-D reactions will be
neglected. However, the tritium produced in the second reaction
can fuse with deuterium through the reaction (5 :15)

$$D + T \qquad {}^4He(3.5 \text{ MeV}) + n(14.1 \text{ MeV}) \qquad\qquad (9)$$
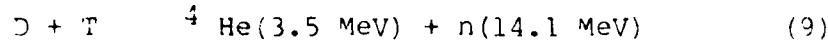
The cross section for this reaction is 50 to 100 times that of
the D-D reaction, so it would normally be assumed the D-T
reaction would occur each time the reaction in equation (8)
occurs. This is not the case here, however. The D-T reaction
will be assumed to not proceed and will be ignored. This is
because if one considers the number of deuterium atoms present
on the order of 10 exp(18)), compared to a reasonable neutron
yield (on the order of 10 exp(9)), one finds that a relatively
small amount of the deuterium present (four in 10 exp(9))

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undergoes fusion. Taking the D-T reaction to be 100 times as likely (four in 10 exp(7)), and 50% of the D-D reactions producing tritium (2 x 10 exp(9)), we see that only 50 D-T reactions will occur. Note that this also means that less than one ^{3}He-D reaction will occur. The neutrons produced are therefore going to be of one energy, 2.45 MeV.

The neutrons will affect the detectors both directly and indirectly. The direct effect is by the neutrons depositing energy into the scintillating material via proton recoil. The proton will deposit energy by ionization, thus causing a signal. This effect is countered by having one of the detectors filtered with two in. of Pb. The Pb will stop virtually all of the incident photons, but will allow virtually all of the neutrons to pass. The signal from this detector is then subtracted from the signals of the other detectors, thus removing the neutron produced component of their signals.

The indirect method mentioned in the preceding paragraph is fast neutron capture followed by the nearly instantaneous (compared to the plasma radiation pulse length) emission of a gamma ray. This neutron capture radiation is going to be much smaller than the direct neutron induced signals, and will show up as a high energy component in the spectrum if it is large enough to be seen.

Bremsstrahlung Radiation

Bremsstrahlung is radiation produced when a fast charged particle collides with matter and is either stopped or

decelerated. The maximum energy photon produced by bremsstrahlung is equal to the relative kinetic energy between the charged particle and the matter stopping the charged particle.

High voltages (greater than 30 kV) are possible between the electrode and anode in the plasma gun. If electrons are accelerated across this gap without making any collisions with any gas that is left in the plasma gun, the energy of the electrons will be equal to the potential between the electrodes. When these electrons collide with and are stopped by the electrode material (copper), then bremsstrahlung will be emitted. The maximum energy of the bremsstrahlung will thus be equal to the potential between the electrodes.

The maximum voltage between the electrodes can be estimated by knowing dI/dt, which is the change in current with respect to time. This is because the voltage, V, across an inductor (which in the present case is the plasma gun) is given by the equation

$$V = \frac{d}{dt}(IL) = L\frac{dI}{dt} + I\frac{dL}{dt} \tag{10}$$

where L is the inductance of the inductor, I is the total current flowing through the inductor, and t is time. If the circuit had a short in place at the lower puff holes instead of a plasma load, the maximum voltage expected would be 30 kV. This is because the inductance would be constant and maximum dI/dt would occur at time equal to zero, so the voltage is the capacitor charging voltage at that time.

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The inductance of the plasma gun is not constant and the load is not a short. The inductance of the load at bank initiation is 20 nH. As the annular plasma reaches the end of the plasma gun, the inductance has risen to 46 nH. As the plasma leaves the plasma gun, the inductance of the circuit rises very quickly. Also, the current drops very rapidly for two reasons. The first is that the current path disappears as the plasma leaves the gun. The second is that the inductance is rapidly increasing.

The dI/dt will be greater than the initial dI/dt. This can be seen by performing a simple calculation. First, the dL/dt term is assumed to be about zero when compared to the dI/dt term because inductance is nearly constant when compared to the change in current expected. Taking 24 in. as the plasma gun length and assuming a constant acceleration time of 3 μsec (3 μsec is chosen because as the inductance of the circuit rises, the current rise time in the circuit also rises), the average plasma velocity would have to be 2.0 x 10 exp(3) cm/sec and a final velocity of 4.0 x 10 exp(3) cm/sec (the timing of the bank and the lower puffer are adjusted so that the maximum current occurs as the plasma leaves the gun). If the plasma is 10 cm long, then it would leave the gun and interrupt the current in only 25 nsec. Taking the maximum current of a short for the bank into 46 nH as 3.2 MA, then dI/dt is 1.28 x 10 exp(14) amps/sec. Multiplying this by 46 nH gives 5.9 MV across the electrodes. This voltage can never be reached, however, because the current will first restrike in the plasma gun. It does make the point though that voltages much higher than the initial 80 kV capacitor bank charge

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are possible, and that those voltages are limited by the hold-off voltage of the low density gas that remains behind in the plasma gun after the plasma has left.

It is anticipated that these voltages will be able to accelerate electrons across the electrode gap to voltages as high as 80 keV or even higher. The energy of the electrons will be determined by the density of the gas in the gap which determines the electron path length between collisions. The bremsstrahlung produced will hopefully be observed by radiation detectors that reside outside the vacuum chamber.

Plastic Scintillators and Photomultiplier Tubes

A material that scintillates is a material that exhibits prompt fluorescence when excited electronically (11:240). The radiation detectors used for the present research use plastic for the scintillator material. Atoms in the plastic are electronically excited by the interaction with radiation passing through the plastic. The way the radiation interacts with the plastic depends upon the type of radiation and the energy of the radiation.

Photons can interact with the plastic scintillator in three different ways. They are photoabsorption, Compton scattering, and pair production. In the cases of photoabsorption and pair production, the incident photon completely disappears (11:308).

Pair production occurs when a photon of sufficient energy ($h\nu > 1.022$ MeV) interacts via an electromagnetic interaction with a nucleus. The interaction produces an electron-positron pair,

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which then proceed to interact further with the plastic scintillator (11:312-313). The positron almost immediately annihilates itself with an electron producing either two 0.511 MeV photons or three photons of varying energies. The dominate method of annihilation produces two 0.511 MeV photons. The two photons then either interact with the plastic via one of the other two interactions, or they escape the plastic. The electron produced can also interact with the plastic. The methods by which it interacts will be considered later in this section.

Compton scattering occurs when a photon scatters off an electron. The energy gained by the electron is equal to the change in energy of the photon (11:309-312). The scattered photon will either again interact with the plastic in one of the three types of interactions mentioned or it will escape. The methods by which the Compton electron interacts will be considered later in this section.

Photoabsorption occurs when a photon interacts with an atom, and the photon completely disappears. In the place of the photon, an energetic photoelectron appears. The energy of the photoelectron is equal to the energy of the incident photon minus the binding energy of the electron to the atom (11:308-309).

The energetic electrons produced by either Compton scattering, photoabsorption, or the electron from pair production interact with the plastic through the electromagnetic interaction. These energetic electrons ionize the plastic and deposit energy as they pass through the plastic. The ionization produces low energy photons that the P-M tube can detect when

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recombination of the ions occur. The number and energy of the low energy photons produced in the plastic is proportional to the energy deposited. Thus the low energy photon output of the plastic is proportional to the energy of the fast electrons.

The plastic scintillator presently used is composed completely of carbon and hydrogen. For photon energies of above a hundred keV or so and below 1.5 to 2 MeV, the dominate interaction with carbon and hydrogen is the Compton effect. Therefore, if a pulse height analysis of the output of a photomultiplier tube that is coupled to the plastic scintillator is performed, the pulse height distribution will consist of a Compton edge and a Compton continuum.

A photomultiplier tube consist of two major structures. They are a photocathode section and an electron multiplier section. The photomultiplier tube acts to first use any photon that strikes it to free an electron from the photocathode. The electron is then used to produce a shower of secondary electrons in the electron multiplier section. The number and energy of photons striking the photocathodes is proportional to the energy that has been deposited in the scintillator material. The number of electrons liberated from the photocathode is proportional to the number and energy of photons striking it. The number of secondary electrons produced at all the amplifying stages of the electron multiplier is also proportional to the number of electrons striking them. Therefore, after the final amplifying stage of the photomultiplier tube (at the anode), the charge deposited will be proportional to the energy deposited in the

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scintillator material.

The photocathode consists of a material that will emit an electron if a photon strikes it. The probability that a material will undergo photoemission for an incident photon is called its quantum efficiency. This process of photoemission occurs in three stages. First, an incident photon is absorbed with the energy of the photon being transfered to an electron in the photoemissive material. Second, the electron migrates to the surface of the photoemissive material. Third, the electron escapes from the surface of the photocathode.

An electron that escapes from the photocathode is "directed" by the use of electric fields into the electron multiplier section of the photomultiplier tube. The electron is attracted to and strikes a dynode. The dynode will then re-emit several electrons. These electrons are again directed by an electric field to the next dynode, where each electron that strikes it causes the re-emission of several electrons. This continues until after the final dynode is reached, where the anode collects a detectable amount of electrons.

This process works exactly the same way when more than one photon strikes the photocathode. Thus, the current produced by the electrons that are collected on the anode is in direct proportion to the energy deposited in the scintillator material. The energy deposited in the scintillator material is a function of the energy and type of radiation that is incident on the scintillator material. Thus, the current produced by the photomultiplier tube is dependent upon the energy and type of

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radiation incident on the scintillator material.

Spectral Deconvolution

As seen in the previous section, the current produced by the type of radiation detectors presently used is proportional to the amount, type, and energy of radiation striking it. It will now be shown that it is possible to use several radiation detectors whose response is like that described above to "back-out" or deconvolute the spectrum of radiation striking them.

In order to do this, first assume we have a radiation source whose spectrum, $\underline{x}$, consists of discrete energy lines. The strenjths of the lines are not known, but their enerjies are. Also assume the detectors being used have a energy response function, $\underline{R}$, and the number of detectors used equals the number of energy lines. Third, assume the detectors used to measure this radiation are filtered by filters of different transmission, whose transmission function is $\underline{T}$. Therefore, the response of the detectors, $\underline{b}$, will be given by the equation

$$\underline{T}\ \underline{R}\ \underline{x} = \underline{b} \tag{11}$$

Taking

$$\underline{A} = \underline{T}\ \underline{R} \tag{12}$$

we have

$$\underline{A}\ \underline{x} = \underline{b} \tag{13}$$

The spectrum was assumed to be discrete, so this is just a simple matrix equation where $\underline{b}$ is measured experimentally, $\underline{R}$ is arrived at by calibrating the detectors, and $\underline{T}$ is taken from photon transmission tables. The solution then is found by multiplying

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both sides of equation (13) by A^{-1}. This gives

$$\underset{\sim}{A}^{-1}\underset{\sim}{A}\underset{\sim}{x} = \underset{\sim}{A}^{-1}\underset{\sim}{b} \qquad (14)$$

or

$$\underset{\sim}{x} = \underset{\sim}{A}^{-1}\underset{\sim}{b} \qquad (15)$$

The solution thus becomes one of simply finding the inverse of $\underset{\sim}{A}$. This is not always trivial, though, as is the case here. The problem is that although $\underset{\sim}{A}$ is not singular, it is "nearly" singular. A problem like this is called ill-posed, and a matrix like $\underset{\sim}{A}$ is called "ill-conditioned" (8:2,48). Therefore, many of the usual methods of matrix inversion will not work, so some type of a specialized technique must be used, such as an iterative technique.

An Iterative Technique

Iterative methods are so called "because each method is designed to generate a sequence of vectors (iterates), $\{\underset{\sim}{x}^{(k)}\}_{k=0}^{\infty}$ which converge to the true solution, $\underset{\sim}{x}_t$, of $\underset{\sim}{A}\,\underset{\sim}{x} = \underset{\sim}{b}$" (8:60). The iterative technique used here is exactly that used by Degnan et al. (3:264-265). The technique works by solving the following two equations,

$$VC_k(t) = \int R_k(E)\, S(E)\, dE \qquad (16)$$

and

$$S'(E) = S(E)\; \frac{\sum_k R'_k(E)\, (V_k(t)/VC_k(t))^{\gamma}}{\sum_k R'_k(E)} \qquad (17)$$

where $V_k(t)$ is the observed signal for the detector k, $VC_k(t)$ is the calculated signal for the detector k, $S(E)$ is the trial spectrum for the present iteration, $S'(E)$ is the corrected

spectrum, $R_k'(E) = R_k(E)/\int R_k(E)\,dE$, E is the photon energy, and gamma is 1.2. Gamma is an overcorrection parameter that is used to speed convergence. Equation (17) is called the corrector equation because it modifies or corrects a guess at the spectrum to produce a new guess at the spectrum.

The equations are solved by first calculating several functions that do not change throughout the iterative process. They are $R_k(E)$ which is the response of the kth detector as a function of energy, $R_k'(E)$ which is the normalized response of the kth detector as a function of energy, and $\sum_k R_k'(E)$ which is the sum over the k detectors of $R_k'(E)$. The iterative process consists of taking a guess at the spectrum, $S(E)$, and computing a calculated detector response, VC_k, for that spectrum by integrating equation (16). The calculated detector response is then used in the corrector equation to produce a new guess at the spectrum. This process is repeated until the desired convergence is reached.

The program DCON that solves equations (16) and (17) is the same program used by Degnan et. al. (3:264-269), except for some restructuring, changes in the input/output routines, and detector response function/filter calculation routine changes. The program is listed in Appendix A, is written in FORTRAN, and as listed will execute on a VAX 11/780 using the UNIX operating system. The program will execute under the VAX/VMS operating system by just modifying the two OPEN statements in subroutine DATIN.

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III. Equipment

PUFF Capacitor Bank

The experimental measurements were taken on a large fast capacitor bank called PUFF which is located at the Simulators and Advanced Weapons concepts Branch of the Air Force Weapons Laboratory (AFWL/NTYP) at Kirtland Air Force Base, New Mexico. The PUFF bank consists of 48 six µF capacitors which are operated in a configuration of two groups of 24 capacitors each. The capacitors in each group are connected in parallel, with the two groups connected in series. The capacitance of the bank is 72 µF in this configuration. The bank is charged so that both groups of capacitors are charged to voltages that are opposite in sign, but equal in magnitude. For the present series of experiments, PUFF was charged to either ±30, 35, or 40 kV. Charging PUFF to these voltages produces stored electrical energies in the bank of either 129.6, 176.4, or 230.4 kJ, respectively. The inductance of the bank, transmission lines, radiation baffles, and plasma gun up to the gas puff holes in the anode is 20 nH, so the quarter cycle rise time of the bank for a short at that point is 1.9 µsec.

Radiation Baffles

The current from the bank is carried from the vacuum-insulator interface to the load region of the apparatus by a large "screw" or helix (see Figure 1, page 7). The helix acts as a set of radiation baffles, and prevents radiation produced by the plasma from reaching the insulator at the vacuum-insulator

interface. The baffles are necessary because the untra-violet and soft x-ray radiation from the plasma is able to remove electrons from the insulator by the photoelectric effect. If enough electrons are removed from the insulator, electrical breakdown occurs at the vacuum-insulator interface. Electrical breakdown would divert current from the load region. The diversion of current from the load region would degrade the performance of the plasma gun.

Coaxial Plasma Gun

The load region of the apparatus consists of a 24 in. long coaxial plasma gun. The inner diameter of the plasma gun (the anode) is 5 in., and the outer diameter (the cathode) is 7 in. The anode is a hollow copper tube that has 96 1/4 in. holes spaced around the base 13 in. from the top. These holes allow deuterium gas to be puffed into the plasma gun just prior to bank initiation. The gas is puffed into the gun through the holes by the lower puffer. The cathode consists of 24 1/2 in. diameter copper rods equally spaced around the anode, with a one in. gap between the anode and the cathode. The rods that make up the cathode stop even with the top of the anode. The anode is covered with a circular copper plate. There is a second gas puffer located six cm above the center of the copper plate.

The plasma gun operates as follows. First, gas is puffed into the plasma gun by the lower puffer at the base of the gun. The bank is then initiated. As the voltage in the plasma gun increases, the gas breaks down electrically and current, $\vec{j}$,

begins to flow through the gas. The gas is heated resistively, which turns the gas into a plasma. The rising current creates a rising magnetic field, $\vec{B}$, beneath the plasma. The $\vec{J} \times \vec{B}$ force accelerates the annular plasma upward. The plasma is accelerated up the plasma gun to the end of the plasma gun where it either pinches on itself in a z-pinch, drives a second puff of gas from the upper puffer into a z-pinch, or just departs as what is known as a plasmoid and impacts the plate at the top of the vacuum chamber.

Gas Puffers

The gas that is puffed into the plasma gun is deuterium. Three cubic cm of deuterium at 900 psi is puffed near the base and is directed radially outward in an annulus. The mass of the gas in the coaxial gap is about $1 \times 10 \exp(-3)$ gm at the time the main capacitor bank discharge is triggered.

A second puff of deuterium comes from a puffer of six cubic cm in volume at 30 psi. This puffer has a single puff hole in it, is aligned along the vertical asix of the anode, and is pointed down at the top of the anode. The mass of deuterium in the second puffer is about $2.7 \times 10 \exp(-4)$ gm. The second puff occurs six cm above the end of the plasma gun. The timing of the second puff is such that the maximum coupling of energy occurs between the first puff and the bank into the deuterium of the second puff. The amount of energy coupled can be measured by examining either the x-ray output of the pinch, the fusion neutron yield of the pinch, or the voltages and currents of the

experimental apparatus at various locations.

The two puffers are activated by the discharge of small capacitor banks. The first puffer operates by a piston being released. The piston is held in place by a small (one in. long by 1/2 in. in diameter) wooden rod. The wooden rod has a 0.005 in. hole along the axis. The hole contains a 0.001 in. diameter copper wire and water. The wire is attached to one of the small capacitor banks. When the capacitor bank fires, the copper wire explodes. The exploding wire heats the water and turns it into steam. The pressure from the steam destroys the wooden rod. The piston is thus released, allowing deuterium to puff into the plasma gun through the puff holes in the anode.

The second puffer is operated by having an electromagnetically driven hammer strike against a re-usable diaphram. The puffers are similar in design to those used at Sandia National Laboratory and Lawrence Livermore National Laboratory (12:65).

The Detectors

The detectors presently used are called NP154X, and are built by EG&G. They are sealed integral units consisting of a 2 and 1/4 in. diameter by 2 and 1/4 in. long plastic NE111 cylindrical scintillator coupled to a ND2020 photomultiplier tube. The NE111 is painted black, except for the end coupled to the photomultiplier tube. The scintillator and the photomultiplier tube are encased in a 3/16 in. aluminum jacket, except for the front end, which is only covered by a three mil aluminum foil. Also, the detectors are wrapped in 5/8 inch of lead which

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cuts down the amount of scattered radiation incident on the detectors.

The operating voltage of the detectors is -2800 V. They were originally acquired as neutron time-of-flight detectors, and have been used as such for SHIVA I, SHIVA I', SHIVA II, and SHIVA STAR experiments at AFWL/NTYP.

IV. Procedure

Detector Calibration

Calibrating the detectors consists of three steps. The first consists of acquiring pulse height distributions for various energy radioactive sources. The second is to determine relative response functions for the various energies. Step three entails using the ^{60}Co calibration data provided by the manufacturer to produce an absolute calibration.

The first step consists of using a single detector that is connected to a pulse height analysis system to obtain pulse height distributions of a set of calibrated radiation sources. This is done by connecting the output of one of the detectors to a linear amplifier which is connected to a multichannel analyzer (MCA). A calibration source is then placed a known distance from the detector, and the MCA is started. Pulse height data from the detector is taken until a sufficient number of pulses have been recorded. This is repeated for several other calibrated sources of different energies. A background pulse height data run is also performed.

In addition to the pulse height data taken, there is also a need to know what pulse voltage corresponds to what channel in the MCA. This is found by connecting a pulse generator into the linear amplifier where the detector was. Voltage pulses of known voltage are then sent to the MCA, and the corresponding channel number is recorded.

The second step involves taking the data acquired in the

first step and analyzing it on a computer. First, all of the
pulse height distributions n(c) are normalized to the same count
time of one second.

$$N(c) = n(c)/t \qquad\qquad (18)$$

where N(c) is the time normalized pulse height distribution, c is
the channel number, and t is the count time for n(c) in seconds.
The time normalized background pulse height distribution B(c) is
then subtracted from the pulse height distributions of the
various sources.

$$M(c) = N(c) - B(c) \qquad\qquad (19)$$

Next, the calibration data of the pulse height analysis system is
used to determine what height voltage pulse corresponds to what
channel for each channel. These voltages are then multiplied by
the number of pulses in each corresponding channel, and the
voltage-pulse number products are all summed.

$$S = \sum_{c} M(c) * V(c) \qquad\qquad (20)$$

where V(c) are the voltages of the pulses corresponding to each
channel c, and S is the sum of all of these voltage-pulse number
products. Using the source to detector distance, an area is
determined for a sphere of radius equal to that distance. This
surface area is divided into the area of the detector presented
to the source. This is multiplied by the present activity of the
source in counts per second.

$$D = A \pi r^2 / 4 \pi R^2 = A r^2 / 4 R^2 \qquad\qquad (21)$$

where r is the radius of the detector face, R is the source to
detector distance, A is the present source activity, and D is the
activity passing through the detector. Now, dividing S by D, we

arrive at an average voltage pulse height per count, P.

$$P = S/D \tag{22}$$

If E is the photon energy of radiation in MeV, and that is divided into P, then we obtain the average pulse height per MeV H, for the energy E.

$$H(E) = P/E \tag{23}$$

The above procedure is repeated for several different energy photon sources, so that the function H(E) is determined for several different photon energies. One of the photon sources used must be ^{60}Co so that an absolute calibration may be performed.

The final step is to calibrate each detector separately to its own absolute ^{60}Co calibration data. The absolute calibration is obtained from data provided by the manufacturer of the detectors. The value of H(E) for ^{60}Co is divided into the calibration sensitivity value which has units of Amps-cm^2-sec/γ-MeV. This ratio is then multiplied by the other values of H(E) to obtain the calibration for the other energies obtained when the relative calibration was performed.

Choosing The Filters

The filters for the detectors are chosen by two simple criteria. The first is determining exactly what filters are available. The second is to use the known cross sections of these filters and a resonable predicted spectrum with varying combinations of the filters to calculate detector voltage signals. These detector signals are then used as input to the

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program DCON. The set of filters that produces the "best" (i.e.
the calculated spectrum is closer to the "original" predicted
spectrum, and the error estimated by the program is smaller)
results should be used. It should also become apparent that as
the original spectrum is changed, the sets of filters that work
best also changes.

Acquisition Of Data

The details of data acquisition presented here will only
concern the operation of the NPM54X detectors and the method of
obtaining current traces for them. The operational procedures of
the PUFF bank and its associated equipment will not be presented.
There are two reasons for this. The main reason is that the data
being acquired by the author was only an ancillary part of a
larger effort. The second reason is preparing the bank for a
shot and firing the bank requires at least four people performing
different tasks simultaneously. Therefore, the author did not
perform all of the tasks.

The detector voltage traces were obtained as follows.
First, the detectors were placed on a stand such that the centers
of all the detectors were in a horizontal plane with an elevation
one in. above the anode surface. Then, the filters were mounted
on the detectors. Next, the detectors were aligned so that they
were pointing at a spot one in. above the center of the anode.
After power and signal cables were connected, detector to center
of anode measurements were obtained. Power was then applied to
the detectors. The oscilloscopes on which the data was taken

were located in a screen room, and personnel in the screen room readied the oscilloscopes for bank firing. The bank was then charged and fired. The voltage traces from the oscilloscopes was recorded on Polaroid film. From knowing the oscilliscope settings and the graticule images on the traces, voltages can be read directly off the traces.

Normally, timing marks or fiducials are added to the voltage traces. It was not possible to add them for this series of experiments because of a serious noise problem the fiducial generator was causing on the traces. For the data presented in the results chapter, pulse timing information is obtained by assuming the voltage peaks in all the traces occurred simultaneously. Also, prior to a shot, fiducials were used to synchronize the traces.

Deconvolution Of Spectrum

The voltages obtained in the manner described in the previous section are entered into a data file that is used as input into a program that acts as a pre-processor to the program DCON. The pre-processor program is called PIN, and it removes any $1/R^2$ dependance the data has and changes the voltages to the appropriate units. PIN then outputs a file that is used as the input file for DCON. The program DCON is then run, and the program outputs a spectrum for the experiment. PIN is listed in Appendix D, and sample input for PIN is listed in Appendix E. Appendix B, the sample input for DCON, is also an example of output from program PIN.

Choosing New Filters

The final step in this process is to choose a new set of radiation filters. This is done by using the spectrum obtained from the spectrum deconvolution step as a predicted spectrum for the program DCON. A program is executed that accepts the spectrum and a set of filters as input. The program outputs detector responses in a format DCON can read. DCON is executed, and it calculates a spectrum. This spectrum is then compared to the experimental spectrum. When the filters used for the experiment are used as input to the program that generates the detector responses, the differences in spectra should give some idea about the error in the experimental spectra that result from non-optimum filter selection. Again, after many different sets of filters are tried, it should become apparent that certain sets of filters work better than others for the actual spectra the plasma gun produces. One of the sets that work best should be used as filters for the next experiment.

V. Results

Detector Calibration

Detector S/N 080-LV was used to perform the relative photon energy calibration. The relative detector response as a function of incident photon energy is shown in Figure 5. The "x's" in the figure is experimental data that has been normalized to a linear least squares fit to the same energy as the ^{60}Co response data. The "dashed" line is the linear regression fit to the data. For the linear regression, the x intercept was calculated to be 0.816 ±0.151, and the slope was calculated to be 0.147 ±0.186 per MeV. The curved "solid" line in Figure 5 shows mass energy absorption coefficients for polystyrene (C8H8) (7:74), which is close in chemical composition to NE111 (NE111 is composed entirely of hydrogen and carbon, and has an H:C atom ratio of 1.096:1) (11:246). The data for polystyrene has been log-log interpolated, and as shown is normalized so that the mass energy absorption coefficient at the ^{60}Co photon energy is equal to one. The "dotted" line in the figure is the flat detector response function. The program DCON is written so that any of three relative detector response functions may be used. They are the linear regression, the flat response, or the log-log interpolated polystyrene mass absorption coefficients.

The absolute detector responses are computed in the program DCON by multiplying the relative detector response for the photon energy in question by the response for ^{60}Co that was provided by the manufacturer. It will be shown later that the relative

NPM54X Relative Detector Responses
Normalized To Cobalt 60
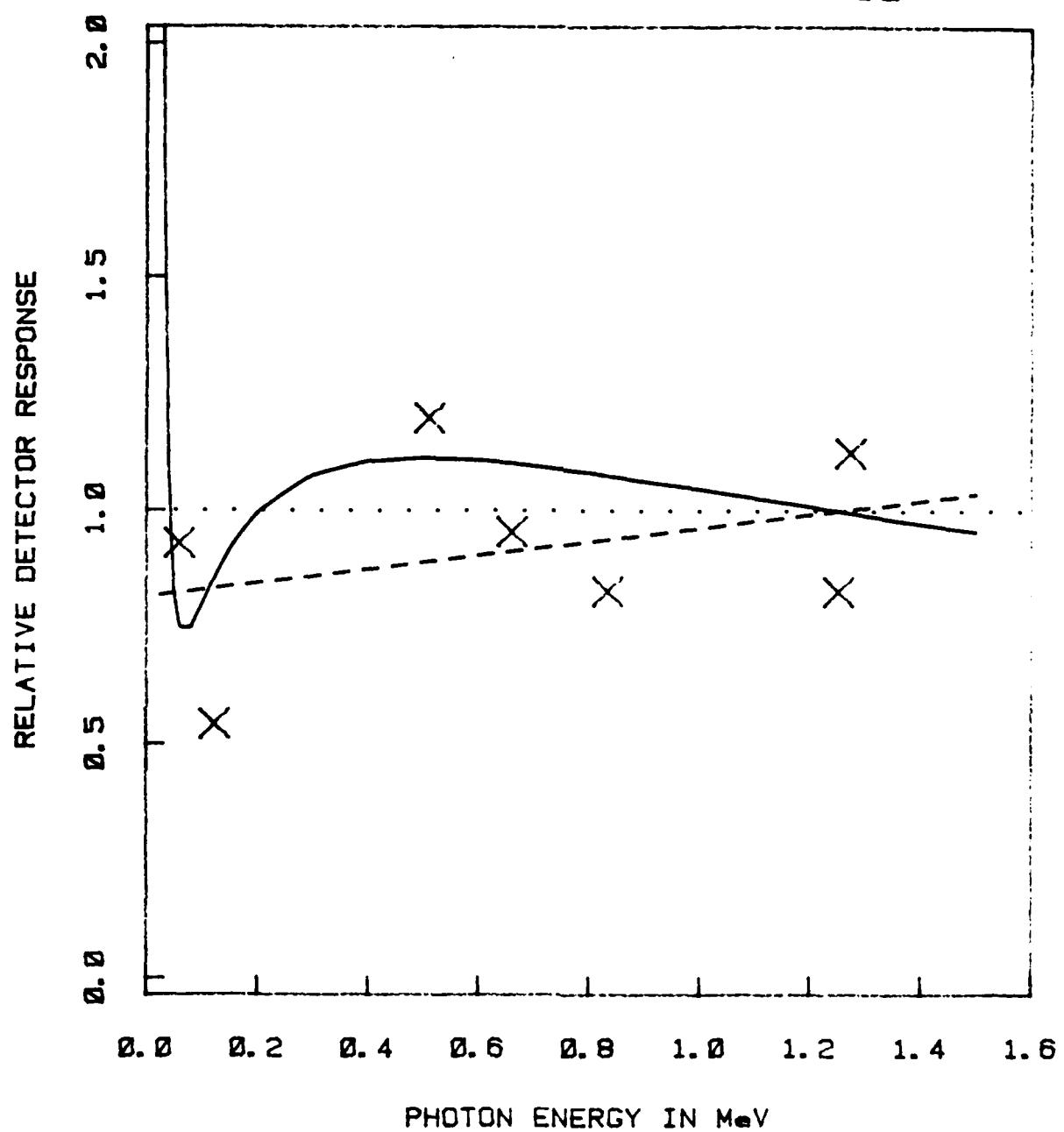
FIGURE 5. NPM54X Relative Detector Response Normalized to Cobalt 60. In this figure, the solid line is the polystyrene (C8H8) mass-energy absorption data (7:74); the dotted line is the flat response; and the dashed line is the linear regression response of the experimental data, which are the x's.

detector response chosen has very little effect upon the photon spectrum determined by DCON.

Filter Selection

The material available for the filters was Al, Cu, and Pb. The sizes that were available were 1/16 in. and larger in 1/16 in. increments for the Pb. For the Al, and Cu, 1/8 in. increments were available.

For the trial spectrum input into DCON, a two temperature exponentially decreasing spectrum was used. The two temperatures chosen were 10 and 50 keV, and were equally weighted. DCON was executed repeatedly using this spectra as input with the filters varied each time. It was found that the spectrum produced by DCON did not vary much (it was within 5%) from the input spectra in the interior of the energy mesh or bins. Near the endpoints of the mesh, the spectra output varied sometimes by as much as a factor of two or three. The worst of these results occurred when a different filter material was used on each detector. The best occurred when all the filters were of the same material. When the same material was used, the errors across the interior of the energy mesh being examined could be brought to within two or three percent and the error at the ends could be brought to within 20%. The smallest errors were achieved when the filter material was the same, the Z of the material was at or higher than that of steel, and the thicknesses of the filters varied by a factor of three from one detector to the next. Also, when using one of the detectors unfiltered, a requirement for the lowest filter size to

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have a thickness of 1/8 in. or so was noticed for filters other than Pb. This is because the supposed "bare" detector is really already shielded by 3/8 in. of steel vacuum chamber, and 1/16 in. of Al or Cu does little additional attenuation.

The first sets of filters used did not take this information into account because they were being used to determine what size of detector signals could be expected from what filters. The data presented in the next section is from one of these shots. The detector serial numbers and filters used are listed in Table I.

Shot D6-4

The shot parameters and experimental data is listed in Table II. The signal from the two in. Pb filtered detector was zero, so the signals were input into the preprocessing program as shown. The deconvoluted spectrum is shown in Figure 6. The spectrum is plotted on a log-log plot. The linear portion of the spectrum extending from 35 keV to 100 keV demonstrates a spectral dependency upon the photon energy raised to some power over that region. There is also an interesting "bump" in the spectrum that is centered at about 500 keV. The neutron yield of the shot was 4.5x10 exp(7). The total radiated power above 35 keV was 650 MW.

The spectrum shown in Figure 6 uses as input all of the detector responses listed in Table II except for the response from detector 079. For this shot, there were two detectors that were identically filtered with 1/8 in. of Cu (detectors 077 and 124). Figure 7 shows a comparison of the spectrum shown in Figure 6 and the spectrum computed when only one of the 1/8 in. Cu filtered

TABLE I.

Detector Filters For PUFF Shot D6-4

Detector Serial Number	Filter Material Thicknesses in CM				
	Fe	Air	Al	Cu	Pb
079-LV	0.9525	352.7	0.0	0.0	5.08
077-LV	0.9525	356.9	0.0	0.3175	0.0
123-LV	0.9525	358.8	0.635	0.3175	0.0
122-LV	0.9525	358.1	0.0	0.0	0.15875
124-LV	0.9525	359.7	0.0	0.3175	0.0
125-LV	0.9525	356.6	0.635	0.0	0.0

TABLE I. Detector Filters for PUFF shot D6-4. The Fe filter is the vacuum chamber wall. Note that dtectors 077 and 123 have the same thickness of Cu filteration.

TABLE II.

Shot Data For PUFF Shot D6-4

Detector Serial Number	079-LV	077-LV	123-LV	122-LV	124-LV	125-LV
Detector Voltage	-0.0 V	-3.9	-5.0	-1.5	-5.0	-27.5
Error in Reading Trace	0.3 V	0.3	0.3	0.3	0.3	0.3
Filter Material	Pb	Cu	Al+Cu	Pb	Cu	Al
Detector Response Amps/(-Mev/Cm**2/Sec)	2.49E-10	1.88E-10	1.56E-10	1.41E-10	1.63E-10	1.62E-10

Detector Bias = -2800 V

First Minute Ag Neutron Activation Counts = n = 7662 Counts/Min

Ag Activation Counter Background = b = 160 Counts/Min

Neutron Yield = 0.6 x 10 exp(6) x (n-b) = 4.5 x 10 exp(7)

Bank Voltage = +/- 40,000 V

Lower Puffer Pressure = 900 psi

Upper Puffer Pressure = 0 (No upper Puff)

Total Radiated Power above 35 keV = 6.5 x 10 exp(8) watts

Table II. Shot Data for PUFF shot D6-4. The detector response listed is for ^{60}Co gamma radiation that was provided by the manufacturer.

PUFF Plasma Gun Spectrum   Shot D6-4
Five Detector Signals Used
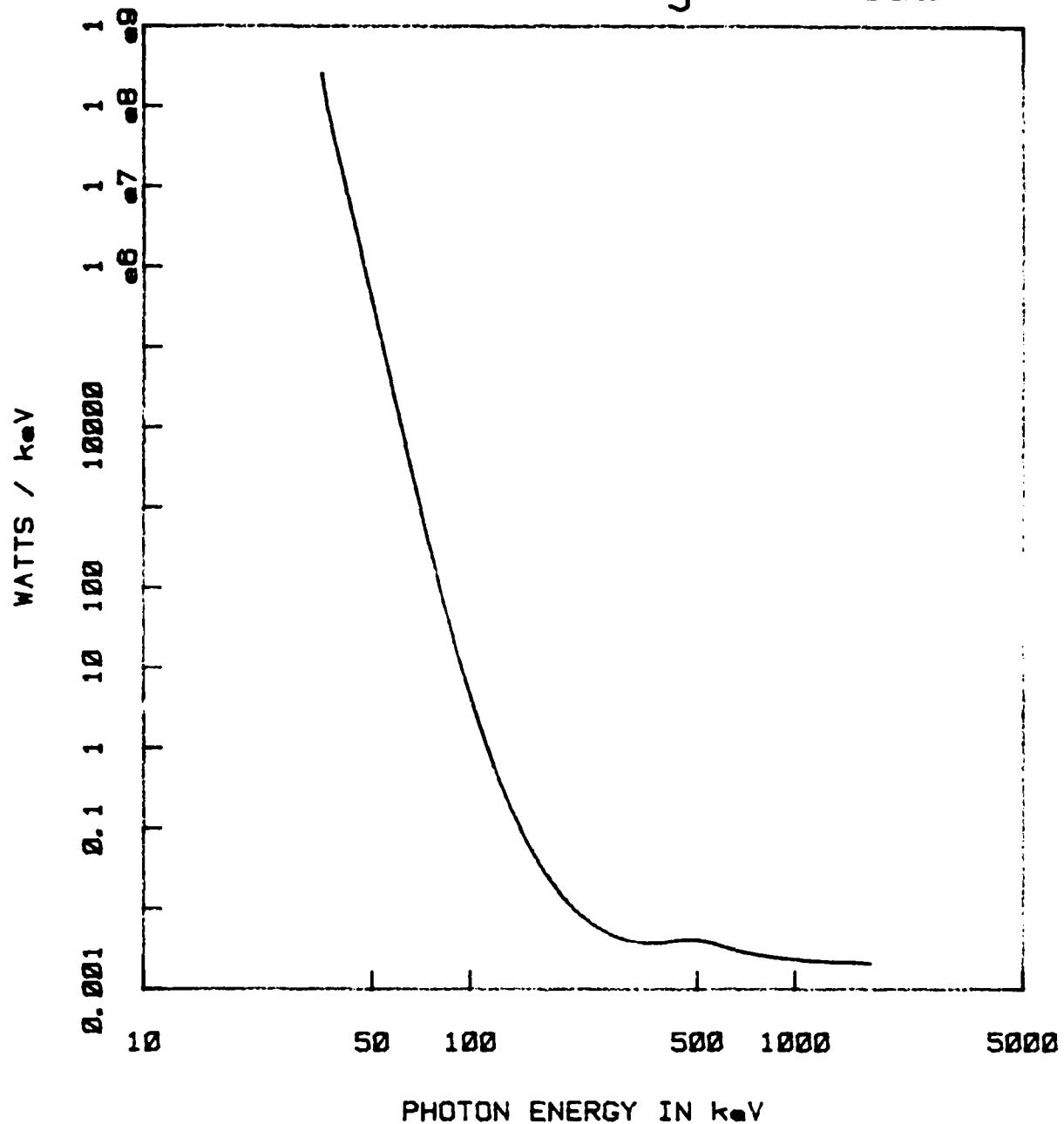
FIGURE 6. PUFF Plasma Gun Spectrum Shot D6-4 Five Detector Signals
Used. This figure shows the results of the deconvolution
of PUFF shot D6-4. All detector signals, except for the
signal from detector 079 (filtered with 2 in. Pb) were
used. The response function used was the C8H8 log-log
interpolation.

PUFF Plasma Gun Spectrum   Shot D6-4
Copper Filtered Detector Signals Varied
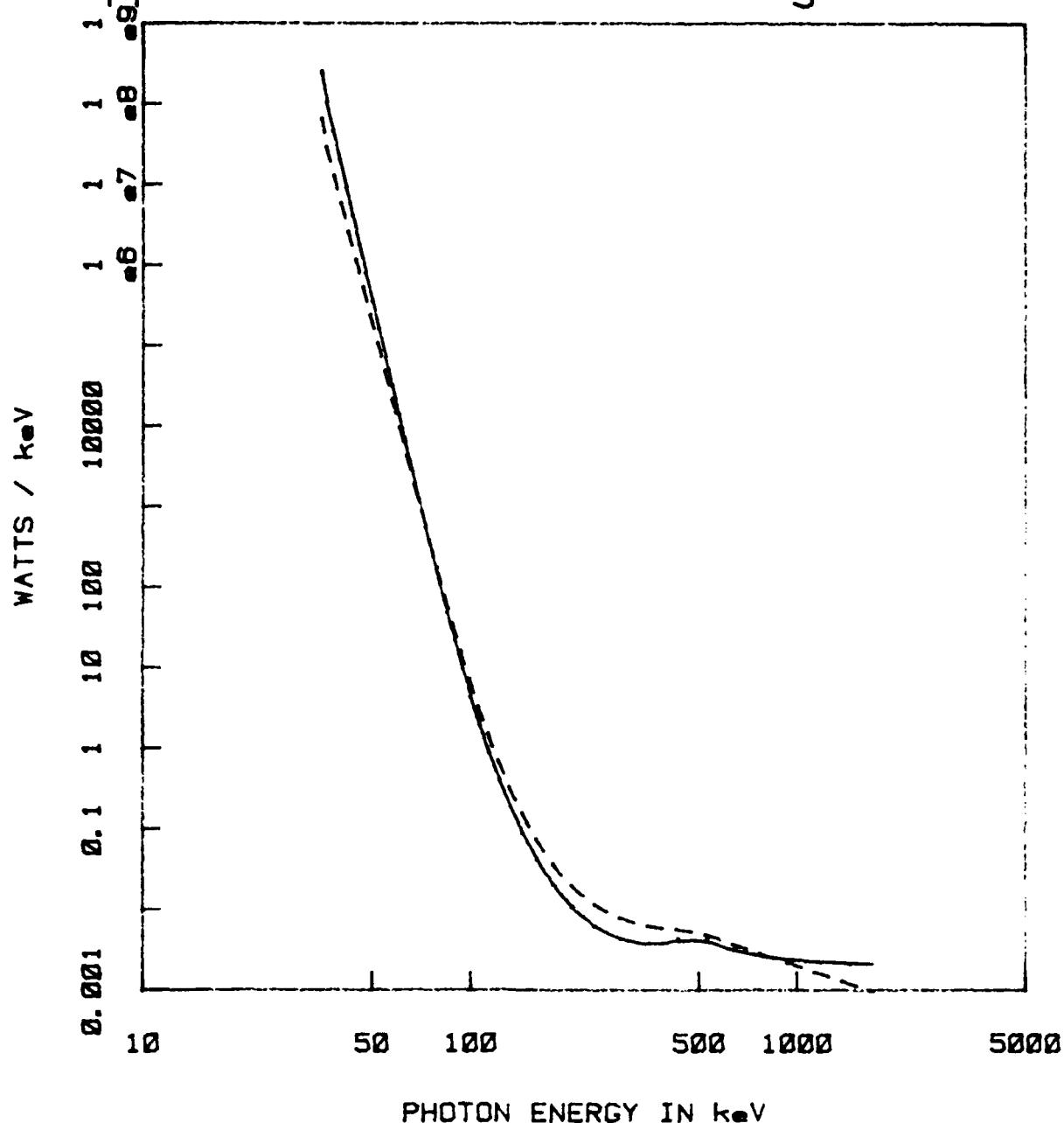
FIGURE 7. PUFF Plasma Gun Spectrum Shot D6-4 Copper Filtered
Detector Signals Varied. The effects on the deconvo-
lution of using only one signal from the two Cu filtered
detectors (Ø77 and 124) is shown. The solid line is the
same spectrum as in Figure 6. The dotted line (hidden
by the solid line) is from neglecting detector Ø77.
The dashed line is from neglecting detector 124. The
C8H8 response function was used in all three cases.

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detectors is used as input. The solid line is the spectrum shown
in Figure 6, the dotted line is the spectrum computed without the
data from detector 124, and the dashed line is the spectrum
computed without the data from detector 077.

As stated above, the spectrum in Figure 6 exhibits a power
law dependency between 35 and 100 keV. The other two spectra
shown in Figure 7 demonstrate this also. This power law
dependency has the form

$$S = \beta(h\nu)^{\alpha} \qquad (24)$$

where S is the spectra, beta is a constant determined from the
height of the linear region at some arbitrary $(h\nu)$, $(h\nu)$ is the
photon energy, and alpha is the slope of the linear region.
Table III shows alpha's and beta's for the three spectra in
addition to the alpha's and beta's for the spectra presented in
Figure 8.

Figure 8 shows the effect of changing the absolute detector
response. Again the solid line is the same spectrum shown in
Figure 6. This spectrum was computed using the polystyrene mass
energy absorption coefficients for the absolute detector
response. The dashed spectrum was computed using the flat
detector response. The dotted spectrum was computed using the
linear regression detector response.

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TABLE III.

Power Law Equation: $S=\beta(h\nu)^{\alpha}$

	Alpha	Beta
All Five Detectors:		
Polystyrene Detector Response	−17.024	2.9895x10exp(34)
Flat Response	−17.020	2.4272x10exp(34)
Linear Regression Response	−17.202	6.9319x10exp(34)
Four Detectors:		
Detector 077 Excluded	−15.403	2.957x10exp(31)
Detector 124 Excluded	−17.141	5.316x10exp(34)

Table III.   Parameters For Spectral Power Law Fit

PUFF Plasma Gun Spectrum   Shot D6-4
Detector Response Function Varied
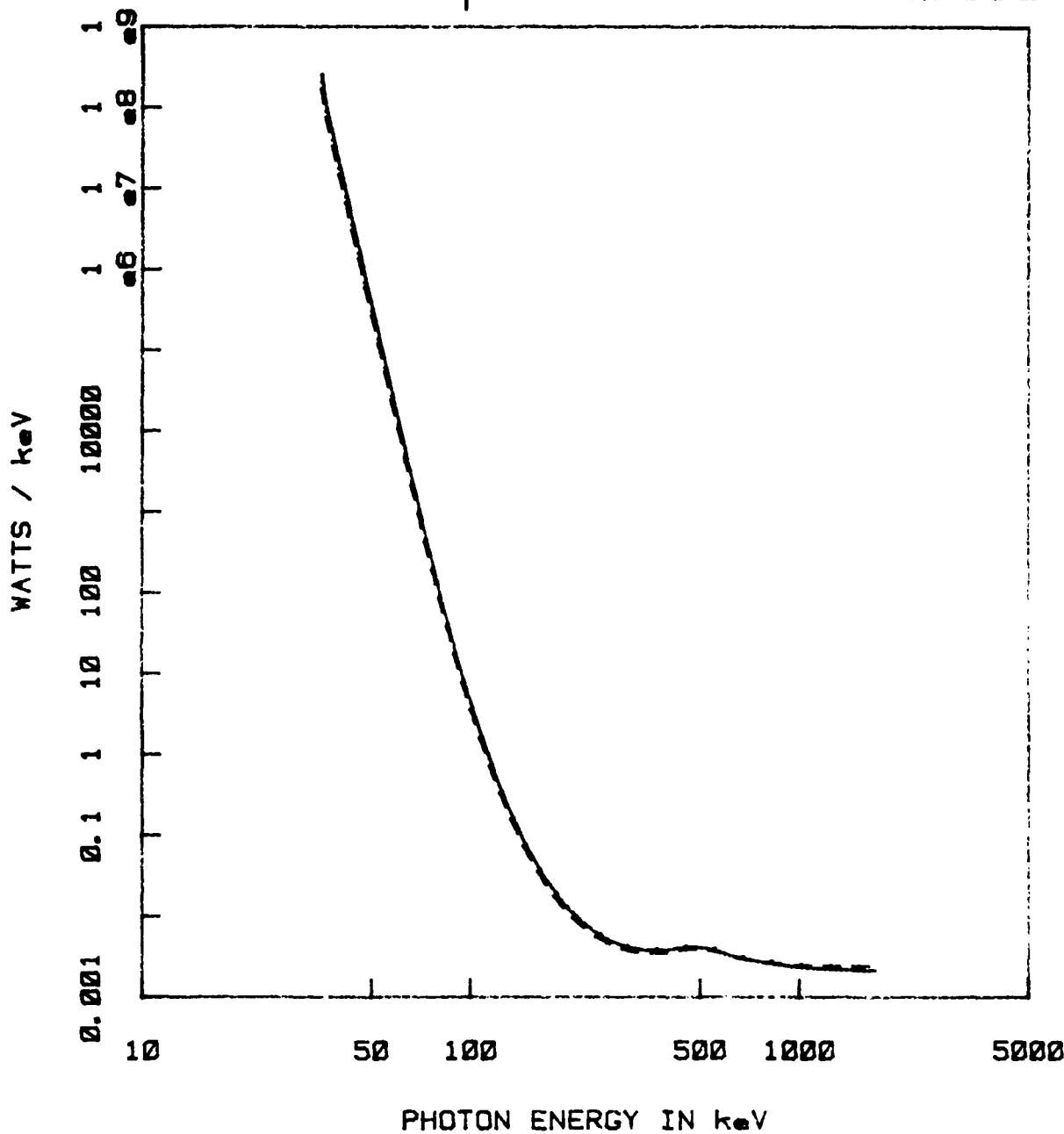
FIGURE 8. PUFF Plasma Gun Spectrum Shot D6-4 Detector Response
Function Varied. This figure shows the effects of varying
the relative detector response function in the program
DCON. The solid line is the same spectrum as in Figure 6,
and used the C8H8 response data. The dashed line used
the flat response. The dotted line (hidden under the
other two spectra) used the linear regression.

Experimental Errors

The sources of experimental errors and their estimated contributions are listed in Table IV. The contribution from the knowledge of the response functions of the detectors is definitely dominant. The contribution is 20% to a figure of 23% overall.

TABLE IV.

Type of Error	Contribution
Error in Reading Trace (Random)	about 3%
Error Due to EMP Change in Bias (Systematic)	about 1 to 2%
Error in Separating EMP Signal From Noise (Random)	0.05 to 1 V
Errors in Scope Calibration (Systematic)	about 1%
Errors in Distance Measurements (Systematic)	less than 0.5%
Error in Knowledge of Response Function (Unfiltered) (Systematic)	about +/- 20%
Error in Knowledge of Filter Transmissions (Systematic)	less than 5 to 10%
Total Error	about 23%

Table IV. List of Experimental Errors

VI. Conclusions And Discussion

Detector Calibration

Figure 5 on page 38 shows that the relative response for the detector was not determined very accurately. Therefore, besides the linear regression fit to the experimental calibration data, two other relative detector response functions were used. They were a flat response and a log-log fit to the mass energy absorption coefficients of polystyrene (C8H8). The polystyrene data from Hubble (7:74) was used because the response of the detector to various energy photons should vary as the mass energy absorption coefficients of the scintillating material varies. As was seen in Figure 8 though, changing the relative response function of the detectors changed the resulting spectrum very little. From this, it is reasonable to conclude further calibration of the dectors to photon energies is probably not necessary. However, it is recommended that the polystyrene data be used for the relative response because it should follow the actual detector responses closer than either the linear regression or the flat response.

Filter Selection

From the result section on filter selection, it would seem that Cu or Fe would make excellent filters. The problem with the Fe (or steel) was availability in a reasonable time frame. The Cu was available, but the problem was one the program DCON does not consider. The problem pertains to the detector signal

strength. The signals from the detectors need to be above one volt or so before the signal is above the electronic noise levels (usually about 50 mV) created by firing the bank. The detector signal should also not be above 20 or 25 volts, because the detector will then start to saturate and the response becomes nonlinear.

When using Cu as a filter material with 1/8 in. for the first filter and a factor of three increase of filter thickness from one detector to the next, there is no problem for the first two filter thicknesses (1/8 in. and 3/8 in). For the next filter thickness, 9/8 in., the signal becomes too low. The dilemma raised is one of how to get various filters of the same material and varying thicknesses, but still get a reasonable signal. This was partially solved by first moving the unfiltered detector to a distance about twice as far away from the source as the other detectors. The filter "multiplication factor" was reduced to two, so three detectors were filtered with 1/8, 1/4, and 1/2 in. of Cu. With one detector dedicated to neutron detection (a two in. Pb filter), and seven detectors available, there were two detectors left that needed filters. Shot D6-4 demonstrated that 1/16 in. of Pb gave a reasonable signal, so that was used for the sixth detector. The seventh detector was then filtered with 1/2 in. of Al because it was available and it was hoped to produce a reasonable signal.

Shot D6-4

PUFF shot D6-4 produced a spectrum that is interesting in two aspects. The first aspect is the spectral output is a power law spectrum between 35 and 100 keV. It is not known by the author the exact significance of this result, but if bremsstrahlung is the major photon radiation production mechanism above 30 keV for PUFF with the coaxial rail gun, then the bremsstrahlung is probably produced by electrons whose energies are also distributed in a power law distribution between 35 and 100 keV.

The other interesting aspect is the small "feature" or "bump" in the spectrum at about 500 keV. The feature was originally thought to be a result of the type of iterative algorithm used because of the "end point error" problem discussed in the results section on filter selection. This was because the energy mesh was only extended up to 400 keV and the energy being radiated at the upper end of the spectrum was increasing. When the energy mesh was enlarged to the size seen, the increase turned out to not be the edge effect it was thought to be.

Figure 7 suggests this feature is due to an anomalous signal from detector 124. This is because the feature disappears when the signal from detector 124 is not included. This signal is anomalous because the signal from detector 124, which is shielded with 1/8 in. of Cu, is as large as the signal from detector 123, which is shielded with 1/4 in. of Al.

Probably the most important result is that the deconvoluted spectrum follows a power law, and the slope is changed by ignoring the anomalous detector signal.

Final Conclusions

Overall, the thesis seems to have been a successful one because a reasonable spectrum was produced by the method that was proposed. Therefore, the problem that was to be solved, that of devising a method to determine the high energy spectrum of SHIVA STAR, was accomplished. There were a few interesting results observed and discussed which hopefully point the way to further research. The only regret is that more time was not available for reducing the data available. It would be very interesting to have the current and voltage data analyzed so that one could check to see that the peak voltage is 100 kV or so the spectrum from shot D6-4 would indicate if the major photon production mechanism above 30 keV is bremsstrahlung. It would also be interesting to deconvolute more of the data. Over one hundred shots were performed, though many used a large magnetic probe array and only one scintillator photomultiplier detector. There also needs to be some further work done in the area of filter selection so that a better idea of an ideal filter set can be determined and used.

APPENDIX A

APPENDIX A

LISTING OF PROGRAM DCON

The following is a listing of the program DCON. As listed
it will execute on a VAX 11/780 under the UNIX operating system.
By changing the two OPEN statements in subroutine DATIN, the
program will execute under the VMS operating system.

Sample input and output for the program are listed in
Appendices B and C.

```
      PROGRAM DCON
C
CCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCC
C  .                                                               C
C      This program deconvolutes the spectrum of a high energy photon   C
C      source.  This program is adopted from the program XSPEC, by      C
C      James H. Degnan,  AFWL/NTYP.  The program DCON was written by     C
C      Charles W. Beason, as part of research for a Masters Thesis       C
C      at the Air Force Inst. of Technology.  The research was           C
C      supported by AFWL/NTYP.                                           C
C                                              Febuary 1984              C
C                                                                        C
CCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCC
C
      COMMON / ALL / TRES , NDET , KIT , KITMAX , E1 , RATIO ,
     1 JLAST2(5) , X(15,12) , LMAT(5) , XMAT(5) , KDET(15) ,
     2 VC(15) , VO(15) , NFILT(15) , MATNUM(15,15) ,
     3 LABLE(2) , THING(5) , IT(3) , RESPSM(15) , CORREC(15) ,
     4 Z(12) , ECOEF(8,100) , COEF(8,100) ,ERES(8,100),RES(8,100) ,
     5 EE(100) , XMU(8,100) , RESPO(8,100) , RESP1(8,100) ,
     6 DE(100) , SPEC(100) , DEL(15,15) , SO ,
     7 CONV(100) , TSPEC(100) , SPEC2(100) , EPRNT(1000) ,
     8 VCPRNT(8,1000) , NSPEC , M ,
     9 JJ1(125) , JJ2(125) , APEA(125) , WIDTH(125) , SBAR(125) ,
     1 JBAR(125) , ISHOT , IDDUG , IFLAG , AA , BB , ND , NE , NP ,
     2 KITP , GAMMA , ERROR , ERMAX ,
     3 IDNUM(8) , DETCAL(8) , ISNUM(8) , DSTCAL(8)
C
C      Perform first initialization...
C
      CALL INIT
C
C      Input data...
C
      CALL DATIN
      IF ( IFLAG .EQ. -2 ) GOTO 799
C
C      Generate bare NFHS4X response tables RES(I,J)...
C
      CALL RESGEN
      IF ( IFLAG .EQ. -3 ) GOTO 799
C
C      Unfiltered responses have been generated,  so now
C      generate standard arrays by log-log interpolation...
C
      CALL SARRAY
      IF ( IFLAG .EQ. 799 ) GOTO 799
C
C      Have now generated compatible standard arrays EE(J),XMU(N,J)
C      and RESPO(I,J) where N=1,NTAB and I=1,ND and J=1,NE .
C      Other useful arrays are RESP1(F,J), DE(J)...
C
C
C      Top of 'time' loop...
C
460   CONTINUE
C
C      Perform initialization needed for each time loop...
C
C      Create initial spectrum
C      Initial spectrum is to start iteration
C      SO.GT.0.0 for flat spectrum SO.LT.0.0 for two temperature
C      black body spectrum...
C
```

```
            CALL SPCGEN
803         continue
C
C           Input observed signals,  VO(I)=(1./50.)*(4.*PI/OMEGA)*(1./TMESH)*
C           (VP  OR  INT(V*DT)) ,VP in VOLTS for SPEC in WATTS/KEV,
C           INT(V*DT) IN VOLT-SEC for SPEC in JOULES/keV
C
            CALL DETINP
            IF ( IFLAG .EQ. 708 ) GOTO 799
C
C           902 / GOTO 902 is iteration/convergence loop...
C
902         CONTINUE
C
C           Generate calculates detector signals...
C
            CALL CALDET
            IF ( IFLAG .EQ. 903 ) GOTO 903
            IF ( IFLAG .EQ. 904 ) GOTO 904
C
C           Smoothing loop...
C
            CALL SMOOTH
C
C           End of iteration/convergence loop...
C
            GO TO 902
C
C
904         CONTINUE
C
C           Convergence to specified max error ERMAX was not achieved...
C
            WRITE(3,299)
299         FORMAT( 1  , // , '                ERROR .GT. ERMAX' , // )
903         CONTINUE
            WRITE(3,1001) IFLAG
C
C           Now output data that has been computed...
C
            CALL FINOUT
C
C           Generate new initial spectrum?
C
            IF ( NSPEC .EQ. 1 ) GOTO 460
C
C           No,  start from end of previous spectrum...
C
            GOTO 803
799         CONTINUE
708         CONTINUE
C
C           All finished,  so close input and output files and stop...
C
            CLOSE(UNIT=1)
            WRITE(3,1000)
            WRITE(3,1001) IFLAG
1000        FORMAT( // ,' *** COMPLETION OF COMPUTATION *** ' // )
1001        FORMAT( ' ********************************** IFLAG=' , I5 )
            CLOSE(UNIT=3)
            STOP
            END
```

```
      SUBROUTINE DETINP
C
CCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCC
C                                                                     C
C      This subroutine inputs observed detector signals              C
C                                                                     C
CCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCC
C
      COMMON / ALL / TRES , NDET , KIT , KITMAX , E1 , RATIO ,
     1 JLAST2(5) , X(15,12) , LMAT(5) , XMAT(5) , KDET(15) ,
     2 VC(15) , VO(15) , NFILT(15) , MATNUM(15,15) ,
     3 LABLE(2) , THING(5) , IT(3) , RESPSM(15) , CORREC(15) ,
     4 Z(12) , ECOEF(8,100) , COEF(8,100) ,EPES(8,100),PES(8,100) ,
     5 EE(100) , XMU(8,100) , RESPO(8,100) , RESP1(8,100) ,
     6 DE(100) , SPEC(100) , DEL(15,15) , SO ,
     7 CONV(100) , TSPEC(100) , SPEC2(100) , EPRNT(1000) ,
     8 VCPRNT(8,1000) , NSPEC , M ,
     9 JJ1(125) , JJ2(125) , AREA(125) , WIDTH(125) , SBAR(125) ,
     1 JBAR(125) , ISHOT , IDBUG , IFLAG , AA , BB , ND , NE , NP ,
     2 KITP , GAMMA , ERROR , ERMAX ,
     3 IDNUM(8) , DETCAL(8) , ISNUM(8) , DSTCAL(8)
C
C      Input observed signals.  VO(I)=(1./50.)*.4.*PI/OMEGA)*(1./TMESH)*
C      (VP   OR   INT(V*DT)) ,VP in VOLTS for SPEC in WATTS/KEV,
C      INT(V*DT) IN VOLT-SEC for SPEC in JOULES/KEV.
C      TRES is time after fiducal(sp?) of response...
C
      READ(1,1,END=708) ( VO(I) , I = 1 , NDET ) , TRES
1     FORMAT(9E16.4)
      WRITE(3,1) ( VO(I) , I = 1 , NDET ) , TRES
      IF ( VO(1) .LE. 0.0E0   GOTO 708
      KIT = 1
      RETURN
708   CONTINUE
      IFLAG = 708
      RETURN
      END
```

```
      SUBROUTINE RESGEN
C
CCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCC
C                                                                     C
C        This subroutine generates the 'bare' (unfiltered) NPM54X     C
C        response function.                                           C
C                                                                     C
CCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCC
C
         COMMON / ALL / TRES , NDET , KIT , KITMAX , E1 , RATIO ,
     1 JLAST2(5) , X(15,12) , LMAT(5) , XMAT(5) , KDET(15) ,
     2 VC(15) , VO(15) , NFILT(15) , MATNUM(15,15) ,
     3 LABLE(2) , THING(5) , IT(3) , RESPSM(15) , CORPEC(15) ,
     4 Z(12) , ECOEF(8,100) , COEF(8,100) ,CRES 8,100),PES(8,100) ,
     5 EE(100) , XMU(8,100) , RESPO(8,100) , RESP1(8,100) ,
     6 DE(100) , SPEC(100) , DEL(15,15) , SO ,
     7 CONV(100) , TSPEC(100) , SPEC2(100) , EPRNT(1000) ,
     8 VCPRNT(8,1000) , NSPEC , M ,
     9 JJ1(125) , JJ2(125) , AREA(125) , WIDTH(125) , SBAP(125) ,
     1 JBAR(125) , ISHOT , IDBUG , IFLAG , AA , BB , ND , NE , NP ,
     2 KITP , GAMMA , ERROR , ERMAX ,
     3 IDNUM(8) , DETCAL(8) , ISNUM(8) , DSTCAL(8)
C
C      Detector serial numbers...
C
      DATA ISNUM /  077 , 078 , 079 , 080 ,
     1              122 , 123 , 124 , 125 /
C
C      Detector response in Amps/(Gamma-MeV/CM**2/SEC)
C
      DATA DSTCAL / 1.88E-10 , 2.00E-10 , 2.49E-10 , 1.71E-10 ,
     1              1.41E-10 , 1.56E-10 , 1.03E-10 , 1.62E-10 /
C
C      Change DSTCAL to keV (multiply by 1 = 1 Mev / 1000 keV)...
C
      IFLAG = 1
      DO 400 I = 1 , 8
      DSTCAL(I) = 0.001 * DSTCAL(I)
400   CONTINUE
C
C      Determine who belongs to who in a detector sort...
C
      DO 300 I = 1 , NDET
      J = 0
250   CONTINUE
      J = J + 1
      IF ( J .GE. 9 ) GOTO 9999
      IF ( IDNUM(I) .NE. ISNUM(J) ) GOTO 250
      DETCAL(I) = DSTCAL(J)
300   CONTINUE
      M = NE - 1
      EE(1)= E1
C
C      RESCAL is relative calibration normalization factor...60Co=1.0...
C
C      VCAL is a function that returns a detector response that needs
C      to be normalized...
C
      RESCAL = VCAL(1253.0,IFLAG)
      IF ( IFLAG .EQ. -1 ) GOTO 9999
      DO 100 J = 1 , NE
      IF ( J .NE. 1 ) EE(J) = EE( J - 1 ) * RATIO
C
C      Calculate relative calibration factor PESi...
```

```
        RES1 = VCAL(EE(J),IFLAG)
        IF ( IFLAG .EQ. -1 ) GOTO 9999
C
C       Normalize relative calibration factor...
C
        RES1 = RES1 / RESCAL
        DO 50 I = 1 , NDET
C
C       RES(I,J) is response of the I'th detector at the energy
C       ERES(I,J).
C
C
C       ERES(I,J) is energy in keV.  RES(I,J) is 'bare' (unfiltered)
C       NPM54X responses in ...
C
        RES(I,J) = RES1 * DETCAL(I)
        RESPO(I,J) = RES(I,J)
        ERES(I,J) = EE(J)
50      CONTINUE
100     CONTINUE
        RETURN
9999    CONTINUE
        IFLAG = -3
        RETURN
        END
```

```fortran
      FUNCTION VCAL(E)
C
CCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCC
C                                                                      C
C     THIS FUNCTION DETERMINES THE RELATIVE RESPONSE OF THE DETECTOR   C
C     TO A GIVEN ENERGY PHOTON.                                        C
C                                                                      C
C     IVCAL...... = 0 ....... THEN USE LINEAR FIT TO CALIBRATION DATA  C
C                  = 1 ....... THEN USE FLAT RESPONSE                  C
C                  = 2 ....... THEN USE MASS ENERGY ABSORPTION COEF'S  C
C                              FROM J.H. HUBBEL, "PHOTON MASS ATTENUATION C
C                      AND MASS ENERGY-ABSORPTION COEFFICIENTS FOR H, C, N, C
C                      O, Ar, AND SEVEN MIXTURES FROM 0.1 keV to 20 MeV", C
C                      RADIATION RESEARCH  70, PAGES 58-81 (1977).     C
C                                                                      C
CCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCC
C
      COMMON / IVIV / IVCAL
      DIMENSION EVCAL(26), RVCAL(26)
C
C     ENERGIES CORRESPONDING TO THE ABSORPTION COEF'S LOADED INTO
C     RVCAL BELOW
C
      DATA EVCAL /
     1    1.0000E1 , 1.5000E1 , 2.0000E1 , 3.0000E1 , 4.0000E1 ,
     2    5.0000E1 , 6.0000E1 , 8.0000E1 , 1.0000E2 , 1.5000E2 ,
     3    2.0000E2 , 3.0000E2 , 4.0000E2 , 5.0000E2 , 6.0000E2 ,
     4    6.6163E2 , 8.0000E2 , 1.0000E3 , 1.1732E3 , 1.2522E3 ,
     5    1.3025E3 , 1.5000E3 , 2.0000E3 , 3.0000E3 , 4.0000E3 ,
     6    5.0000E3 /
C
C     MASS-ENERGY ABSORPTION COEFFICIENTS IN M**2/KG FOR
C     POLYSTYRENE (C8H8) (NE111 C:H RATIO IS 1:1.092)
C
      DATA RVCAL /
     1    1.849E-1 , 5.014E-2 , 2.002E-2 , 6.056E-3 , 3.190E-3 ,
     2    2.387E-3 , 2.153E-3 , 2.152E-3 , 2.292E-3 , 2.631E-3 ,
     3    2.856E-3 , 3.088E-3 , 3.174E-3 , 3.195E-3 , 3.181E-3 ,
     4    3.163E-3 , 3.106E-3 , 3.005E-3 , 2.913E-3 , 2.871E-3 ,
     5    2.828E-3 , 2.744E-3 , 2.522E-3 , 2.196E-3 , 1.978E-3 ,
     6    1.822E-3 /
      IVCP1 = IVCAL + 1
      GOTO ( 100 , 200 , 300 ) , IVCP1
100   CONTINUE
C
C     LINEAR RESPONSE FROM CALIBRATION DATA...
C
      VCAL = 7.67e-13 + E * 1.384e-16
      RETURN
200   CONTINUE
C
C     FLAT RESPONSE...
C
      VCAL = 1.0
      RETURN
300   CONTINUE
C
C     RESPONSES DETERMINED FROM HUBBELL'S DATA...
C     (DATA IS LOG-LOG INTERPOLATED)...
C
C     Perform table look-up...
C
      IFLAG = 1
      J = 1
```

```fortran
      IF ( E .LT. EVCAL(1) ) GOTO 9999
1     CONTINUE
      IF ( E .LT. EVCAL(J) ) GOTO 2
C
C     Test for out of bounds...
C
      IF ( J .GE. 26 ) GOTO 9999
      J = J + 1
      GOTO 1
2     CONTINUE
      JM1 = J - 1
      ENGTOP = ALOG( EVCAL(J) )
      ENGBOT = ALOG( EVCAL(JM1) )
      RHOTOP = ALOG( RVCAL(J) )
      RHOBOT = ALOG( RVCAL(JM1) )
      ENGL = ALOG( E ) - ENGBOT
      SLOPE = ( RHOTOP - RHOBOT ) / ( ENGTOP - ENGBOT )
      VCAL = RHOBOT + ENGL * SLOPE
      VCAL = EXP( VCAL )
      RETURN
9999  CONTINUE
      IFLAG = -1
      RETURN
      END
```

```
      SUBROUTINE FINOUT
C
CCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCC
C                                                                  C
C     This subroutine outputs the computed spectrum when the technique C
C     either converges or fails to converge.                       C
C                                                                  C
CCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCC
C
      COMMON / ALL / TRES , NDET , KIT , KITMAX , E1 , RATIO ,
     1 JLAST2(5) , X(15,12) , LMAT(5) , XMAT(5) , KDET(15) ,
     2 VC(15) , VO(15) , NFILT(15) , MATNUM(15,15) ,
     3 LABLE(2) , THING(5) , IT(3) , RESPSM(15) , CORREC(15) ,
     4 Z(12) , ECOEF(8,100) , COEF(8,100) ,ERES(8,100),RES(8,100) ,
     5 EE(100) , XMU(8,100) , RESPO(8,100) , RESP1(8,100) ,
     6 DE(100) , SPEC(100) , DEL(15,15) , SO ,
     7 CONV(100) , TSPEC(100) , SPEC2(100) , EPRNT(1000) ,
     8 VCPRNT(8,1000) , NSPEC , M ,
     9 JJ1(125) , JJ2(125) , AREA(125) , WIDTH(125) , SBAR(125) ,
     1 JBAR(125) , ISHOT , IDBUG , IFLAG , AA , BB , ND , NE , NP ,
     2 KITP , GAMMA , ERROR , EPMAX ,
     3 IDNUM(8) , DETCAL(8) , ISNUM(8) , DSTCAL(8)
      ITCC = 0
      LL2 = 12
      IF ( NDET .LE. 12) LL2 = NDET
350   CONTINUE
      WRITE(3,351) ( K , K = 1 , LL2 )
351   FORMAT( // , 1X ,' K    ERROR(K)', 17 , 11I10 , // )
352   FORMAT( // , 1X ,' K' , 3I10 , // )
      WRITE(3,891)
891   FORMAT( 1X , // )
      DO 556 K = 1 , KITP
      WRITE(3,303) K , EPRNT(K) , ( VCPRNT(I,K) , I = 1 , LL2 )
556   CONTINUE
      IF ( NDET .LE. 12 ) GOTO 354
      WRITE(3,352) ( K , K = 13 , NDET )
      WRITE(3,891)
      DO 557 K = 1 , KITP
      WRITE(3,353) K , ( VCPRNT(I,K) , I = 13 , NDET )
557   CONTINUE
353   FORMAT( 1X , I5 , 3E10.3 )
303   FORMAT( 1X , I5 , 13E10.3 )
354   CONTINUE
      WRITE(3,297)
297   FORMAT( 1H0 , // , 4X , 'DET' , 7X , 'OBS.  SIGNAL' ,
     1   8X , 'CALC. SIGNAL' ,// )
      DO 56 I = 1 , NDET
      WRITE(3,296) I , VO(I) , VC(I)
296   FORMAT( 1X , I5 , 4X , E15.4 , 5X , E15.4 )
56    CONTINUE
      SSPEC = 0.0E0
      DO 54 J = 1 , NE
      SSPEC = SSPEC + SPEC(J) * DE(J)
      TSPEC(J) = SSPEC
54    CONTINUE
      WRITE(3,304)
304   FORMAT( // , 1X , '      EE(J) (keV)' , 5X , 'SPEC(J)' ,
     1   ' (Watts/keV)' , 3X ,
     1              'TSPEC(J) (Watts)' , // )
      DO 114 J = 1 , NE
C
C     Change from Watts/Joule to Watts/keV...
C
      SPEC(J) = SPEC(J) * 1.602E-16
```

63

```
      TSPEC(J) = TSPEC(J) * 1.602E-16
      WRITE(3,305) EE(J) , SPEC(J) , TSPEC(J)
305   FORMAT( 1X , F10.4 , 2E20.4 )
114   CONTINUE
      RETURN
      END
```

```
      SUBROUTINE CALDET
C
CCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCC
C                                                                    C
C                                                                    C
C      This subroutine generates calculated detector signals from a  C
C      predicted spectrum.                                           C
C                                                                    C
CCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCC
C
       COMMON / ALL / TRES , NDET , KIT , KITMAX , E1 , RATIO ,
      1 JLAST2(5) , X(15,12) , LMAT(5) , XMAT(5) , KDET(15) ,
      2 VC(15) , VO(15) , NFILT(15) , MATNUM(15,15) ,
      3 LABLE(2) , THING(5) , IT(3) , RESPSM(15) , CORREC(15) ,
      4 Z(12) , ECOEF(8,100) , COEF(8,100) ,ERES(8,100),RES(8,100) ,
      5 EE(100) , XMU(8,100) , RESPO(8,100) , RESP1(8,100) ,
      6 DE(100) , SPEC(100) , DEL(15,15) , SO ,
      7 CONV(100) , TSPEC(100) , SPEC2(100) , EPRNT(1000) ,
      8 VCPRNT(8,1000) , NSPEC , M ,
      9 JJ1(125) , JJ2(125) , AREA(125) , WIDTH(125) , SBAR(125) ,
      1 JBAR(125) , ISHOT , IDBUG , IFLAG , AA , BB , ND , NE , NP ,
      2 KITP , GAMMA , ERROR , ERMAX ,
      3 IDNUM(8) , DETCAL(8) , ISNUM(8) , DSTCAL(8)
C
C      NDET is the number of detectors...
C
       DO 106 I = 1 , NDET
C
C      Initialize SUM...
C
       SUM = 0.0E0
C
C      M is NE - 1 (the number of energy groups being computed
C      minus one)
C
       DO 107 J = 1 , M
       DSUM = 0.25E0 * ( RESP1(I,J) + RESP1(I,J+1) ) *
      1       ( SPEC(J) + SPEC(J+1) ) * DE(J)
       SUM = SUM + DSUM
   62  FORMAT(5X,2I5,4E16.8)
  107  CONTINUE
C
C      VC(I) is the computed detector voltage output for
C      this iteration...
C
       VC(I) = SUM
C
C      CORREC(I) is the correction to the detector voltage
C      of the previous guess...
C
       IF ( IDBUG .GE. 4 ) WRITE(6,63) I , VO(I) , VC(I)
   63  FORMAT(5X,I5,2E16.8)
       CORREC(I) = ( VO(I) / VC(I) ) ** GAMMA
  106  CONTINUE
C
C      Calculate error in convergence...
C
       ERROR = 0.0E0
       DO 52 I = 1 , NDET
C
C      Sum errors for each detector...
C
       ERROR = ERROR + ( (VO(I) - VC(I) ) * ( VO(I) - VC(I) )
      1       / ( VO(I) * VO(I) ) )
```

```
52      CONTINUE
C
C       Divide by number of detectors NDET to get error per detector...
C
        ERROR = ERROR / DFLOAT(NDET)
C
C       Store error,  iteration number -vs- detector
C
        EPRNT(KIT) = ERROR
        DO 519 I = 1 , NDET
        VCPRNT(I,KIT) = VC(I)
519     CONTINUE
        KITP = KIT
C
C       KIT is number of iterations.  At end of present iteration,
C       so increment by one...
C
        KIT = KIT + 1
C
C       Check for convergence...
C
        IF ( ERROR .LT. ERMAX ) GOTO 903
C
C       Check to see if iteration limit KITMAX has been reached...
C
        IF ( KIT .GT. KITMAX ) GOTO 904
C
C       Have not reached error or iteration limits,
C       so ...
C
        DO 108 J = 1 , NE
        SUM1 = 0.0E0
        SUM2 = 0.0E0
        DO 109 I = 1 , NDET
        SUM1 = SUM1 + ( RESP1(I,J) / RESPSM(I) ) * CORREC(I)
        SUM2 = SUM2 + ( RESP1(I,J) / RESPSM(I) )
109     CONTINUE
        XCONV = 0.0E0
        IF ( SUM1 .EQ. 0.0E0 ) GOTO 120
        XCONV = ALOG(SUM1) - ALOG(SUM2)
120     CONTINUE
        IF ( XCONV .LT. -82.89E0 ) XCONV = -82.89E0
        IF ( XCONV .GT. 82.89E0 ) XCONV = 82.89E0
        XTEMP = ALOG( SPEC(J) ) + XCONV
        IF ( XTEMP .LT. -82.89E0 ) XTEMP = -82.89E0
        IF ( XTEMP .GT. 82.89E0 ) XTEMP = 82.89E0
        SPEC(J) = EXP(XTEMP)
        IF ( SPEC(J) .LT. 1.0E-36 ) SPEC(J) = 1.0E-36
108     CONTINUE
        RETURN
903     CONTINUE
C
C       Have achieved specified convergence,  so ...
C
        IFLAG = 903
        RETURN
904     CONTINUE
C
C       Have reached iteration limit...
C
        IFLAG = 904
        RETURN
        END
```

```
      SUBROUTINE SMOOTH
C
CCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCC
C                                                                   C
C      This subroutine 'smooths' the spectrum when the energy range C
C      being examined contains an absorption edge.  This is needed  C
C      so the edge does not show up as a part of the spectrum. (The C
C      edge will produce features that are really not there.)       C
C                                                                   C
CCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCC
C
      COMMON / ALL / TRES , NDET , KIT , KITMAX , E1 , RATIO ,
     1 JLAST2(5) , X(15,12) , LMAT(5) , XMAT(5) , KDET(15) ,
     2 VC(15) , VO(15) , NFILT(15) , MATNUM(15,15) ,
     3 LABLE(2) , THING(5) , IT(3) , RESPSM(15) , CORREC(15) ,
     4 Z(12) , ECOEF(8,100) , COEF(8,100) ,ERES(8,100),RES(8,100) ,
     5 EE(100) , XMU(8,100) , RESPO(8,100) , RESP1(8,100) ,
     6 DE(100) , SPEC(100) , DEL(15,15) , SO ,
     7 CONV(100) , TSPEC(100) , SPEC2(100) , EPRNT(1000) ,
     8 VCPRNT(8,1000) , NSPEC , M ,
     9 JJ1(125) , JJ2(125) , AREA(125) , WIDTH(125) , SBAR(125) ,
     1 JBAR(125) , ISHOT , IDBUG , IFLAG , AA , BB , ND , NE , NP ,
     2 KITP , GAMMA , ERROR , ERMAX ,
     3 IDNUM(8) , DETCAL(8) , ISNUM(8) , DSTCAL(8)
C
C      If NP = 0,  then do not perform smoothing...
C
      IF ( NP .EQ. 0 ) GOTO 1107
      NP = 10
      IF ( KIT .GT. 140 ) NP = 5
      IF ( KIT .GT. 180 ) NP = 3
      NCTR = 5
      NCTR2 = 1
      NTEST = ( KIT / NCTR ) * NCTR
      IF ( NTEST .EQ. KIT ) GOTO 1107
      IF ( KIT .GE. ( KITMAX - NCTR2 ) ) GOTO 1107
C
      NMAX = NE / NP
      DO 1201 N = 1 , NMAX
      JJ1(N) = ( ( N - 1 ) * NP ) + 1
      JJ2(N) = N * NP
1201  CONTINUE
      JJ2( NMAX ) = NE
      WIDTH(1) = EE( JJ2(1) ) - ( EE(1) / RATIO )
      DO 1202 N = 1 , NMAX
      IF ( N .GT. 1 ) WIDTH(N) = EE( JJ2(N) ) - EE ( JJ2( N - 1 ) )
      L1 = JJ1(N)
      L2 = JJ2(N)
      ASUM = 0.0E0
      PEXP = 0.0E0
      DO 1203 J = L1 , L2
      ASUM = ASUM + 0.5E0 * ( SPEC(J) + SPEC( J + 1 ) ) * DE(J)
      PEXP = PEXP + ALOG( SPEC(J) )
1203  CONTINUE
      AREA(N) = ASUM
      Y = DFLOAT( L2 - L1 + 1 )
      SBAR(N) = AREA(N) / WIDTH(N)
      IF(( N .NE. 1 ) .AND. ( N .NE. NMAX ) .AND. ( KIT.GT.50 )
     1    .AND. ( PEXP .GE. -207.23E0 ) ) SBAR(N)=EXP(PEXP/Y)
C
C      PEXP for geometric averaging.  ASUM for arithmetic averaging...
C
      EBAR = 0.5E0 * ( EE(L1) + EE( L2 + 1 ) )
      IF ( L2 .EQ. NE ) EBAR = 0.5E0 * ( EE(L1) + EE(NE) )
```

```
              XBAR = ( ALOG ( EBAR / EE(L1) ) / ALOG(RATIO) )
              JBAR(N) = L1 + INT(XBAR)
1202    CONTINUE
              DO 1200  N = 1 , NMAX
              SPEC2( JBAR(N) ) = SBAR(N)
1200    CONTINUE
C
              SPEC2(1) = SPEC(1)
              SPEC2(NE) = SPEC(NE)
              KMAX = NMAX
              DO 1204 K = 1 , KMAX
              K1 = 1
              K2 = NE
              IF ( K .NE. 1 ) K1 = JBAR(K-I)
              IF ( K .NE. KMAX ) K2 = JBAR(K)
              DO 1205 J = K1 , K2
              XTEMP = ALOG( SPEC2(K1) ) + ( ALOG( SPEC2(K2) ) -
1                   ALOG( SPEC2(K1) ) )
2             * ( ALOG( EE(J) ) - ALOG( EE(K1) ) ) /
3                ( ALOG( EE(K2) ) - ALOG( EE(K1) ) )
              IF ( XTEMP .LT. -82.89E0 ) XTEMP = -82.89E0
              IF ( XTEMP .GT. 82.89E0 ) XTEMP = 82.89E0
              SPEC2(J) = EXP(XTEMP)
1205    CONTINUE
1204    CONTINUE
              DO 1206 J = 1 , NE
              SPEC(J) = SPEC2(J)
1206    CONTINUE
1107    CONTINUE
              RETURN
              END
```

```
      SUBROUTINE SPCGEN
C
CCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCC
C                                                                     C
C     This subroutine generates the initial spectrum.                 C
C                                                                     C
CCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCC
C
      COMMON / ALL / TRES , NDET , KIT , KITMAX , E1 , RATIO ,
     1 JLAST2(5) , X(15,12) , LMAT(5) , XMAT(5) , KDET(15) ,
     2 VC(15) , VO(15) , NFILT(15) , MATNUM(15,15) ,
     3 LABLE(2) , THING(5) , IT(3) , RESPSM(15) , CORREC(15) ,
     4 Z(12) , ECOEF(8,100) , COEF(8,100) ,ERES(8,100),RES(8,100) ,
     5 EE(100) , XMU(8,100) , RESPO(8,100) , RESP1(8,100) ,
     6 DE(100) , SPEC(100) , DEL(15,15) , SO ,
     7 CONV(100) , TSPEC(100) , SPEC2(100) , EPRNT(1000) ,
     8 VCPRNT(8,1000) , NSPEC , M ,
     9 JJ1(125) , JJ2(125) , AREA(125) , WIDTH(125) , SBAR(125) ,
     1 JBAR(125) , ISHOT , IDBUG , IFLAG , AA , BB , ND , NE , NP ,
     2 KITP , GAMMA , ERROR , ERMAX ,
     3 IDNUM(8) , DETCAL(8) , ISNUM(8) , DSTCAL(8)
C
C     If SO < 0.  then generate blackbody as initial spectrum.
C     If not,  then use flat spectrum as initial spectrum.
C
      IF ( SO .GT. 0.0E0 ) GOTO 801
C
C     T1 and T2 are kT in keV...
C     AA and BB are
      DO 2 J = 1 , NE
C
C     SPEC(J) is initial two temperature spectrum for
C     temperature = EE(J)...
C
      SPEC(J) = AA * EXP( T1 / EE(J) ) + BB * EXP( T2 / EE(J) )
2     CONTINUE
      RETURN
801   CONTINUE
C
C     Flat initial spectrum chosen...
C
      DO 851 J = 1 , NE
      SPEC(J) = SO
851   CONTINUE
      RETURN
      END
```

```
      SUBROUTINE SARRAY
C
CCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCC
C                                                                     C
C     This subroutine calculates the detector responses and the sums  C
C     Rk(E) and R(E)                                                   C
C                                                                     C
CCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCC
C
      COMMON / ALL / TRES , NDET , KIT , KITMAX , E1 , RATIO ,
     1 JLAST2(5) , X(15,12) , LMAT(5) , XMAT(5) , KDET(15) ,
     2 VC(15) , VO(15) , NFILT(15) , MATNUM(15,15) ,
     3 LABLE(2) , THING(5) , IT(3) , RESPSM(15) , CORREC(15) ,
     4 Z(12) , ECOEF(8,100) , COEF(8,100) ,ERES(8,100),RES(8,100) ,
     5 EE(100) , XMU(8,100) , RESPO(8,100) , RESP1(8,100) ,
     6 DE(100) , SPEC(100) , DEL(15,15) , SO ,
     7 CONV(100) , TSPEC(100) , SPEC2(100) , EPRNT(1000) ,
     8 VCPRNT(8,1000) , NSPEC , M ,
     9 JJ1(125) , JJ2(125) , AREA(125) , WIDTH(125) , SBAR(125) ,
     1 JBAR(125) , ISHOT , IDBUG , IFLAG , AA , BB , ND , NE , NP ,
     2 KITP , GAMMA , ERROR , ERMAX ,
     3 IDNUM(8) , DETCAL(8) , ISNUM(8) , DSTCAL(8)
      DO 1 J = 2 , NE
C
C     DE is the 'width' of each energy group...
C
      DE( J - 1 ) = EE(J) - EE( J - 1 )
1     CONTINUE
      DE(NE) = DE( NE - 1 )
C
C     Loop over each energy group...
C
      DO 5 J = 1 , NE
C
C     ENG is the energy of group J
C
      ENG = EE(J)
C
C     Loop over each detector...
C
      DO 5 I = 1 , NDET
C
C     Loop over the filters for each detector...
C
      SUM = 0.0E0
      DO 3 KFILT = 1 , NFILT(I)
C
C     MAT is the material number of filter KFILT of detector I...
C
      MAT = MATNUM(I,KFILT)
C
C     DELT is the thickness in cm of filter KFILT of detector I...
C
      DELT = DEL(I,KFILT)
      CALL CROSS(MAT,ENG,CRO,DELT,IFLAG,IDBUG)
      IF ( IFLAG .EQ. -1 ) GOTO 999
C
C     CRO is total cross section of filter KFILT of detector I
C     times thickness DELT times the density of filter KFILT...
C
      SUM = SUM + CRO
C
C     SUM is the sumation of all the cross sections of all the filters
C     for detector I at energy ENG...
```

```
C
3      CONTINUE
C
C      XMU(I,J) is the sumation of all the cross sections of all
C      the filters for detector I at energy ENG...
C
       XMU(I,J) = SUM
5      CONTINUE
       DO 10 I = 1 , NDET
       DO 10 J = 1 , NE
C
C      RESP1(I,J) is the response of detector I at energy EE(J)...
C
       RESP1(I,J) = RESPO(I,J) * EXP( - XMU(I,J) )
10     CONTINUE
       DO 102 I = 1 , NDET
       RSUM = 0.0E0
       DO 101 J = 1 , M
       RSUM = RSUM + 0.5E0 * ( RESP1(I,J) + RESP1( I , J + 1 ) ) * DE(J)
101    CONTINUE
       RESPSM(I) = RSUM
102    CONTINUE
       CALL BEGOUT
       RETURN
999    CONTINUE
       IFLAG = 799
       RETURN
       END
```

```fortran
      SUBROUTINE CROSS(MAT,ENG,CRO,DEL,IFLAG,IDBUG)
      DIMENSION DENSTY(5)
      DATA JFLAG / 1 /
C
C     Material densities in gm/cm**3
C
      DATA DENSTY / 11.35E0 , 2.6989E0 , 7.874E0 ,
     1              8.96E0 , 1.29E-3 /
C
C     Convert air density to altitude in Albuquerque (83% of sea-level)
C
      IF ( JFLAG .EQ. 1 ) DENSTY(5) = DENSTY(5) * 0.83E0
      JFLAG = JFLAG + 1
C
C     Get RHO, which is the cross section of material MAT in gm/cm**2
C     at energy ENG which is in keV.
C
      CALL CSECT(MAT,ENG,RHO,IFLAG,IDBUG)
C
C     If IFLAG = -1, then cross section was not computed, so exit...
C
      IF ( IFLAG .EQ. -1 ) GOTO 9999
C
C     CRO is unitless, and is the density of the filter (DENSTY) in
C     gm/cm**3 times the thickness of the filter (DEL) in cm times
C     the material cross section (RHO) in cm**2/gm
C
      CRO = DENSTY(MAT) * DEL * RHO
      RETURN
9999  CONTINUE
      WRITE(6,9998) MAT , ENG
9998  FORMAT(' IFLAG = -1, MAT=',I2,' ENG=',E16.8)
      STOP
      END
```

```
      SUBROUTINE BEGOUT
C
CCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCC
C                                                                 C
C       This subroutine outputs data computed once (detector responses,  C
C       shot number, etc.)                                        C
C                                                                 C
CCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCC
C
      COMMON / ALL / TRES , NDET , KIT , KITMAX , E1 , RATIO ,
     1 JLAST2(5) , X(15,12) , LMAT(5) , XMAT(5) , KDET(15) ,
     2 VC(15) , VO(15) , NFILT(15) , MATNUM(15,15) ,
     3 LABLE(2) , THING(5) , IT(3) , RESPSM(15) , CORREC(15) ,
     4 Z(12) , ECOEF(8,100) , COEF(8,100) ,ERES(8,100),RES(8,100) ,
     5 EE(100) , XMU(8,100) , RESPO(8,100) , RESP1(8,100) ,
     6 DE(100) , SPEC(100) , DEL(15,15) , SO ,
     7 CONV(100) , TSPEC(100) , SPEC2(100) , EPRNT(1000) ,
     8 VCPRNT(8,1000) , NSPEC , M ,
     9 JJ1(125) , JJ2(125) , AREA(125) , WIDTH(125) , SBAR(125) ,
     1 JBAR(125) , ISHOT , IDBUG , IFLAG , AA , BB , ND , NE , NP ,
     2 KITP , GAMMA , ERROR , ERMAX ,
     3 IDNUM(8) , DETCAL(8) , ISNUM(8) , DSTCAL(8)
      WRITE(3,93) ISHOT
93    FORMAT( 1H1 , // , ' ISHOT' , 1X , I4 )
      WRITE(3,508)
508   FORMAT( // , 1X , 50X , ' RESPONSE IN VOLTS/Joule' , / )
450   FORMAT( ' DETECTOR' , I7 , 11I10 , / )
520   FORMAT( / , ' E(keV)' , // )
      ITCC = 0
      LL1 = 1
      LL2 = 15
      IF ( NDET .LE. 12 ) LL2 = NDET
515   CONTINUE
      WRITE(3,450) ( K , K = LL1 , LL2 )
      WRITE(3,520)
      DO 420 J = 1 , NE
      WRITE(3,509) EE(J) , ( RESP1(K,J) , K = LL1 , LL2 )
509   FORMAT( 1X , 13E10.3 )
420   CONTINUE
      IF ( NDET .LE. 12 .OR. ITCC .EQ. 1 ) GOTO 516
      LL1 = 13
      LL2 = NDET
      ITCC = 1
      GOTO 515
516   CONTINUE
      RETURN
      END
```

73

```fortran
      SUBROUTINE CSECT(MAT,ENGK,RHO,IFLAG,IDBUG)
C
CCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCC
C                                                                    C
C      This subroutine returns the cross section for the material    C
C      specified at the energy specified.  The photon energy (ENGK)  C
C      must be provided in keV.  The material is selected by the     C
C      variable MAT.  Where:                                         C
C                       MAT = 1.....Lead                             C
C                           = 2.....Aluminum                         C
C                           = 3.....Iron                             C
C                           = 4.....Copper                           C
C                           = 5.....Air (NBS Handbook 85(1964) Composition  C
C                                                                    C
C      The data is taken from:                                       C
C      Hubbel, J. H. 'Photon Cross Sections, Attenuation Coeficients, C
C      and Energy Absorption Coefficients From 10 keV to 100 GeV'    C
C      NSRDS-NBS 29 ( Augest 1969 ).  The data is loaded into the    C
C      arrays CROSS and ENERGY in data statements in subroutine INIT. C
C                                                                    C
CCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCC
C
      COMMON / CCCC / CROSS(30,5) , ENERGY(30,5)
C
C      Change input photon energy from keV to MeV...
C
      ENG = ENGK * 0.001E0
C
C      Perform table look-up...
C
      IFLAG = 1
      J = 1
      IF ( ENG .LT. ENERGY(1,MAT) ) GOTO 9999
1     CONTINUE
      IF ( ENG .LT. ENERGY(J,MAT) ) GOTO 2
C
C      Test for out of bounds...
C
      IF ( J .GE. 30 ) GOTO 9999
      J = J + 1
      GOTO 1
2     CONTINUE
      JM1 = J - 1
      ENGTOP = ALOG( ENERGY(J,MAT) )
      ENGBOT = ALOG( ENERGY(JM1,MAT) )
      RHOTOP = ALOG( CROSS(J,MAT) )
      RHOBOT = ALOG( CROSS(JM1,MAT) )
      ENGL = ALOG( ENG ) - ENGBOT
      SLOPE = ( RHOTOP - RHOBOT ) / ( ENGTOP - ENGBOT )
      RHO = RHOBOT + ENGL * SLOPE
      RHO = EXP( RHO )
      RETURN
9999  CONTINUE
      IFLAG = -1
      RETURN
      END
```

74

```
      SUBROUTINE INIT
C
CCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCC
C                                                                   C
C      This subroutine initializes the cross sections               C
C                                                                   C
CCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCC
C
       COMMON / ALL / TRES , NDET , KIT , KITMAX , E1 , RATIO ,
      1 JLAST2(5) , X(15,12) , LMAT(5) , XMAT(5) , KDET(15) ,
      2 VC(15) , VO(15) , NFILT(15) , MATNUM(15,15) ,
      3 LABLE(2) , THING(5) , IT(3) , RESPSM(15) , CORREC(15) ,
      4 Z(12) , ECOEF(8,100) , COEF(8,100) ,ERES(8,100),RES(8,100) ,
      5 EE(100) , XMU(8,100) , RESPO(8,100) , RESP1(8,100) ,
      6 DE(100) , SPEC(100) , DEL(15,15) , SO ,
      7 CONV(100) , TSPEC(100) , SPEC2(100) , EPRNT(1000) ,
      8 VCPRNT(8,1000) , NSPEC , M ,
      9 JJ1(125) , JJ2(125) , AREA(125) , WIDTH(125) , SBAR(125) ,
      1 JBAR(125) , ISHOT , IDBUG , IFLAG , AA , BB , ND , NE , NP ,
      2 KITP , GAMMA , ERROR , ERMAX ,
      3 IDNUM(8) , DETCAL(8) , ISNUM(8) , DSTCAL(8)
       COMMON / CCCC / CROSS(30,5) , ENERGY(30,5)
       DIMENSION CROSS1(30,2),CROSS2(30,3),ENERG1(30,2),ENERG2(30,3)
       EQUIVALENCE (CROSS(1,1),CROSS1(1,1)),(CROSS(1,3),CROSS2(1,1)),
      1   (ENERGY(1,1),ENERG1(1,1)),(ENERGY(1,3),ENERG2(1,1))
C
C      Total cross sections (in cm**2/gm)
C
       DATA CROSS1 /
C
C      Lead,  MAT = 1
C
      1  1.33E+2 , 6.78E+1 , 1.66E+2 , 1.15E+2 , 1.12E+5 , 1.46E+2 ,
      2  1.30E+0 , 1.57E+2 , 8.57E+1 , 2.97E+1 , 1.40E+1 , 7.81E+0 ,
      3  4.87E+0 , 2.33E+0 , 1.03E+0 , 7.45E+0 , 5.40E+0 , 1.97E+0 ,
      4  9.91E-1 , 4.04E-1 , 2.31E-1 , 1.61E-1 , 1.25E-1 , 8.85E-2 ,
      5  7.08E-2 , 5.17E-2 , 4.55E-2 , 4.18E-2 , 4.16E-2 , 4.24E-2 ,
C
C      Aluminum,  MAT = 2
C
      1  2.63E+1 , 7.93E+0 , 3.41E+0 , 1.12E+0 , 5.67E-1 , 3.69E-1 ,
      2  2.80E-1 , 2.03E-1 , 1.71E-1 , 1.38E-1 , 1.22E-1 , 1.04E-1 ,
      3  9.27E-2 , 8.44E-4 , 7.80E-2 , 6.84E-2 , 6.13E-2 , 5.00E-2 ,
      4  4.32E-2 , 3.54E-2 , 3.11E-2 , 2.84E-2 , 8 * 0.0 /
       DATA CROSS2 /
C
C      Iron,  MAT = 3
C
      1  1.73E+2 , 5.64E+1 , 2.55E+1 , 8.13E+0 , 3.62E+0 , 1.94E+0 ,
      2  1.20E+0 , 5.95E-1 , 3.70E-1 , 1.96E-1 , 1.46E-1 , 1.10E-1 ,
      3  9.40E-2 , 8.40E-2 , 7.69E-2 , 6.69E-2 , 5.99E-2 , 4.86E-2 ,
      4  4.25E-2 , 3.62E-2 , 3.31E-2 , 3.14E-2 , 8 * 0.0 ,
C
C      Copper,  MAT = 4
C
      1  2.24E+2 , 7.42E+1 , 3.25E+1 , 1.09E+1 , 4.89E+0 , 2.62E+0 ,
      2  1.62E+0 , 7.72E-1 , 4.61E-1 , 2.23E-1 , 1.57E-1 , 1.12E-1 ,
      3  9.41E-2 , 8.36E-2 , 7.62E-2 , 6.60E-2 , 5.89E-2 , 4.80E-2 ,
      4  4.20E-2 , 3.60E-2 , 3.32E-2 , 3.18E-2 , 8 * 0.0 ,
C
C      Air,  MAT = 5
C
      1  4.99E+0 , 1.55E+0 , 7.52E-1 , 3.49E-1 , 2.48E-1 , 2.06E-1 ,
      2  1.88E-1 , 1.67E-1 , 1.54E-1 , 1.36E-1 , 1.23E-1 , 1.07E-1 ,
```

```
   3  9.54E-2 , 8.70E-2 , 8.05E-2 , 7.07E-2 , 6.36E-2 , 5.18E-2 ,
   4  4.45E-2 , 3.58E-2 , 3.08E-2 , 2.75E-2 , 8 * 0.0 /
C
C
C     Table energies in Mev
C
      DATA ENERG1 /
C
C     Lead
C
   1  1.00E-2 , 1.30E-2 , 1.50E-2 , 1.52E-2 , 1.52E-2 , 1.59E-2 ,
   2  1.59E-2 , 1.59E-2 , 2.00E-2 , 3.00E-2 , 4.00E-2 , 5.00E-2 ,
   3  6.00E-2 , 8.00E-2 , 8.80E-2 , 8.80E-2 , 1.00E-1 , 1.50E-1 ,
   4  2.00E-1 , 3.00E-1 , 4.00E-1 , 5.00E-1 , 6.00E-1 , 8.00E-1 ,
   5  1.00E+0 , 1.50E+0 , 2.00E+0 , 3.00E+0 , 4.00E+0 , 5.00E+0 ,
C
C     Aluminum,  MAT = 2
C
   1  1.00E-2 , 1.50E-2 , 2.00E-2 , 3.00E-2 , 4.00E-2 , 5.00E-2 ,
   2  6.00E-2 , 8.00E-2 , 1.00E-1 , 1.50E-1 , 2.00E-1 , 3.00E-1 ,
   3  4.00E-1 , 5.00E-1 , 6.00E-1 , 8.00E-1 , 1.00E+0 , 1.50E+0 ,
   4  2.00E+0 , 3.00E+0 , 4.00E+0 , 5.00E+0 , 8 * 0.0 /
C
C     Iron ,  MAT = 3
C
      DATA ENERG2 /
   1  1.00E-2 , 1.50E-2 , 2.00E-2 , 3.00E-2 , 4.00E-2 , 5.00E-2 ,
   2  6.00E-2 , 8.00E-2 , 1.00E-1 , 1.50E-1 , 2.00E-1 , 3.00E-1 ,
   3  4.00E-1 , 5.00E-1 , 6.00E-1 , 8.00E-1 , 1.00E+0 , 1.50E+0 ,
   4  2.00E+0 , 3.00E+0 , 4.00E+0 , 5.00E+0 , 8 * 0.0 ,
C
C     Copper,  MAT = 4
C
   1  1.00E-2 , 1.50E-2 , 2.00E-2 , 3.00E-2 , 4.00E-2 , 5.00E-2 ,
   2  6.00E-2 , 8.00E-2 , 1.00E-1 , 1.50E-1 , 2.00E-1 , 3.00E-1 ,
   3  4.00E-1 , 5.00E-1 , 6.00E-1 , 8.00E-1 , 1.00E+0 , 1.50E+0 ,
   4  2.00E+0 , 3.00E+0 , 4.00E+0 , 5.00E+0 , 8 * 0.0 ,
C
C     Air,  MAT = 5
C
   1  1.00E-2 , 1.50E-2 , 2.00E-2 , 3.00E-2 , 4.00E-2 , 5.00E-2 ,
   2  6.00E-2 , 8.00E-2 , 1.00E-1 , 1.50E-1 , 2.00E-1 , 3.00E-1 ,
   3  4.00E-1 , 5.00E-1 , 6.00E-1 , 8.00E-1 , 1.00E+0 , 1.50E+0 ,
   4  2.00E+0 , 3.00E+0 , 4.00E+0 , 5.00E+0 , 8 * 0.0 /
      DATA GAMMA / 1.2E0 /
      RETURN
      END
```

```
      SUBROUTINE DATIN
C
CCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCC
C                                                                    C
C      This subroutine reads in the detector responses and other data C
C      needed for the deconvolution.                                  C
C                                                                    C
CCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCC
C
      COMMON / ALL / TRES , NDET , KIT , KITMAX , E1 , RATIO ,
     1 JLAST2(5) , X(15,12) , LMAT(5) , XMAT(5) , KDET(15) ,
     2 VC(15) , VO(15) , NFILT(15) , MATNUM(15,15) ,
     3 LABLE(2) , THING(5) , IT(3) , RESPSM(15) , CORREC(15) ,
     4 Z(12) , ECOEF(8,100) , COEF(8,100) ,ERES(8,100),RES(8,100) ,
     5 EE(100) , XMU(8,100) , RESPO(8,100) , RESP1(8,100) ,
     6 DE(100) , SPEC(100) , DEL(15,15) , SO ,
     7 CONV(100) , TSPEC(100) , SPEC2(100) , EPRNT(1000) ,
     8 VCPRNT(8,1000) , NSPEC , M ,
     9 JJ1(125) , JJ2(125) , AREA(125) , WIDTH(125) , SBAR(125) ,
     1 JBAR(125) , ISHOT , IDBUG , IFLAG , AA , BB , ND , NE , NP ,
     2 KITP , GAMMA , ERROR , ERMAX ,
     3 IDNUM(8) , DETCAL(8) , ISNUM(8) , DSTCAL(8)
      COMMON / IVIV / IVCAL
      CHARACTER*20 FILNAM
      WRITE(6,60)
60    FORMAT(' Enter debug code:')
      READ(5,61) IDBUG
61    FORMAT(I2)
C
C     Get functional form of response function...
C
      WRITE(6,*) "Enter Response Function Flag"
      READ(5,*) IVCAL
C
C     Get name of file detector response data is in...
C
      WRITE(6,1)
1     FORMAT(' Enter name of file detector responses are in:')
      READ(5,2) FILNAM
2     FORMAT(A20)
C
C     Open file detector responses are in...
C
      OPEN( UNIT = 1 , FILE = FILNAM , STATUS = 'OLD' )
      REWIND 1
C
C     Get name of output file...
C
      WRITE(6,3)
3     FORMAT(' Enter name of output file:')
      READ(5,2) FILNAM
C
C     Open output file...
C
C     Unit 3 is the output file being used...
C
      OPEN( UNIT = 3 , FILE = FILNAM , STATUS = 'NEW' )
C
C     ISHOT is ISHOT number for the experiment the data was taken from...
C
      READ(1,5) ISHOT
5     FORMAT(I4)
      IF ( ISHOT .EQ. 0 ) GOTO 799
      WRITE(3,7)
```

```
7         FORMAT( 1H1 , 25X , ' BEGIN NPM54X SPECTRUM UNFOLD CODE' , // )
C
C         Read in how many detectors were used...
C
          READ(1,10) NDET
10        FORMAT(I2)
C
C         Loop over the number of detectors
C
          DO 100 I = 1 , NDET
C
C         Read in detector serial numbers...
C
          READ(1,12) IDNUM(I)
12        FORMAT(I3)
C
C         Read number of filters used for detector I...
C
          READ(1,10) NFILT(I)
C
C         Loop over the number of filters for detector I...
C
          DO 20 J = 1 , NFILT(I)
C
C         Read in material number and thickness in centimeters for
C         filter J of detector I...
C
          READ(1,15) MATNUM(I,J) , DEL(I,J)
15        FORMAT(I1,E16.8)
20        CONTINUE
100       CONTINUE
C
C         Input initial spectrum.  iteration/convergence parameters
C
          READ(1,200) SO , NP , ERMAX , KITMAX , NSPEC , T1 , T2
C
C         SO is value for initial flat spectrum.  if < 0 then
C         two temperature exponential is used instead.
C         NP is
C         ERMAX is average error allowed per detector for the
C         completion of the convergence
C         KITMAX is the maximum number of iterations allowed
C         T1 and T2 are the temperatures to be used for the
C         two temperature exponential initial iteration
C
200       FORMAT( E16.8 , I5 , E16.8 , 2I5 , 2E16.8 )
          WRITE(3,210) NDET , SO , NP , NSPEC , ERMAX , T1 , T2
210       FORMAT( 1H0 , // , 1X , 'ND=' , I5 , / , 1X , 'SO=' , E14.4 , / ,
         1           1X , 'NP=' , I5 , ' NSPEC=', I2 , / , 1X ,
         2           'ERMAX=' , E11.4 , / , 1X ,
         3           'T1=' , E16.8 , ' keV      T2=' , E16.8 ,' keV' , // )
C
C         Input energy limits/number of groups/etc...
C
          READ(1,220) NE , E1 , RATIO , AA , BB
220       FORMAT(I5,4E16.8)
          WRITE(3,230) NE , E1 , RATIO , AA , BB
230       FORMAT( ' NE=' , I5 , ' RATIO=' , E16.8 , ' keV      E2=' , E16.8 ,
         1           ' AA=' , E16.8 , ' BB=' , E16.8 )
          RETURN
799       CONTINUE
          IFLAG = -2
          END
```

APPENDIX B

Appendix B

Sample Input for Program DCON

The listing on the next page is sample input for the program
DCON. This input was generated by the program PIN, which is
listed in Appendix D. The input for program PIN to generate this
data is listed in Appendix E. Using this data as input for
program DCON will generate the output listed in Appendix C.

```
604
  5
125
  3
    3   .95249999e+00
    5   .36555200e+03
    2   .63199999e+00
 77
  3
    3   .95249999e+00
    5   .35687000e+03
    4   .31750000e+00
122
  3
    3   .95249999e+00
    5   .35014001e+03
    1   .15675000e+00
123
  4
    3   .95249999e+00
    5   .35077499e+03
    4   .31750000e+00
    2   .63199399e+00
124
  3
    3   .95249999e+00
    5   .35972751e+03
    4   .31750000e+00
100   .10000000e+01        .10000000e-05   1000
      .35000000e+02        .10000000e+01           .10000000e+00   .00000000e+00   .00000000e+00
      .3920e+05            .5300e+04        .2051e+04   .6000e+04   .00000000e+00   .00000000e+00
                                                                   .6000e+04   .6899e+04   .0000e+00
```

APPENDIX C

APPENDIX C

Sample Output for Program DCON

This is a listing of sample output generated by program DCON. This output was generated by using the sample input data listed in Appendix B.

BEGIN NPM54X SPECTRUM UNFOLD CODE

ND=   5
SO=   .1000e+01
NP=   1 NSPEC= 1
ERMAX= .1000e-05
T1=   .00000000e+00 keV   T2=   .00000000e+00 keV

NE=   100 RATIO=   .35000000e+02 keV   E2=   .10400000e+01

ISHOT 604   AA=   .00000000e+00 BB=   .00000000e+00

RESPONSE IN VOLTS/Joule

DETECTOR	1	2	3	4	5
E(keV)					
.350e+02	.390e-30	.296e-38	.000e+00	.000e+00	.000e+00
.364e+02	.251e-28	.137e-35	.000e+00	.336e-36	.118e-35
.379e+02	.104e-26	.330e-33	.000e+00	.905e-34	.286e-33
.394e+02	.289e-25	.446e-31	.263e-36	.135e-31	.387e-31
.409e+02	.564e-24	.364e-29	.615e-34	.119e-29	.315e-29
.426e+02	.807e-23	.188e-27	.832e-32	.660e-2	.163e-27
.443e+02	.873e-22	.644e-26	.682e-30	.240e-26	.558e-26
.461e+02	.735e-21	.152e-24	.357e-28	.600e-25	.131e-24
.479e+02	.495e-20	.256e-23	.125e-26	.107e-23	.222e-23
.498e+02	.272e-19	.321e-22	.303e-25	.141e-22	.278e-22
.518e+02	.118e-18	.282e-21	.499e-24	.128e-21	.244e-21
.539e+02	.439e-18	.197e-20	.620e-23	.931e-21	.171e-20
.560e+02	.144e-17	.114e-19	.602e-22	.556e-20	.988e-20
.583e+02	.418e-17	.553e-19	.467e-21	.278e-19	.479e-19
.606e+02	.108e-16	.226e-18	.292e-20	.117e-18	.196e-18
.630e+02	.246e-16	.777e-18	.148e-19	.409e-18	.673e-18
.656e+02	.519e-16	.238e-17	.641e-19	.128e-17	.206e-17
.682e+02	.103e-15	.656e-17	.243e-18	.359e-17	.569e-17
.709e+02	.191e-15	.165e-16	.812e-18	.918e-17	.143e-16
.737e+02	.335e-15	.380e-16	.242e-17	.215e-16	.329e-16
.767e+02	.561e-15	.811e-16	.654e-17	.467e-16	.703e-16
.798e+02	.895e-15	.161e-15	.161e-16	.943e-16	.140e-15
.829e+02	.132e-14	.285e-16	.349e-16	.168e-15	.247e-15
.863e+02	.187e-14	.478e-15	.707e-16	.286e-15	.414e-15
.897e+02	.260e-14	.770e-15	.871e-20	.464e-15	.667e-15
.933e+02	.351e-14	.119e-14	.388e-19	.726e-15	.103e-14
.970e+02	.464e-14	.170e-14	.151e-18	.110e-14	.155e-14
.101e+03	.592e-14	.253e-14	.511e-18	.157e-14	.220e-14
.105e+03	.711e-14	.330e-14	.148e-17	.206e-14	.266e-14
.109e+03	.845e-14	.424e-14	.369e-17	.266e-14	.367e-14
.114e+03	.990e-14	.535e-14	.944e-17	.338e-14	.464e-14
.118e+03	.116e-13	.668e-14	.212e-16	.403e-14	.579e-14
.123e+03	.135e-13	.822e-14	.444e-16	.524e-14	.712e-14
.128e+03	.155e-13	.999e-14	.874e-16	.641e-14	.866e-14
.133e+03	.177e-13	.120e-13	.162e-15	.774e-14	.104e-13
.138e+03	.200e-13	.143e-13	.286e-15	.926e-14	.124e-13
.144e+03	.226e-13	.168e-13	.481e-15	.110e-13	.146e-13
.149e+03	.253e-13	.196e-13	.773e-15	.129e-13	.170e-13
.155e+03	.273e-13	.218e-13	.115e-14	.143e-13	.189e-13

84

.162e+03	.293e-13	.240e-13	.164e-14	.158e-13	.208e-13
.168e+03	.314e-13	.263e-13	.229e-14	.174e-13	.228e-13
.175e+03	.336e-13	.288e-13	.311e-14	.191e-13	.249e-13
.182e+03	.359e-13	.314e-13	.412e-14	.209e-13	.272e-13
.189e+03	.382e-13	.341e-13	.534e-14	.228e-13	.295e-13
.197e+03	.406e-13	.369e-13	.679e-14	.248e-13	.320e-13
.204e+03	.427e-13	.393e-13	.835e-14	.265e-13	.341e-13
.213e+03	.444e-13	.414e-13	.998e-14	.280e-13	.359e-13
.221e+03	.462e-13	.435e-13	.118e-13	.295e-13	.377e-13
.230e+03	.481e-13	.457e-13	.138e-13	.311e-13	.396e-13
.239e+03	.499e-13	.480e-13	.159e-13	.327e-13	.416e-13
.249e+03	.518e-13	.502e-13	.182e-13	.344e-13	.436e-13
.259e+03	.537e-13	.526e-13	.206e-13	.361e-13	.456e-13
.269e+03	.557e-13	.550e-13	.231e-13	.379e-13	.477e-13
.280e+03	.577e-13	.574e-13	.258e-13	.397e-13	.498e-13
.291e+03	.597e-13	.599e-13	.286e-13	.415e-13	.519e-13
.303e+03	.616e-13	.623e-13	.313e-13	.433e-13	.540e-13
.315e+03	.632e-13	.641e-13	.338e-13	.447e-13	.556e-13
.327e+03	.647e-13	.660e-13	.362e-13	.461e-13	.572e-13
.340e+03	.663e-13	.679e-13	.387e-13	.475e-13	.588e-13
.354e+03	.678e-13	.698e-13	.412e-13	.490e-13	.605e-13
.368e+03	.694e-13	.717e-13	.437e-13	.505e-13	.621e-13
.383e+03	.710e-13	.736e-13	.462e-13	.520e-13	.638e-13
.398e+03	.726e-13	.756e-13	.487e-13	.535e-13	.655e-13
.414e+03	.801e-13	.772e-13	.508e-13	.593e-13	.669e-13
.431e+03	.849e-13	.788e-13	.529e-13	.632e-13	.683e-13
.448e+03	.878e-13	.804e-13	.549e-13	.658e-13	.697e-13
.466e+03	.899e-13	.821e-13	.569e-13	.676e-13	.711e-13
.485e+03	.915e-13	.837e-13	.589e-13	.692e-13	.725e-13
.504e+03	.929e-13	.853e-13	.608e-13	.706e-13	.739e-13
.524e+03	.937e-13	.866e-13	.625e-13	.715e-13	.751e-13
.545e+03	.940e-13	.880e-13	.641e-13	.721e-13	.763e-13
.567e+03	.931e-13	.894e-13	.657e-13	.718e-13	.775e-13
.589e+03	.892e-13	.908e-13	.672e-13	.691e-13	.787e-13
.613e+03	.860e-13	.920e-13	.687e-13	.669e-13	.798e-13
.638e+03	.870e-13	.933e-13	.700e-13	.679e-13	.808e-13
.663e+03	.880e-13	.945e-13	.713e-13	.690e-13	.819e-13
.690e+03	.888e-13	.955e-13	.725e-13	.699e-13	.828e-13
.717e+03	.896e-13	.965e-13	.736e-13	.708e-13	.837e-13
.746e+03	.904e-13	.976e-13	.747e-13	.717e-13	.846e-13
.776e+03	.912e-13	.986e-13	.758e-13	.726e-13	.855e-13
.807e+03	.919e-13	.995e-13	.768e-13	.735e-13	.863e-13
.839e+03	.926e-13	.100e-12	.776e-13	.742e-13	.870e-13
.873e+03	.931e-13	.101e-12	.784e-13	.750e-13	.876e-13
.907e+03	.937e-13	.102e-12	.792e-13	.757e-13	.883e-13
.944e+03	.943e-13	.103e-12	.799e-13	.764e-13	.889e-13
.982e+03	.948e-13	.103e-12	.806e-13	.771e-13	.896e-13
.102e+04	.952e-13	.104e-12	.811e-13	.777e-13	.901e-13
.106e+04	.956e-13	.104e-12	.816e-13	.782e-13	.905e-13
.110e+04	.959e-13	.105e-12	.820e-13	.788e-13	.909e-13
.115e+04	.962e-13	.105e-12	.824e-13	.793e-13	.913e-13
.119e+04	.964e-13	.106e-12	.827e-13	.797e-13	.917e-13
.124e+04	.966e-13	.106e-12	.830e-13	.801e-13	.919e-13
.129e+04	.967e-13	.106e-12	.831e-13	.804e-13	.921e-13
.134e+04	.967e-13	.106e-12	.833e-13	.806e-13	.922e-13
.140e+04	.967e-13	.106e-12	.834e-13	.808e-13	.923e-13
.145e+04	.966e-13	.107e-12	.834e-13	.810e-13	.924e-13
.151e+04	.966e-13	.107e-12	.834e-13	.812e-13	.924e-13
.157e+04	.963e-13	.106e-12	.832e-13	.812e-13	.922e-13
.163e+04	.960e-13	.106e-12	.830e-13	.811e-13	.920e-13
.170e+04	.958e-13	.106e-12	.827e-13	.811e-13	.918e-13

.3920e+05   .5300e+04   .2051e+04   .6880e+04

.6899e+04   .0000e+00

ERROR .GT. ERMAX

******************************** IFLAG= 904

K	ERROR(K)	1	2	3	4	5
1	.100e+01	.136e-09	.144e-09	.106e-09	.108e-09	.125e-09
2	.344e+07	.902e+07	.952e+07	.693e+07	.714e+07	.825e+07
3	.318e+00	.394e+04	.413e+04	.296e+04	.309e+04	.358e+04
4	.863e+01	.188e+05	.195e+05	.137e+05	.146e+05	.169e+05
5	.342e+01	.135e+05	.138e+05	.953e+04	.103e+05	.120e+05
6	.427e+01	.155e+05	.155e+05	.102e+05	.115e+05	.134e+05
7	.367e+01	.161e+05	.153e+05	.949e+04	.112e+05	.133e+05
8	.319e+01	.177e+05	.156e+05	.670e+04	.113e+05	.135e+05
9	.253e+01	.201e+05	.159e+05	.749e+04	.113e+05	.137e+05
10	.133e+01	.198e+05	.138e+05	.566e+04	.966e+04	.120e+05
11	.172e+01	.282e+05	.168e+05	.505e+04	.114e+05	.146e+05
12	.123e+01	.308e+05	.162e+05	.344e+04	.107e+05	.141e+05
13	.116e+01	.336e+05	.163e+05	.250e+04	.106e+05	.141e+05
14	.113e+01	.351e+05	.162e+05	.191e+04	.105e+05	.141e+05
15	.405e+00	.275e+05	.119e+05	.136e+04	.764e+04	.103e+05
16	.139e+01	.409e+05	.173e+05	.160e+04	.110e+05	.150e+05
17	.109e+01	.380e+05	.160e+05	.129e+04	.101e+05	.138e+05
18	.114e+01	.386e+05	.162e+05	.120e+04	.103e+05	.140e+05
19	.114e+01	.384e+05	.162e+05	.113e+04	.102e+05	.140e+05
20	.397e+00	.288e+05	.116e+05	.913e+03	.732e+04	.100e+05
21	.143e+01	.432e+05	.174e+05	.124e+04	.110e+05	.151e+05
22	.109e+01	.392e+05	.159e+05	.108e+04	.101e+05	.138e+05
23	.115e+01	.396e+05	.162e+05	.106e+04	.102e+05	.140e+05
24	.115e+01	.392e+05	.162e+05	.104e+04	.102e+05	.140e+05
25	.390e+00	.291e+05	.115e+05	.852e+03	.723e+04	.993e+04
26	.145e+01	.440e+05	.175e+05	.119e+04	.110e+05	.151e+05
27	.109e+01	.397e+05	.159e+05	.104e+04	.100e+05	.138e+05
28	.116e+01	.401e+05	.162e+05	.103e+04	.102e+05	.141e+05
29	.115e+01	.397e+05	.162e+05	.102e+04	.102e+05	.140e+05
30	.384e+00	.293e+05	.114e+05	.833e+03	.718e+04	.986e+04
31	.146e+01	.446e+05	.175e+05	.116e+04	.110e+05	.152e+05
32	.109e+01	.402e+05	.159e+05	.103e+04	.100e+05	.138e+05
33	.116e+01	.405e+05	.162e+05	.102e+04	.102e+05	.141e+05
34	.115e+01	.401e+05	.162e+05	.101e+04	.102e+05	.140e+05
35	.378e+00	.294e+05	.113e+05	.823e+03	.714e+04	.981e+04
36	.147e+01	.450e+05	.175e+05	.117e+04	.110e+05	.152e+05
37	.110e+01	.404e+05	.159e+05	.102e+04	.100e+05	.138e+05
38	.117e+01	.408e+05	.162e+05	.101e+04	.102e+05	.141e+05
39	.116e+01	.403e+05	.162e+05	.100e+04	.102e+05	.140e+05
40	.375e+00	.294e+05	.113e+05	.816e+03	.711e+04	.978e+04
41	.147e+01	.453e+05	.175e+05	.116e+04	.110e+05	.152e+05
42	.110e+01	.406e+05	.159e+05	.101e+04	.100e+05	.138e+05
43	.117e+01	.410e+05	.162e+05	.101e+04	.102e+05	.141e+05
44	.116e+01	.405e+05	.162e+05	.995e+03	.102e+05	.140e+05
45	.373e+00	.295e+05	.113e+05	.811e+03	.709e+04	.976e+04
46	.148e+01	.454e+05	.176e+05	.116e+04	.111e+05	.152e+05
47	.110e+01	.407e+05	.159e+05	.101e+04	.100e+05	.138e+05
48	.117e+01	.411e+05	.162e+05	.101e+04	.102e+05	.141e+05
49	.116e+01	.406e+05	.162e+05	.991e+03	.102e+05	.140e+05
50	.372e+00	.295e+05	.112e+05	.808e+03	.708e+04	.975e+04
51	.687e+00	.359e+05	.108e+05	.105e+04	.872e+04	.120e+05
52	.616e+00	.347e+05	.134e+05	.107e+04	.845e+04	.116e+05
53	.638e+00	.352e+05	.135e+05	.111e+04	.856e+04	.117e+05
54	.638e+00	.354e+05	.136e+05	.113e+04	.857e+04	.117e+05

55	.344e+00	.309e+05	.112e+05	.102e+04	.710e+04	.974e+04
56	.660e+00	.377e+05	.137e+05	.122e+04	.868e+04	.119e+05
57	.586e+00	.363e+05	.133e+05	.117e+04	.837e+04	.115e+05
58	.605e+00	.367e+05	.134e+05	.117e+04	.845e+04	.116e+05
59	.605e+00	.367e+05	.134e+05	.116e+04	.844e+04	.116e+05
60	.307e+00	.316e+05	.109e+05	.103e+04	.687e+04	.944e+04
61	.630e+00	.390e+05	.136e+05	.124e+04	.856e+04	.118e+05
62	.555e+00	.374e+05	.131e+05	.117e+04	.823e+04	.113e+05
63	.577e+00	.377e+05	.132e+05	.117e+04	.831e+04	.114e+05
64	.578e+00	.377e+05	.132e+05	.115e+04	.831e+04	.114e+05
65	.285e+00	.321e+05	.107e+05	.102e+04	.671e+04	.924e+04
66	.607e+00	.400e+05	.134e+05	.124e+04	.845e+04	.116e+05
67	.532e+00	.382e+05	.129e+05	.116e+04	.812e+04	.112e+05
68	.556e+00	.385e+05	.131e+05	.116e+04	.821e+04	.113e+05
69	.558e+00	.384e+05	.131e+05	.115e+04	.822e+04	.113e+05
70	.270e+00	.324e+05	.105e+05	.101e+04	.660e+04	.910e+04
71	.590e+00	.407e+05	.133e+05	.124e+04	.838e+04	.115e+05
72	.517e+00	.388e+05	.128e+05	.116e+04	.804e+04	.111e+05
73	.541e+00	.391e+05	.130e+05	.115e+04	.814e+04	.112e+05
74	.544e+00	.390e+05	.130e+05	.114e+04	.815e+04	.112e+05
75	.260e+00	.327e+05	.104e+05	.101e+04	.652e+04	.899e+04
76	.578e+00	.412e+05	.132e+05	.124e+04	.832e+04	.115e+05
77	.505e+00	.392e+05	.127e+05	.116e+04	.798e+04	.110e+05
78	.530e+00	.395e+05	.129e+05	.115e+04	.809e+04	.112e+05
79	.534e+00	.393e+05	.129e+05	.114e+04	.810e+04	.112e+05
80	.253e+00	.329e+05	.103e+05	.101e+04	.646e+04	.892e+04
81	.570e+00	.415e+05	.132e+05	.124e+04	.828e+04	.114e+05
82	.497e+00	.395e+05	.127e+05	.116e+04	.794e+04	.110e+05
83	.523e+00	.398e+05	.128e+05	.115e+04	.805e+04	.111e+05
84	.527e+00	.396e+05	.128e+05	.114e+04	.807e+04	.111e+05
85	.248e+00	.330e+05	.102e+05	.100e+04	.642e+04	.887e+04
86	.564e+00	.418e+05	.132e+05	.124e+04	.825e+04	.114e+05
87	.492e+00	.397e+05	.126e+05	.115e+04	.791e+04	.109e+05
88	.518e+00	.400e+05	.128e+05	.115e+04	.803e+04	.111e+05
89	.522e+00	.398e+05	.128e+05	.113e+04	.804e+04	.111e+05
90	.244e+00	.331e+05	.102e+05	.100e+04	.639e+04	.883e+04
91	.560e+00	.420e+05	.131e+05	.124e+04	.823e+04	.114e+05
92	.488e+00	.399e+05	.126e+05	.115e+04	.789e+04	.109e+05
93	.514e+00	.401e+05	.128e+05	.115e+04	.801e+04	.111e+05
94	.519e+00	.399e+05	.128e+05	.113e+04	.802e+04	.111e+05
95	.242e+00	.331e+05	.102e+05	.100e+04	.637e+04	.881e+04
96	.557e+00	.421e+05	.131e+05	.123e+04	.822e+04	.114e+05
97	.485e+00	.400e+05	.126e+05	.115e+04	.788e+04	.109e+05
98	.511e+00	.402e+05	.127e+05	.115e+04	.799e+04	.110e+05
99	.516e+00	.400e+05	.128e+05	.113e+04	.801e+04	.111e+05
100	.240e+00	.332e+05	.101e+05	.999e+03	.636e+04	.879e+04
101	.555e+00	.422e+05	.131e+05	.123e+04	.821e+04	.113e+05
102	.483e+00	.400e+05	.126e+05	.115e+04	.787e+04	.109e+05
103	.510e+00	.403e+05	.127e+05	.114e+04	.798e+04	.110e+05
104	.515e+00	.401e+05	.128e+05	.113e+04	.800e+04	.111e+05
105	.239e+00	.332e+05	.101e+05	.998e+03	.635e+04	.877e+04
106	.554e+00	.423e+05	.131e+05	.123e+04	.820e+04	.113e+05
107	.482e+00	.401e+05	.125e+05	.115e+04	.786e+04	.109e+05
108	.509e+00	.404e+05	.127e+05	.114e+04	.798e+04	.110e+05
109	.514e+00	.402e+05	.128e+05	.113e+04	.800e+04	.111e+05
110	.238e+00	.332e+05	.101e+05	.997e+03	.634e+04	.877e+04
111	.553e+00	.423e+05	.131e+05	.123e+04	.620e+04	.113e+05
112	.481e+00	.401e+05	.125e+05	.115e+04	.786e+04	.109e+05
113	.508e+00	.404e+05	.127e+05	.114e+04	.797e+04	.110e+05
114	.513e+00	.402e+05	.127e+05	.113e+04	.799e+04	.110e+05
115	.238e+00	.333e+05	.101e+05	.996e+03	.634e+04	.876e+04
116	.553e+00	.423e+05	.131e+05	.123e+04	.819e+04	.113e+05
117	.481e+00	.402e+05	.125e+05	.115e+04	.786e+04	.109e+05
118	.507e+00	.404e+05	.127e+05	.114e+04	.797e+04	.110e+05

119	.513e+00	.402e+05	.127e+05	.113e+04	.799e+04	.110e+05
120	.238e+00	.333e+05	.101e+05	.995e+03	.633e+04	.876e+04
121	.552e+00	.424e+05	.131e+05	.123e+04	.819e+04	.113e+05
122	.481e+00	.402e+05	.125e+05	.115e+04	.785e+04	.109e+05
123	.507e+00	.404e+05	.127e+05	.114e+04	.797e+04	.110e+05
124	.512e+00	.402e+05	.127e+05	.113e+04	.799e+04	.110e+05
125	.237e+00	.333e+05	.101e+05	.995e+03	.633e+04	.875e+04
126	.552e+00	.424e+05	.131e+05	.123e+04	.819e+04	.113e+05
127	.481e+00	.402e+05	.125e+05	.115e+04	.785e+04	.109e+05
128	.507e+00	.404e+05	.127e+05	.114e+04	.797e+04	.110e+05
129	.512e+00	.402e+05	.127e+05	.113e+04	.799e+04	.110e+05
130	.237e+00	.333e+05	.101e+05	.994e+03	.633e+04	.875e+04
131	.552e+00	.424e+05	.131e+05	.123e+04	.819e+04	.113e+05
132	.481e+00	.402e+05	.125e+05	.115e+04	.785e+04	.109e+05
133	.507e+00	.405e+05	.127e+05	.114e+04	.797e+04	.110e+05
134	.512e+00	.403e+05	.127e+05	.113e+04	.799e+04	.110e+05
135	.237e+00	.333e+05	.101e+05	.994e+03	.633e+04	.875e+04
136	.552e+00	.424e+05	.131e+05	.123e+04	.819e+04	.113e+05
137	.481e+00	.402e+05	.125e+05	.115e+04	.785e+04	.109e+05
138	.507e+00	.405e+05	.127e+05	.114e+04	.797e+04	.110e+05
139	.512e+00	.403e+05	.127e+05	.112e+04	.799e+04	.110e+05
140	.237e+00	.333e+05	.101e+05	.993e+03	.633e+04	.875e+04
141	.251e+00	.351e+05	.104e+05	.111e+04	.654e+04	.903e+04
142	.226e+00	.349e+05	.102e+05	.115e+04	.639e+04	.881e+04
143	.214e+00	.352e+05	.101e+05	.121e+04	.633e+04	.873e+04
144	.202e+00	.353e+05	.997e+04	.125e+04	.627e+04	.864e+04
145	.182e+00	.358e+05	.974e+04	.128e+04	.613e+04	.844e+04
146	.176e+00	.358e+05	.968e+04	.132e+04	.609e+04	.839e+04
147	.170e+00	.359e+05	.963e+04	.135e+04	.607e+04	.835e+04
148	.165e+00	.359e+05	.959e+04	.138e+04	.604e+04	.831e+04
149	.161e+00	.360e+05	.954e+04	.140e+04	.602e+04	.827e+04
150	.148e+00	.365e+05	.937e+04	.142e+04	.591e+04	.812e+04
151	.145e+00	.364e+05	.934e+04	.144e+04	.589e+04	.809e+04
152	.144e+00	.364e+05	.934e+04	.146e+04	.589e+04	.809e+04
153	.142e+00	.364e+05	.932e+04	.148e+04	.589e+04	.806e+04
154	.140e+00	.364e+05	.931e+04	.149e+04	.588e+04	.807e+04
155	.131e+00	.369e+05	.917e+04	.150e+04	.579e+04	.795e+04
156	.130e+00	.367e+05	.915e+04	.152e+04	.578e+04	.793e+04
157	.130e+00	.367e+05	.917e+04	.153e+04	.580e+04	.794e+04
158	.129e+00	.367e+05	.917e+04	.154e+04	.580e+04	.795e+04
159	.129e+00	.367e+05	.917e+04	.155e+04	.580e+04	.795e+04
160	.121e+00	.371e+05	.904e+04	.156e+04	.572e+04	.783e+04
161	.121e+00	.369e+05	.903e+04	.157e+04	.572e+04	.783e+04
162	.121e+00	.369e+05	.906e+04	.158e+04	.573e+04	.785e+04
163	.122e+00	.369e+05	.907e+04	.158e+04	.574e+04	.786e+04
164	.122e+00	.369e+05	.907e+04	.158e+04	.575e+04	.786e+04
165	.114e+00	.373e+05	.894e+04	.159e+04	.566e+04	.775e+04
166	.115e+00	.371e+05	.896e+04	.160e+04	.568e+04	.776e+04
167	.116e+00	.371e+05	.898e+04	.161e+04	.569e+04	.779e+04
168	.117e+00	.371e+05	.900e+04	.161e+04	.570e+04	.780e+04
169	.117e+00	.371e+05	.901e+04	.161e+04	.571e+04	.781e+04
170	.110e+00	.374e+05	.887e+04	.162e+04	.562e+04	.769e+04
171	.111e+00	.373e+05	.890e+04	.162e+04	.565e+04	.772e+04
172	.112e+00	.373e+05	.893e+04	.163e+04	.566e+04	.774e+04
173	.113e+00	.373e+05	.895e+04	.163e+04	.567e+04	.775e+04
174	.113e+00	.373e+05	.896e+04	.163e+04	.568e+04	.776e+04
175	.106e+00	.375e+05	.882e+04	.163e+04	.559e+04	.764e+04
176	.108e+00	.375e+05	.886e+04	.164e+04	.562e+04	.768e+04
177	.109e+00	.375e+05	.889e+04	.164e+04	.564e+04	.770e+04
178	.110e+00	.375e+05	.890e+04	.164e+04	.565e+04	.772e+04
179	.111e+00	.375e+05	.892e+04	.164e+04	.566e+04	.773e+04
180	.103e+00	.376e+05	.877e+04	.165e+04	.556e+04	.760e+04
181	.101e+00	.378e+05	.873e+04	.166e+04	.564e+04	.757e+04
182	.982e-01	.378e+05	.868e+04	.167e+04	.551e+04	.753e+04
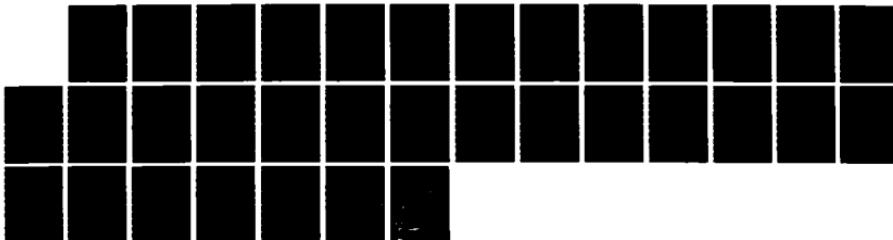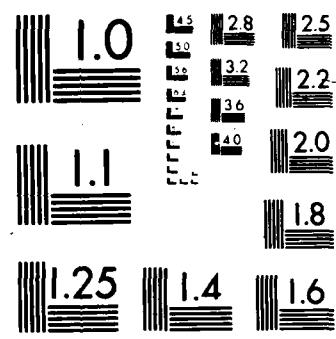
183	.960e-01	.379e+05	.865e+04	.169e+04	.549e+04	.749e+04
184	.940e-01	.380e+05	.861e+04	.170e+04	.547e+04	.746e+04
185	.902e-01	.381e+05	.854e+04	.172e+04	.543e+04	.740e+04
186	.891e-01	.381e+05	.853e+04	.173e+04	.542e+04	.739e+04
187	.879e-01	.381e+05	.851e+04	.175e+04	.542e+04	.737e+04
188	.869e-01	.382e+05	.849e+04	.176e+04	.541e+04	.736e+04
189	.860e-01	.382e+05	.848e+04	.177e+04	.540e+04	.735e+04
190	.833e-01	.383e+05	.842e+04	.178e+04	.537e+04	.730e+04
191	.828e-01	.383e+05	.842e+04	.180e+04	.537e+04	.729e+04
192	.823e-01	.383e+05	.841e+04	.181e+04	.537e+04	.729e+04
193	.818e-01	.383e+05	.841e+04	.182e+04	.536e+04	.728e+04
194	.814e-01	.383e+05	.840e+04	.182e+04	.536e+04	.728e+04
195	.792e-01	.384e+05	.835e+04	.183e+04	.533e+04	.724e+04
196	.790e-01	.384e+05	.835e+04	.184e+04	.533e+04	.724e+04
197	.789e-01	.384e+05	.835e+04	.185e+04	.534e+04	.724e+04
198	.787e-01	.384e+05	.835e+04	.186e+04	.534e+04	.724e+04
199	.785e-01	.384e+05	.835e+04	.186e+04	.534e+04	.723e+04
200	.766e-01	.385e+05	.830e+04	.187e+04	.531e+04	.719e+04
201	.766e-01	.384e+05	.631e+04	.188e+04	.531e+04	.720e+04
202	.767e-01	.384e+05	.831e+04	.188e+04	.532e+04	.720e+04
203	.767e-01	.384e+05	.831e+04	.189e+04	.532e+04	.720e+04
204	.766e-01	.384e+05	.831e+04	.189e+04	.532e+04	.720e+04
205	.749e-01	.385e+05	.827e+04	.189e+04	.529e+04	.717e+04
206	.751e-01	.385e+05	.827e+04	.190e+04	.530e+04	.717e+04
207	.752e-01	.385e+05	.828e+04	.190e+04	.530e+04	.718e+04
208	.753e-01	.385e+05	.828e+04	.191e+04	.531e+04	.718e+04
209	.754e-01	.385e+05	.829e+04	.191e+04	.531e+04	.718e+04
210	.738e-01	.386e+05	.824e+04	.191e+04	.528e+04	.714e+04
211	.740e-01	.385e+05	.825e+04	.192e+04	.529e+04	.715e+04
212	.742e-01	.385e+05	.826e+04	.192e+04	.529e+04	.716e+04
213	.744e-01	.385e+05	.826e+04	.192e+04	.530e+04	.716e+04
214	.745e-01	.385e+05	.827e+04	.192e+04	.530e+04	.716e+04
215	.729e-01	.386e+05	.822e+04	.192e+04	.527e+04	.713e+04
216	.733e-01	.386e+05	.824e+04	.193e+04	.528e+04	.714e+04
217	.735e-01	.385e+05	.824e+04	.193e+04	.529e+04	.714e+04
218	.737e-01	.385e+05	.825e+04	.193e+04	.529e+04	.715e+04
219	.738e-01	.385e+05	.825e+04	.193e+04	.529e+04	.715e+04
220	.723e-01	.386e+05	.821e+04	.193e+04	.527e+04	.712e+04
221	.727e-01	.386e+05	.822e+04	.194e+04	.528e+04	.713e+04
222	.730e-01	.386e+05	.823e+04	.194e+04	.528e+04	.713e+04
223	.732e-01	.386e+05	.824e+04	.194e+04	.528e+04	.714e+04
224	.733e-01	.385e+05	.824e+04	.194e+04	.529e+04	.714e+04
225	.718e-01	.386e+05	.820e+04	.194e+04	.526e+04	.711e+04
226	.723e-01	.386e+05	.822e+04	.194e+04	.527e+04	.712e+04
227	.725e-01	.386e+05	.822e+04	.195e+04	.526e+04	.713e+04
228	.728e-01	.386e+05	.823e+04	.194e+04	.528e+04	.713e+04
229	.729e-01	.386e+05	.823e+04	.194e+04	.528e+04	.714e+04
230	.714e-01	.386e+05	.819e+04	.195e+04	.526e+04	.710e+04
231	.720e-01	.386e+05	.821e+04	.195e+04	.527e+04	.711e+04
232	.722e-01	.386e+05	.821e+04	.195e+04	.527e+04	.712e+04
233	.725e-01	.386e+05	.822e+04	.195e+04	.528e+04	.713e+04
234	.726e-01	.386e+05	.823e+04	.195e+04	.528e+04	.713e+04
235	.711e-01	.386e+05	.818e+04	.195e+04	.525e+04	.709e+04
236	.717e-01	.307e+05	.820e+04	.195e+04	.527e+04	.711e+04
237	.720e-01	.386e+05	.821e+04	.195e+04	.527e+04	.711e+04
238	.722e-01	.386e+05	.822e+04	.195e+04	.527e+04	.712e+04
239	.724e-01	.386e+05	.822e+04	.195e+04	.528e+04	.712e+04
240	.708e-01	.386e+05	.818e+04	.195e+04	.525e+04	.709e+04
241	.715e-01	.387e+05	.820e+04	.196e+04	.526e+04	.710e+04
242	.717e-01	.386e+05	.820e+04	.196e+04	.527e+04	.711e+04
243	.720e-01	.386e+05	.821e+04	.196e+04	.527e+04	.712e+04
244	.722e-01	.386e+05	.822e+04	.195e+04	.528e+04	.712e+04
245	.706e-01	.387e+05	.817e+04	.196e+04	.525e+04	.708e+04
246	.713e-01	.387e+05	.819e+04	.196e+04	.526e+04	.710e+04

247	.715e-01	.387e+05	.820e+04	.196e+04	.527e+04	.711e+04
248	.718e-01	.387e+05	.821e+04	.196e+04	.527e+04	.711e+04
249	.720e-01	.387e+05	.821e+04	.196e+04	.527e+04	.712e+04
250	.704e-01	.387e+05	.816e+04	.196e+04	.524e+04	.708e+04
251	.712e-01	.387e+05	.819e+04	.196e+04	.526e+04	.710e+04
252	.714e-01	.387e+05	.819e+04	.196e+04	.526e+04	.710e+04
253	.716e-01	.387e+05	.820e+04	.196e+04	.527e+04	.711e+04
254	.718e-01	.387e+05	.821e+04	.196e+04	.527e+04	.711e+04
255	.702e-01	.387e+05	.816e+04	.196e+04	.524e+04	.707e+04
256	.710e-01	.387e+05	.819e+04	.197e+04	.526e+04	.710e+04
257	.712e-01	.387e+05	.819e+04	.196e+04	.526e+04	.710e+04
258	.715e-01	.387e+05	.820e+04	.196e+04	.527e+04	.710e+04
259	.716e-01	.387e+05	.820e+04	.196e+04	.527e+04	.711e+04
260	.700e-01	.387e+05	.815e+04	.196e+04	.524e+04	.707e+04
261	.709e-01	.387e+05	.818e+04	.197e+04	.526e+04	.709e+04
262	.711e-01	.387e+05	.819e+04	.197e+04	.526e+04	.710e+04
263	.713e-01	.387e+05	.819e+04	.196e+04	.527e+04	.710e+04
264	.715e-01	.387e+05	.820e+04	.196e+04	.527e+04	.711e+04
265	.698e-01	.387e+05	.815e+04	.196e+04	.524e+04	.706e+04
266	.708e-01	.388e+05	.818e+04	.197e+04	.526e+04	.709e+04
267	.709e-01	.387e+05	.818e+04	.197e+04	.526e+04	.709e+04
268	.712e-01	.387e+05	.819e+04	.197e+04	.526e+04	.710e+04
269	.714e-01	.387e+05	.820e+04	.196e+04	.527e+04	.710e+04
270	.697e-01	.387e+05	.815e+04	.196e+04	.523e+04	.706e+04
271	.707e-01	.388e+05	.818e+04	.197e+04	.526e+04	.709e+04
272	.708e-01	.387e+05	.818e+04	.197e+04	.526e+04	.709e+04
273	.711e-01	.387e+05	.819e+04	.197e+04	.526e+04	.710e+04
274	.712e-01	.387e+05	.819e+04	.197e+04	.526e+04	.710e+04
275	.695e-01	.387e+05	.814e+04	.197e+04	.523e+04	.706e+04
276	.706e-01	.388e+05	.818e+04	.197e+04	.526e+04	.709e+04
277	.707e-01	.388e+05	.818e+04	.197e+04	.526e+04	.709e+04
278	.710e-01	.388e+05	.819e+04	.197e+04	.526e+04	.709e+04
279	.711e-01	.387e+05	.819e+04	.197e+04	.526e+04	.710e+04
280	.694e-01	.387e+05	.814e+04	.197e+04	.523e+04	.705e+04
281	.705e-01	.388e+05	.817e+04	.197e+04	.525e+04	.708e+04
282	.706e-01	.388e+05	.818e+04	.197e+04	.526e+04	.709e+04
283	.708e-01	.388e+05	.818e+04	.197e+04	.526e+04	.709e+04
284	.710e-01	.388e+05	.819e+04	.197e+04	.526e+04	.710e+04
285	.692e-01	.387e+05	.813e+04	.197e+04	.523e+04	.705e+04
286	.704e-01	.388e+05	.817e+04	.198e+04	.525e+04	.708e+04
28	.705e-01	.388e+05	.817e+04	.197e+04	.525e+04	.708e+04
288	.707e-01	.388e+05	.818e+04	.197e+04	.526e+04	.709e+04
289	.709e-01	.388e+05	.818e+04	.197e+04	.526e+04	.709e+04
290	.691e-01	.387e+05	.813e+04	.197e+04	.523e+04	.705e+04
291	.703e-01	.388e+05	.817e+04	.198e+04	.525e+04	.708e+04
292	.704e-01	.388e+05	.817e+04	.197e+04	.525e+04	.708e+04
293	.706e-01	.388e+05	.818e+04	.197e+04	.526e+04	.709e+04
294	.708e-01	.388e+05	.818e+04	.197e+04	.526e+04	.709e+04
295	.690e-01	.387e+05	.813e+04	.197e+04	.523e+04	.704e+04
296	.703e-01	.389e+05	.817e+04	.198e+04	.525e+04	.708e+04
297	.703e-01	.388e+05	.817e+04	.198e+04	.525e+04	.708e+04
298	.705e-01	.388e+05	.818e+04	.197e+04	.526e+04	.709e+04
299	.707e-01	.388e+05	.818e+04	.197e+04	.526e+04	.709e+04
300	.688e-01	.387e+05	.812e+04	.197e+04	.522e+04	.704e+04
301	.702e-01	.389e+05	.817e+04	.198e+04	.525e+04	.708e+04
302	.702e-01	.388e+05	.817e+04	.198e+04	.525e+04	.708e+04
303	.704e-01	.388e+05	.817e+04	.196e+04	.525e+04	.708e+04
304	.706e-01	.388e+05	.818e+04	.197e+04	.526e+04	.709e+04
305	.687e-01	.387e+05	.812e+04	.197e+04	.522e+04	.704e+04
306	.701e-01	.389e+05	.816e+04	.198e+04	.525e+04	.708e+04
307	.701e-01	.388e+05	.816e+04	.198e+04	.525e+04	.708e+04
308	.704e-01	.388e+05	.817e+04	.198e+04	.525e+04	.708e+04
309	.705e-01	.388e+05	.818e+04	.198e+04	.526e+04	.708e+04
310	.686e-01	.387e+05	.812e+04	.197e+04	.522e+04	.703e+04

311	.700e-01	.389e+05	.816e+04	.198e+04	.525e+04	.707e+04
312	.700e-01	.388e+05	.816e+04	.198e+04	.525e+04	.707e+04
313	.703e-01	.388e+05	.817e+04	.198e+04	.525e+04	.708e+04
314	.704e-01	.388e+05	.817e+04	.198e+04	.525e+04	.708e+04
315	.685e-01	.387e+05	.811e+04	.198e+04	.522e+04	.703e+04
316	.700e-01	.389e+05	.816e+04	.198e+04	.525e+04	.707e+04
317	.700e-01	.389e+05	.816e+04	.198e+04	.525e+04	.707e+04
318	.702e-01	.389e+05	.817e+04	.198e+04	.525e+04	.708e+04
319	.703e-01	.389e+05	.817e+04	.198e+04	.525e+04	.708e+04
320	.684e-01	.387e+05	.811e+04	.198e+04	.522e+04	.703e+04
321	.699e-01	.389e+05	.816e+04	.199e+04	.525e+04	.707e+04
322	.699e-01	.389e+05	.816e+04	.198e+04	.525e+04	.707e+04
323	.701e-01	.389e+05	.816e+04	.198e+04	.525e+04	.708e+04
324	.703e-01	.389e+05	.817e+04	.198e+04	.525e+04	.708e+04
325	.683e-01	.387e+05	.811e+04	.198e+04	.522e+04	.703e+04
326	.698e-01	.389e+05	.816e+04	.199e+04	.525e+04	.707e+04
327	.698e-01	.389e+05	.816e+04	.198e+04	.525e+04	.707e+04
328	.700e-01	.389e+05	.816e+04	.198e+04	.525e+04	.707e+04
329	.702e-01	.389e+05	.817e+04	.198e+04	.525e+04	.708e+04
330	.682e-01	.388e+05	.811e+04	.198e+04	.521e+04	.702e+04
331	.698e-01	.389e+05	.816e+04	.199e+04	.525e+04	.707e+04
332	.697e-01	.389e+05	.815e+04	.198e+04	.525e+04	.707e+04
333	.700e-01	.389e+05	.816e+04	.198e+04	.525e+04	.707e+04
334	.701e-01	.389e+05	.816e+04	.198e+04	.525e+04	.708e+04
335	.681e-01	.388e+05	.810e+04	.198e+04	.521e+04	.702e+04
336	.697e-01	.390e+05	.816e+04	.199e+04	.525e+04	.707e+04
337	.697e-01	.389e+05	.815e+04	.199e+04	.524e+04	.707e+04
338	.699e-01	.389e+05	.816e+04	.198e+04	.525e+04	.707e+04
339	.700e-01	.389e+05	.816e+04	.198e+04	.525e+04	.707e+04
340	.680e-01	.388e+05	.810e+04	.198e+04	.521e+04	.702e+04
341	.697e-01	.390e+05	.815e+04	.199e+04	.525e+04	.707e+04
342	.696e-01	.389e+05	.815e+04	.199e+04	.524e+04	.706e+04
343	.698e-01	.389e+05	.816e+04	.198e+04	.525e+04	.707e+04
344	.700e-01	.389e+05	.816e+04	.198e+04	.525e+04	.707e+04
345	.679e-01	.388e+05	.810e+04	.198e+04	.521e+04	.702e+04
346	.696e-01	.390e+05	.815e+04	.199e+04	.525e+04	.707e+04
347	.695e-01	.389e+05	.815e+04	.199e+04	.524e+04	.706e+04
348	.698e-01	.389e+05	.816e+04	.199e+04	.525e+04	.707e+04
349	.699e-01	.389e+05	.816e+04	.198e+04	.525e+04	.707e+04
350	.678e-01	.388e+05	.810e+04	.198e+04	.521e+04	.702e+04
351	.695e-01	.390e+05	.815e+04	.199e+04	.525e+04	.706e+04
352	.695e-01	.389e+05	.815e+04	.199e+04	.524e+04	.706e+04
353	.697e-01	.389e+05	.815e+04	.199e+04	.525e+04	.707e+04
354	.698e-01	.389e+05	.816e+04	.198e+04	.525e+04	.707e+04
355	.677e-01	.388e+05	.809e+04	.196e+04	.521e+04	.701e+04
356	.695e-01	.390e+05	.815e+04	.199e+04	.524e+04	.706e+04
357	.694e-01	.389e+05	.815e+04	.199e+04	.524e+04	.706e+04
358	.696e-01	.389e+05	.815e+04	.199e+04	.525e+04	.707e+04
359	.698e-01	.389e+05	.816e+04	.199e+04	.525e+04	.707e+04
360	.677e-01	.388e+05	.809e+04	.198e+04	.521e+04	.701e+04
361	.694e-01	.390e+05	.815e+04	.199e+04	.524e+04	.706e+04
362	.693e-01	.389e+05	.814e+04	.199e+04	.524e+04	.706e+04
363	.696e-01	.389e+05	.815e+04	.199e+04	.524e+04	.706e+04
364	.697e-01	.389e+05	.815e+04	.199e+04	.525e+04	.707e+04
365	.676e-01	.388e+05	.809e+04	.198e+04	.521e+04	.701e+04
366	.694e-01	.390e+05	.815e+04	.200e+04	.524e+04	.706e+04
367	.693e-01	.390e+05	.814e+04	.199e+04	.524e+04	.706e+04
368	.695e-01	.390e+05	.815e+04	.199e+04	.524e+04	.706e+04
369	.696e-01	.390e+05	.815e+04	.199e+04	.525e+04	.707e+04
370	.675e-01	.388e+05	.809e+04	.198e+04	.520e+04	.701e+04
371	.694e-01	.390e+05	.815e+04	.200e+04	.524e+04	.706e+04
372	.692e-01	.390e+05	.814e+04	.199e+04	.524e+04	.706e+04
373	.694e-01	.390e+05	.815e+04	.199e+04	.524e+04	.706e+04
374	.696e-01	.390e+05	.815e+04	.199e+04	.524e+04	.706e+04

375	.674e-01	.388e+05	.808e+04	.199e+04	.520e+04	.701e+04
376	.693e-01	.390e+05	.815e+04	.200e+04	.524e+04	.706e+04
377	.692e-01	.390e+05	.814e+04	.199e+04	.524e+04	.705e+04
378	.694e-01	.390e+05	.815e+04	.199e+04	.524e+04	.706e+04
379	.695e-01	.390e+05	.815e+04	.199e+04	.524e+04	.706e+04
380	.673e-01	.388e+05	.808e+04	.199e+04	.520e+04	.700e+04
381	.693e-01	.390e+05	.814e+04	.200e+04	.524e+04	.706e+04
382	.691e-01	.390e+05	.814e+04	.199e+04	.524e+04	.705e+04
383	.693e-01	.390e+05	.815e+04	.199e+04	.524e+04	.706e+04
384	.695e-01	.390e+05	.815e+04	.199e+04	.524e+04	.706e+04
385	.673e-01	.388e+05	.808e+04	.199e+04	.520e+04	.700e+04
386	.692e-01	.390e+05	.814e+04	.200e+04	.524e+04	.706e+04
387	.691e-01	.390e+05	.814e+04	.199e+04	.524e+04	.705e+04
388	.693e-01	.390e+05	.814e+04	.199e+04	.524e+04	.706e+04
389	.694e-01	.390e+05	.815e+04	.199e+04	.524e+04	.706e+04
390	.672e-01	.388e+05	.808e+04	.199e+04	.520e+04	.700e+04
391	.692e-01	.391e+05	.814e+04	.200e+04	.524e+04	.706e+04
392	.690e-01	.390e+05	.814e+04	.200e+04	.524e+04	.705e+04
393	.692e-01	.390e+05	.814e+04	.199e+04	.524e+04	.706e+04
394	.694e-01	.390e+05	.815e+04	.199e+04	.524e+04	.706e+04
395	.671e-01	.388e+05	.808e+04	.199e+04	.520e+04	.700e+04
396	.691e-01	.391e+05	.814e+04	.200e+04	.524e+04	.706e+04
397	.690e-01	.390e+05	.813e+04	.200e+04	.524e+04	.705e+04
398	.692e-01	.390e+05	.814e+04	.199e+04	.524e+04	.706e+04
399	.693e-01	.390e+05	.814e+04	.199e+04	.524e+04	.706e+04
400	.671e-01	.388e+05	.807e+04	.199e+04	.520e+04	.700e+04
401	.691e-01	.391e+05	.814e+04	.200e+04	.524e+04	.706e+04
402	.689e-01	.390e+05	.813e+04	.200e+04	.524e+04	.705e+04
403	.691e-01	.390e+05	.814e+04	.199e-04	.524e+04	.705e+04
404	.693e-01	.390e+05	.814e+04	.199e+04	.524e+04	.706e+04
405	.670e-01	.388e+05	.807e+04	.199e+04	.520e+04	.700e+04
406	.691e-01	.391e+05	.814e+04	.200e-04	.524e+04	.705e+04
407	.689e-01	.390e+05	.813e+04	.200e+04	.524e+04	.705e+04
408	.691e-01	.390e+05	.814e+04	.200e+04	.524e+04	.705e+04
409	.692e-01	.390e+05	.814e+04	.199e+04	.524e+04	.706e+04
410	.670e-01	.388e+05	.807e+04	.199e+04	.520e+04	.699e+04
411	.690e-01	.391e+05	.814e+04	.200e+04	.524e+04	.705e+04
412	.688e-01	.390e+05	.813e+04	.200e+04	.524e+04	.705e+04
413	.691e-01	.390e+05	.814e+04	.200e-04	.524e+04	.705e+04
414	.692e-01	.390e+05	.814e+04	.199e-04	.524e+04	.706e+04
415	.669e-01	.388e+05	.807e+04	.199e+04	.520e+04	.699e+04
416	.690e-01	.391e+05	.814e+04	.200e+04	.524e+04	.705e+04
417	.688e-01	.390e+05	.813e+04	.200e+04	.523e+04	.705e+04
418	.690e-01	.390e+05	.814e+04	.200e+04	.524e+04	.705e+04
419	.691e-01	.390e+05	.814e+04	.199e+04	.524e+04	.705e+04
420	.668e-01	.388e+05	.807e+04	.199e+04	.520e+04	.699e+04
421	.690e-01	.391e+05	.814e+04	.200e+04	.524e+04	.705e+04
422	.687e-01	.390e+05	.812e+04	.200e+04	.523e+04	.705e+04
423	.690e-01	.390e+05	.814e+04	.200e+04	.524e+04	.705e+04
424	.691e-01	.390e+05	.814e+04	.200e-04	.524e+04	.705e+04
425	.668e-01	.388e+05	.807e+04	.199e+04	.519e+04	.699e+04
426	.689e-01	.391e+05	.814e+04	.201e+04	.524e+04	.705e+04
427	.687e-01	.390e+05	.813e+04	.200e+04	.523e-04	.704e+04
428	.689e-01	.390e+05	.813e+04	.200e+04	.524e+04	.705e+04
429	.690e-01	.390e+05	.814e+04	.200e+04	.524e+04	.705e+04
430	.667e-01	.388e+05	.806e+04	.199e+04	.519e+04	.699e+04
431	.689e-01	.391e+05	.814e+04	.201e+04	.524e+04	.705e+04
432	.687e-01	.390e+05	.812e+04	.200e+04	.523e+04	.704e+04
433	.689e-01	.391e+05	.813e+04	.200e+04	.524e+04	.705e+04
434	.690e-01	.390e+05	.814e+04	.200e+04	.524e+04	.705e+04
435	.667e-01	.388e+05	.806e+04	.199e+04	.519e+04	.699e+04
436	.689e-01	.391e+05	.814e+04	.201e+04	.524e+04	.705e+04
437	.686e-01	.391e+05	.813e+04	.200e+04	.523e+04	.704e+04
438	.689e-01	.391e+05	.813e+04	.200e+04	.524e+04	.705e+04

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439    .690e-01    .391e+05    .814e+04    .200e+04    .524e+04    .705e+04
440    .666e-01    .388e+05    .806e+04    .199e+04    .519e+04    .699e+04
441    .689e-01    .391e+05    .813e+04    .201e+04    .524e+04    .705e+04
442    .686e-01    .391e+05    .813e+04    .200e+04    .523e+04    .704e+04
443    .688e-01    .391e+05    .813e+04    .200e+04    .524e+04    .705e+04
444    .689e-01    .391e+05    .813e+04    .200e+04    .524e+04    .705e+04
445    .666e-01    .388e+05    .806e+04    .199e+04    .519e+04    .699e+04
446    .688e-01    .391e+05    .813e+04    .201e+04    .524e+04    .705e+04
447    .686e-01    .391e+05    .812e+04    .200e+04    .523e+04    .704e+04
448    .688e-01    .391e+05    .813e+04    .200e+04    .524e+04    .705e+04
449    .689e-01    .391e+05    .813e+04    .200e+04    .524e+04    .705e+04
450    .665e-01    .388e+05    .806e+04    .199e+04    .519e+04    .698e+04
451    .688e-01    .391e+05    .813e+04    .201e+04    .524e+04    .705e+04
452    .685e-01    .391e+05    .812e+04    .200e+04    .523e+04    .704e+04
453    .688e-01    .391e+05    .813e+04    .200e+04    .524e+04    .705e+04
454    .689e-01    .391e+05    .813e+04    .200e+04    .524e+04    .705e+04
455    .665e-01    .388e+05    .806e+04    .199e+04    .519e+04    .698e+04
456    .688e-01    .391e+05    .813e+04    .201e+04    .524e+04    .705e+04
457    .685e-01    .391e+05    .812e+04    .200e+04    .523e+04    .704e+04
458    .687e-01    .391e+05    .813e+04    .200e+04    .523e+04    .705e+04
459    .688e-01    .391e+05    .813e+04    .200e+04    .524e+04    .705e+04
460    .664e-01    .388e+05    .806e+04    .200e+04    .519e+04    .698e+04
461    .687e-01    .392e+05    .813e+04    .201e+04    .524e+04    .705e+04
462    .685e-01    .391e+05    .812e+04    .200e+04    .523e+04    .704e+04
463    .687e-01    .391e+05    .813e+04    .200e+04    .523e+04    .704e+04
464    .688e-01    .391e+05    .813e+04    .200e+04    .524e+04    .705e+04
465    .664e-01    .388e+05    .805e+04    .200e+04    .519e+04    .698e+04
466    .687e-01    .392e+05    .813e+04    .201e+04    .524e+04    .705e+04
467    .684e-01    .391e+05    .812e+04    .200e+04    .523e+04    .704e+04
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469    .688e-01    .391e+05    .813e+04    .200e+04    .524e+04    .705e+04
470    .663e-01    .388e+05    .805e+04    .200e+04    .519e+04    .698e+04
471    .687e-01    .392e+05    .813e+04    .201e+04    .524e+04    .705e+04
472    .684e-01    .391e+05    .812e+04    .201e+04    .523e+04    .704e+04
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474    .687e-01    .391e+05    .813e+04    .200e+04    .523e+04    .705e+04
475    .663e-01    .388e+05    .805e+04    .200e+04    .519e+04    .698e+04
476    .687e-01    .392e+05    .813e+04    .201e+04    .524e+04    .705e+04
477    .684e-01    .391e+05    .812e+04    .201e+04    .523e+04    .704e+04
478    .686e-01    .391e+05    .813e+04    .200e+04    .523e+04    .704e+04
479    .687e-01    .391e+05    .813e+04    .200e+04    .523e+04    .704e+04
480    .663e-01    .388e+05    .805e+04    .200e+04    .519e+04    .698e+04
481    .686e-01    .392e+05    .813e+04    .201e+04    .524e+04    .705e+04
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483    .686e-01    .391e+05    .813e+04    .200e+04    .523e+04    .704e+04
484    .687e-01    .391e+05    .813e+04    .200e+04    .523e+04    .704e+04
485    .662e-01    .388e+05    .805e+04    .200e+04    .519e+04    .698e+04
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487    .683e-01    .391e+05    .812e+04    .201e+04    .523e+04    .704e+04
488    .685e-01    .391e+05    .812e+04    .201e+04    .523e+04    .704e+04
489    .686e-01    .391e+05    .813e+04    .200e+04    .523e+04    .704e+04
490    .662e-01    .388e+05    .805e+04    .200e+04    .519e+04    .698e+04
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492    .683e-01    .391e+05    .812e+04    .201e+04    .523e+04    .703e+04
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500    .661e-01    .388e+05    .805e+04    .200e+04    .518e+04    .697e+04
501    .686e-01    .392e+05    .813e+04    .202e+04    .524e+04    .704e+04
502    .682e-01    .391e+05    .812e+04    .201e+04    .523e+04    .703e+04
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503	.685e-01	.391e+05	.812e+04	.201e+04	.523e+04	.704e+04
504	.686e-01	.391e+05	.812e+04	.200e+04	.523e+04	.704e+04
505	.661e-01	.388e+05	.805e+04	.200e+04	.518e+04	.697e+04
506	.685e-01	.392e+05	.813e+04	.202e+04	.524e+04	.704e+04
507	.682e-01	.391e+05	.812e+04	.201e+04	.523e+04	.703e+04
508	.684e-01	.391e+05	.812e+04	.201e+04	.523e+04	.704e+04
509	.685e-01	.391e+05	.812e+04	.200e+04	.523e+04	.704e+04
510	.660e-01	.388e+05	.804e+04	.200e+04	.518e+04	.697e+04
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512	.682e-01	.391e+05	.811e+04	.201e+04	.523e+04	.703e+04
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515	.660e-01	.388e+05	.804e+04	.200e+04	.518e+04	.697e+04
516	.685e-01	.392e+05	.813e+04	.202e+04	.524e+04	.704e+04
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521	.685e-01	.392e+05	.813e+04	.202e+04	.524e+04	.704e+04
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523	.684e-01	.391e+05	.812e+04	.201e+04	.523e+04	.704e+04
524	.685e-01	.391e+05	.812e+04	.201e+04	.523e+04	.704e+04
525	.659e-01	.388e+05	.804e+04	.200e+04	.518e+04	.697e+04
526	.685e-01	.392e+05	.812e+04	.202e+04	.524e+04	.704e+04
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528	.683e-01	.391e+05	.812e+04	.201e+04	.523e+04	.704e+04
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530	.659e-01	.388e+05	.804e+04	.200e+04	.518e+04	.697e+04
531	.685e-01	.392e+05	.812e+04	.202e+04	.523e+04	.704e+04
532	.681e-01	.391e+05	.811e+04	.201e+04	.523e+04	.703e+04
533	.683e-01	.391e+05	.812e+04	.201e+04	.523e+04	.704e+04
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546	.684e-01	.392e+05	.812e+04	.202e+04	.523e+04	.704e+04
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550	.658e-01	.388e+05	.804e+04	.200e+04	.518e+04	.697e+04
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556	.684e-01	.392e+05	.812e+04	.202e+04	.523e+04	.704e+04
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563	.682e-01	.392e+05	.812e+04	.201e+04	.523e+04	.703e+04
564	.683e-01	.392e+05	.812e+04	.201e+04	.523e+04	.704e+04
565	.657e-01	.388e+05	.804e+04	.200e+04	.518e+04	.696e+04
566	.683e-01	.392e+05	.812e+04	.202e+04	.523e+04	.704e+04

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570	.657e-01	.388e+05	.803e+04	.200e+04	.518e+04	.696e+04
571	.683e-01	.392e+05	.812e+04	.202e+04	.523e+04	.704e+04
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573	.682e-01	.392e+05	.811e+04	.201e+04	.523e+04	.703e+04
574	.683e-01	.392e+05	.812e+04	.201e+04	.523e+04	.703e+04
575	.657e-01	.388e+05	.803e+04	.200e+04	.518e+04	.696e+04
576	.683e-01	.392e+05	.812e+04	.202e+04	.523e+04	.704e+04
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578	.681e-01	.392e+05	.811e+04	.201e+04	.523e+04	.703e+04
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581	.683e-01	.392e+05	.812e+04	.202e+04	.523e+04	.704e+04
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583	.681e-01	.392e+05	.811e+04	.201e+04	.523e+04	.703e+04
584	.682e-01	.392e+05	.812e+04	.201e+04	.523e+04	.703e+04
585	.656e-01	.388e+05	.803e+04	.201e+04	.518e+04	.696e+04
586	.683e-01	.393e+05	.812e+04	.202e+04	.523e+04	.704e+04
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588	.681e-01	.392e+05	.811e+04	.201e+04	.523e+04	.703e+04
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590	.656e-01	.388e+05	.803e+04	.201e+04	.518e+04	.696e+04
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596	.683e-01	.393e+05	.812e+04	.202e+04	.523e+04	.704e+04
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602	.678e-01	.392e+05	.811e+04	.202e+04	.522e+04	.702e+04
603	.681e-01	.392e+05	.811e+04	.201e+04	.523e+04	.703e+04
604	.682e-01	.392e+05	.811e+04	.201e+04	.523e+04	.703e+04
605	.655e-01	.389e+05	.803e+04	.201e+04	.518e+04	.696e+04
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607	.678e-01	.392e+05	.811e+04	.202e+04	.522e+04	.702e+04
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609	.681e-01	.392e+05	.811e+04	.201e+04	.523e+04	.703e+04
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614	.681e-01	.392e+05	.811e+04	.201e+04	.523e+04	.703e+04
615	.655e-01	.389e+05	.803e+04	.201e+04	.518e+04	.696e+04
616	.682e-01	.393e+05	.812e+04	.203e+04	.523e+04	.704e+04
617	.678e-01	.392e+05	.810e+04	.202e+04	.522e+04	.702e+04
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619	.681e-01	.392e+05	.811e+04	.201e+04	.523e+04	.703e+04
620	.655e-01	.389e+05	.803e+04	.201e+04	.518e+04	.696e+04
621	.682e-01	.393e+05	.812e+04	.202e+04	.523e+04	.704e+04
622	.678e-01	.392e+05	.810e+04	.202e+04	.522e+04	.702e+04
623	.680e-01	.392e+05	.811e+04	.202e+04	.523e+04	.703e+04
624	.681e-01	.392e+05	.811e+04	.201e+04	.523e+04	.703e+04
625	.655e-01	.389e+05	.803e+04	.201e+04	.517e+04	.696e+04
626	.682e-01	.393e+05	.812e+04	.203e+04	.523e+04	.704e+04
627	.678e-01	.392e+05	.810e+04	.202e+04	.522e+04	.702e+04
628	.680e-01	.392e+05	.811e+04	.202e+04	.523e+04	.703e+04
629	.681e-01	.392e+05	.811e+04	.201e+04	.523e+04	.703e+04
630	.655e-01	.389e+05	.803e+04	.201e+04	.517e+04	.696e+04

631	.682e-01	.393e+05	.812e+04	.203e+04	.523e+04	.704e+04
632	.678e-01	.392e+05	.810e+04	.202e+04	.522e+04	.702e+04
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636	.682e-01	.393e+05	.812e+04	.203e+04	.523e+04	.704e+04
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650	.654e-01	.389e+05	.803e+04	.201e+04	.517e+04	.696e+04
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663	.679e-01	.392e+05	.811e+04	.202e+04	.522e+04	.703e+04
664	.680e-01	.392e+05	.811e+04	.202e+04	.523e+04	.703e+04
665	.653e-01	.389e+05	.802e+04	.201e+04	.517e+04	.695e+04
666	.681e-01	.393e+05	.812e+04	.203e+04	.523e+04	.703e+04
667	.677e-01	.392e+05	.810e+04	.202e+04	.522e+04	.702e+04
668	.679e-01	.392e+05	.811e+04	.202e+04	.522e+04	.703e+04
669	.680e-01	.392e+05	.811e+04	.202e+04	.522e+04	.703e+04
670	.653e-01	.389e+05	.802e+04	.201e+04	.517e+04	.695e+04
671	.681e-01	.393e+05	.812e+04	.203e+04	.523e+04	.703e+04
672	.677e-01	.392e+05	.810e+04	.202e+04	.522e+04	.702e+04
673	.679e-01	.392e+05	.811e+04	.202e+04	.522e+04	.703e+04
674	.680e-01	.392e+05	.811e+04	.202e+04	.522e+04	.703e+04
675	.653e-01	.389e+05	.802e+04	.201e+04	.517e+04	.695e+04
676	.681e-01	.393e+05	.812e+04	.203e+04	.523e+04	.703e+04
677	.677e-01	.392e+05	.810e+04	.202e+04	.522e+04	.702e+04
678	.679e-01	.392e+05	.811e+04	.202e+04	.522e+04	.703e+04
679	.680e-01	.392e+05	.811e+04	.202e+04	.522e+04	.703e+04
680	.653e-01	.389e+05	.802e+04	.201e+04	.517e+04	.695e+04
681	.681e-01	.393e+05	.812e+04	.203e+04	.523e+04	.703e+04
682	.677e-01	.392e+05	.810e+04	.202e+04	.522e+04	.702e+04
683	.679e-01	.392e+05	.811e+04	.202e+04	.522e+04	.703e+04
684	.680e-01	.392e+05	.811e+04	.202e+04	.522e+04	.703e+04
685	.653e-01	.389e+05	.802e+04	.201e+04	.517e+04	.695e+04
686	.681e-01	.393e+05	.812e+04	.203e+04	.523e+04	.703e+04
687	.676e-01	.392e+05	.810e+04	.202e+04	.522e+04	.702e+04
688	.679e-01	.392e+05	.811e+04	.202e+04	.522e+04	.703e+04
689	.679e-01	.392e+05	.811e+04	.202e+04	.522e+04	.703e+04
690	.653e-01	.389e+05	.802e+04	.201e+04	.517e+04	.695e+04
691	.681e-01	.393e+05	.811e+04	.203e+04	.523e+04	.703e+04
692	.676e-01	.392e+05	.810e+04	.202e+04	.522e+04	.702e+04
693	.679e-01	.392e+05	.811e+04	.202e+04	.522e+04	.703e+04
694	.679e-01	.392e+05	.811e+04	.202e+04	.522e+04	.703e+04

695	.653e-01	.389e+05	.802e+04	.201e+04	.517e+04	.695e+04
696	.681e-01	.393e+05	.811e+04	.203e+04	.523e+04	.703e+04
697	.676e-01	.392e+05	.810e+04	.202e+04	.522e+04	.702e+04
698	.678e-01	.392e+05	.811e+04	.202e+04	.522e+04	.703e+04
699	.679e-01	.392e+05	.811e+04	.202e+04	.522e+04	.703e+04
700	.652e-01	.389e+05	.802e+04	.201e+04	.517e+04	.695e+04
701	.681e-01	.393e+05	.811e+04	.203e+04	.523e+04	.703e+04
702	.676e-01	.392e+05	.810e+04	.202e+04	.522e+04	.702e+04
703	.678e-01	.392e+05	.811e+04	.202e+04	.522e+04	.702e+04
704	.679e-01	.392e+05	.811e+04	.202e+04	.522e+04	.703e+04
705	.652e-01	.389e+05	.802e+04	.201e+04	.517e+04	.695e+04
706	.681e-01	.393e+05	.811e+04	.203e+04	.523e+04	.703e+04
707	.676e-01	.392e+05	.810e+04	.202e+04	.522e+04	.702e+04
708	.678e-01	.392e+05	.811e+04	.202e+04	.522e+04	.702e+04
709	.679e-01	.392e+05	.811e+04	.202e+04	.522e+04	.703e+04
710	.652e-01	.389e+05	.802e+04	.201e+04	.517e+04	.695e+04
711	.681e-01	.393e+05	.811e+04	.203e+04	.523e+04	.703e+04
712	.676e-01	.392e+05	.810e+04	.202e+04	.522e+04	.702e+04
713	.678e-01	.392e+05	.811e+04	.202e+04	.522e+04	.702e+04
714	.679e-01	.392e+05	.811e+04	.202e+04	.522e+04	.703e+04
715	.652e-01	.389e+05	.802e+04	.201e+04	.517e+04	.695e+04
716	.681e-01	.393e+05	.811e+04	.203e+04	.522e+04	.703e+04
717	.676e-01	.392e-05	.810e+04	.202e+04	.522e+04	.702e+04
718	.678e-01	.392e-05	.810e+04	.202e+04	.522e+04	.702e+04
719	.679e-01	.392e+05	.811e+04	.202e+04	.522e+04	.703e+04
720	.652e-01	.389e+05	.802e+04	.201e+04	.517e+04	.695e+04
721	.681e-01	.393e+05	.811e+04	.203e+04	.523e+04	.703e+04
722	.676e-01	.392e+05	.810e+04	.202e+04	.522e+04	.702e+04
723	.678e-01	.392e+05	.810e+04	.202e+04	.522e+04	.702e+04
724	.679e-01	.392e+05	.811e+04	.202e+04	.522e+04	.703e+04
725	.652e-01	.389e+05	.802e+04	.201e+04	.517e+04	.695e+04
726	.680e-01	.393e+05	.811e+04	.203e+04	.523e+04	.703e+04
727	.676e-01	.392e+05	.810e+04	.202e+04	.522e+04	.702e+04
728	.678e-01	.392e+05	.810e+04	.202e+04	.522e+04	.702e+04
729	.679e-01	.392e+05	.811e+04	.202e+04	.522e+04	.703e+04
730	.652e-01	.389e+05	.802e+04	.201e+04	.517e+04	.695e+04
731	.680e-01	.393e+05	.811e+04	.203e+04	.523e+04	.703e+04
732	.676e-01	.392e+05	.810e+04	.202e+04	.522e+04	.702e+04
733	.678e-01	.392e+05	.810e+04	.202e+04	.522e+04	.702e+04
734	.679e-01	.392e+05	.811e+04	.202e+04	.522e+04	.703e+04
735	.652e-01	.389e+05	.802e+04	.201e+04	.517e+04	.695e+04
736	.680e-01	.393e+05	.811e+04	.203e+04	.523e+04	.703e+04
737	.676e-01	.392e+05	.810e+04	.202e+04	.522e+04	.702e+04
738	.678e-01	.392e+05	.810e+04	.202e+04	.522e+04	.702e+04
739	.679e-01	.392e+05	.811e+04	.202e+04	.522e+04	.703e+04
740	.652e-01	.389e+05	.802e+04	.201e+04	.517e+04	.695e+04
741	.680e-01	.393e+05	.811e+04	.203e+04	.523e+04	.703e+04
742	.676e-01	.392e+05	.810e+04	.202e+04	.522e+04	.702e+04
743	.678e-01	.392e+05	.810e+04	.202e+04	.522e+04	.702e+04
744	.678e-01	.392e+05	.811e+04	.202e+04	.522e+04	.702e+04
745	.651e-01	.389e+05	.802e+04	.201e+04	.517e+04	.695e+04
746	.680e-01	.393e+05	.811e+04	.203e+04	.523e+04	.703e+04
747	.675e-01	.392e+05	.810e+04	.202e+04	.522e+04	.702e+04
748	.678e-01	.392e+05	.810e+04	.202e+04	.522e+04	.702e+04
749	.678e-01	.392e+05	.811e+04	.202e+04	.522e+04	.702e+04
750	.651e-01	.389e+05	.802e+04	.201e+04	.517e+04	.695e+04
751	.680e-01	.393e+05	.811e+04	.203e+04	.523e+04	.703e+04
752	.675e-01	.392e+05	.810e+04	.202e+04	.522e+04	.702e+04
753	.678e-01	.392e+05	.810e+04	.202e+04	.522e+04	.702e+04
754	.678e-01	.392e+05	.811e+04	.202e+04	.522e+04	.702e+04
755	.651e-01	.389e+05	.802e+04	.201e+04	.517e+04	.695e+04
756	.680e-01	.393e+05	.811e+04	.203e+04	.523e+04	.703e+04
757	.675e-01	.392e+05	.810e+04	.203e+04	.522e+04	.702e+04
758	.678e-01	.392e+05	.810e+04	.202e+04	.522e+04	.702e+04

759	.678e-01	.392e+05	.810e+04	.202e+04	.522e+04	.702e+04
760	.651e-01	.389e+05	.802e+04	.201e+04	.517e+04	.695e+04
761	.680e-01	.393e+05	.811e+04	.203e+04	.523e+04	.703e+04
762	.675e-01	.392e+05	.810e+04	.203e+04	.522e+04	.702e+04
763	.677e-01	.392e+05	.810e+04	.202e+04	.522e+04	.702e+04
764	.678e-01	.392e+05	.810e+04	.202e+04	.522e+04	.702e+04
765	.651e-01	.389e+05	.802e+04	.201e+04	.517e+04	.695e+04
766	.680e-01	.393e+05	.811e+04	.203e+04	.523e+04	.703e+04
767	.675e-01	.392e+05	.810e+04	.203e+04	.522e+04	.702e+04
768	.677e-01	.392e+05	.810e+04	.202e+04	.522e+04	.702e+04
769	.678e-01	.392e+05	.810e+04	.202e+04	.522e+04	.702e+04
770	.651e-01	.389e+05	.802e+04	.201e+04	.517e+04	.695e+04
771	.680e-01	.393e+05	.811e+04	.203e+04	.523e+04	.703e+04
772	.675e-01	.392e+05	.810e+04	.203e+04	.522e+04	.702e+04
773	.677e-01	.392e+05	.810e+04	.202e+04	.522e+04	.702e+04
774	.678e-01	.392e+05	.810e+04	.202e+04	.522e+04	.702e+04
775	.651e-01	.389e+05	.802e+04	.201e+04	.517e+04	.695e+04
776	.680e-01	.393e+05	.811e+04	.203e+04	.523e+04	.703e+04
777	.675e-01	.392e+05	.810e+04	.203e+04	.522e+04	.702e+04
778	.677e-01	.392e+05	.810e+04	.202e+04	.522e+04	.702e+04
779	.678e-01	.392e+05	.810e+04	.202e+04	.522e+04	.702e+04
780	.651e-01	.389e+05	.802e+04	.201e+04	.517e+04	.695e+04
781	.680e-01	.393e+05	.811e+04	.203e+04	.523e+04	.703e+04
782	.675e-01	.392e+05	.810e+04	.203e+04	.522e+04	.702e+04
783	.677e-01	.392e+05	.810e+04	.202e+04	.522e+04	.702e+04
784	.678e-01	.392e+05	.810e+04	.202e+04	.522e+04	.702e+04
785	.651e-01	.389e+05	.802e+04	.201e+04	.517e+04	.695e+04
786	.680e-01	.393e+05	.811e+04	.203e+04	.523e+04	.703e+04
787	.675e-01	.392e+05	.810e+04	.203e+04	.522e+04	.702e+04
788	.677e-01	.392e+05	.810e+04	.202e+04	.522e+04	.702e+04
789	.678e-01	.392e+05	.810e+04	.202e+04	.522e+04	.702e+04
790	.651e-01	.389e+05	.802e+04	.201e+04	.517e+04	.695e+04
791	.680e-01	.393e+05	.811e+04	.203e+04	.523e+04	.703e+04
792	.675e-01	.392e+05	.810e+04	.203e+04	.522e+04	.702e+04
793	.677e-01	.392e+05	.810e+04	.202e+04	.522e+04	.702e+04
794	.678e-01	.392e+05	.810e+04	.202e+04	.522e+04	.702e+04
795	.651e-01	.389e+05	.802e+04	.201e+04	.517e+04	.695e+04
796	.680e-01	.393e+05	.811e+04	.203e+04	.523e+04	.703e+04
797	.675e-01	.392e+05	.810e+04	.203e+04	.522e+04	.702e+04
798	.677e-01	.392e+05	.810e+04	.202e+04	.522e+04	.702e+04
799	.678e-01	.392e+05	.810e+04	.202e+04	.522e+04	.702e+04
800	.651e-01	.389e+05	.802e+04	.202e+04	.517e+04	.695e+04
801	.680e-01	.393e+05	.811e+04	.203e+04	.523e+04	.703e+04
802	.675e-01	.392e+05	.810e+04	.203e+04	.522e+04	.702e+04
803	.677e-01	.392e+05	.810e+04	.203e+04	.522e+04	.702e+04
804	.678e-01	.392e+05	.810e+04	.202e+04	.522e+04	.702e+04
805	.650e-01	.389e+05	.801e+04	.202e+04	.517e+04	.695e+04
806	.680e-01	.393e+05	.811e+04	.204e+04	.523e+04	.703e+04
807	.675e-01	.392e+05	.810e+04	.203e+04	.522e+04	.702e+04
808	.677e-01	.392e+05	.810e+04	.203e+04	.522e+04	.702e+04
809	.678e-01	.392e+05	.810e+04	.202e+04	.522e+04	.702e+04
810	.650e-01	.389e+05	.801e+04	.202e+04	.517e+04	.695e+04
811	.680e-01	.393e+05	.811e+04	.204e+04	.523e+04	.703e+04
812	.675e-01	.392e+05	.809e+04	.203e+04	.522e+04	.702e+04
813	.677e-01	.392e+05	.810e+04	.203e+04	.522e+04	.702e+04
814	.678e-01	.392e+05	.810e+04	.202e+04	.522e+04	.702e+04
815	.650e-01	.389e+05	.801e+04	.202e+04	.517e+04	.695e+04
816	.680e-01	.393e+05	.811e+04	.204e+04	.523e+04	.703e+04
817	.675e-01	.392e+05	.809e+04	.203e+04	.522e+04	.702e+04
818	.677e-01	.392e+05	.810e+04	.203e+04	.522e+04	.702e+04
819	.677e-01	.392e+05	.810e+04	.202e+04	.522e+04	.702e+04
820	.650e-01	.389e+05	.801e+04	.202e+04	.517e+04	.695e+04
821	.680e-01	.393e+05	.811e+04	.204e+04	.523e+04	.703e+04
822	.675e-01	.392e+05	.809e+04	.203e+04	.522e+04	.702e+04

823	.677e-01	.392e+05	.810e+04	.203e+04	.522e+04	.702e+04
824	.677e-01	.392e+05	.810e+04	.202e+04	.522e+04	.702e+04
825	.650e-01	.389e+05	.801e+04	.202e+04	.517e+04	.695e+04
826	.679e-01	.393e+05	.811e+04	.204e+04	.523e+04	.703e+04
827	.675e-01	.392e+05	.809e+04	.203e+04	.522e+04	.702e+04
828	.677e-01	.392e+05	.810e+04	.203e+04	.522e+04	.702e+04
829	.677e-01	.392e+05	.810e+04	.202e+04	.522e+04	.702e+04
830	.650e-01	.389e+05	.801e+04	.202e+04	.517e+04	.695e+04
831	.679e-01	.393e+05	.811e+04	.204e+04	.523e+04	.703e+04
832	.674e-01	.392e+05	.809e+04	.203e+04	.522e+04	.702e+04
833	.677e-01	.392e+05	.810e+04	.203e+04	.522e+04	.702e+04
834	.677e-01	.392e+05	.810e+04	.202e+04	.522e+04	.702e+04
835	.650e-01	.389e+05	.801e+04	.202e+04	.517e+04	.695e+04
836	.679e-01	.393e+05	.811e+04	.204e+04	.523e+04	.703e+04
837	.674e-01	.392e+05	.809e+04	.203e+04	.522e+04	.701e+04
838	.677e-01	.392e+05	.810e+04	.203e+04	.522e+04	.702e+04
839	.677e-01	.392e+05	.810e+04	.202e+04	.522e+04	.702e+04
840	.650e-01	.389e+05	.801e+04	.202e+04	.517e+04	.695e+04
841	.679e-01	.393e+05	.811e+04	.204e+04	.523e+04	.703e+04
842	.674e-01	.392e+05	.809e+04	.203e+04	.522e+04	.701e+04
843	.677e-01	.393e+05	.810e+04	.203e+04	.522e+04	.702e+04
844	.677e-01	.392e+05	.810e+04	.202e+04	.522e+04	.702e+04
845	.650e-01	.389e+05	.801e+04	.202e+04	.517e+04	.694e+04
846	.679e-01	.393e+05	.811e+04	.204e+04	.523e+04	.703e+04
847	.674e-01	.392e+05	.809e+04	.203e+04	.522e+04	.701e+04
848	.676e-01	.393e+05	.810e+04	.203e+04	.522e+04	.702e+04
849	.677e-01	.392e+05	.810e+04	.202e+04	.522e+04	.702e+04
850	.650e-01	.389e+05	.801e+04	.202e+04	.517e+04	.694e+04
851	.679e-01	.393e+05	.811e+04	.204e+04	.523e+04	.703e+04
852	.674e-01	.392e+05	.809e+04	.203e+04	.522e+04	.701e+04
853	.676e-01	.393e+05	.810e+04	.203e+04	.522e+04	.702e+04
854	.677e-01	.392e+05	.810e+04	.202e+04	.522e+04	.702e+04
855	.650e-01	.389e+05	.801e+04	.202e+04	.517e+04	.694e+04
856	.679e-01	.393e+05	.811e+04	.204e+04	.523e+04	.703e+04
857	.674e-01	.392e+05	.809e+04	.203e+04	.522e+04	.701e+04
858	.676e-01	.393e+05	.810e+04	.203e+04	.522e+04	.702e+04
859	.677e-01	.392e+05	.810e+04	.202e+04	.522e+04	.702e+04
860	.650e-01	.389e+05	.801e+04	.202e+04	.517e+04	.694e+04
861	.679e-01	.393e+05	.811e+04	.204e+04	.523e+04	.703e+04
862	.674e-01	.392e+05	.809e+04	.203e+04	.522e+04	.701e+04
863	.676e-01	.393e+05	.810e+04	.203e+04	.522e+04	.702e+04
864	.677e-01	.392e+05	.810e+04	.202e+04	.522e+04	.702e+04
865	.650e-01	.389e+05	.801e+04	.202e+04	.517e+04	.694e+04
866	.679e-01	.393e+05	.811e+04	.204e+04	.523e+04	.703e+04
867	.674e-01	.392e+05	.809e+04	.203e+04	.522e+04	.701e+04
868	.676e-01	.393e+05	.810e+04	.203e+04	.522e+04	.702e+04
869	.677e-01	.392e+05	.810e+04	.202e+04	.522e+04	.702e+04
870	.650e-01	.389e+05	.801e+04	.202e+04	.517e+04	.694e+04
871	.679e-01	.393e+05	.811e+04	.204e+04	.523e+04	.703e+04
872	.674e-01	.392e+05	.809e+04	.203e+04	.522e+04	.701e+04
873	.676e-01	.393e+05	.810e+04	.203e+04	.522e+04	.702e+04
874	.677e-01	.392e+05	.810e+04	.202e+04	.522e+04	.702e+04
875	.650e-01	.389e+05	.801e+04	.202e+04	.517e+04	.694e+04
876	.679e-01	.393e+05	.811e+04	.204e+04	.523e+04	.703e+04
877	.674e-01	.392e+05	.809e+04	.202e+04	.522e+04	.701e+04
878	.676e-01	.393e+05	.810e+04	.203e+04	.522e+04	.702e+04
879	.677e-01	.392e+05	.810e+04	.203e+04	.522e+04	.702e+04
880	.650e-01	.389e+05	.801e+04	.202e+04	.517e+04	.694e+04
881	.679e-01	.393e+05	.811e+04	.204e+04	.523e+04	.703e+04
882	.674e-01	.392e+05	.809e+04	.202e+04	.522e+04	.701e+04
883	.676e-01	.393e+05	.810e+04	.203e+04	.522e+04	.702e+04
884	.677e-01	.392e+05	.810e+04	.202e+04	.522e+04	.702e+04
885	.649e-01	.389e+05	.801e+04	.202e+04	.517e+04	.694e+04
886	.679e-01	.393e+05	.811e+04	.204e+04	.523e+04	.703e+04

```
951   .679e-01   .393e+05   .811e+04   .204e+04   .523e+04   .703e+04
952   .674e-01   .392e+05   .809e+04   .203e+04   .521e+04   .701e+04
953   .676e-01   .393e+05   .810e+04   .203e+04   .522e+04   .702e+04
954   .676e-01   .393e+05   .810e+04   .203e+04   .522e+04   .732e+04
955   .649e-01   .389e+05   .801e+04   .202e+04   .516e+04   .694e+04
956   .679e-01   .393e+05   .811e+04   .204e+04   .523e+04   .703e+04
957   .674e-01   .392e+05   .809e+04   .203e+04   .521e+04   .701e+04
958   .676e-01   .393e+05   .810e+04   .203e+04   .522e+04   .702e+04
959   .676e-01   .393e+05   .810e+04   .203e+04   .522e+04   .702e+04
960   .649e-01   .389e+05   .801e+04   .202e+04   .516e+04   .694e+04
961   .679e-01   .393e+05   .811e+04   .204e+04   .523e+04   .703e+04
962   .674e-01   .392e+05   .809e+04   .203e+04   .521e+04   .701e+04
963   .676e-01   .393e+05   .810e+04   .203e+04   .522e+04   .702e+04
964   .676e-01   .393e+05   .810e+04   .203e+04   .522e+04   .702e+04
965   .649e-01   .389e+05   .801e+04   .202e+04   .516e+04   .694e+04
966   .679e-01   .393e+05   .811e+04   .204e+04   .523e+04   .703e+04
967   .673e-01   .392e+05   .809e+04   .203e+04   .521e+04   .701e+04
968   .676e-01   .393e+05   .810e+04   .203e+04   .522e+04   .702e+04
969   .676e-01   .393e+05   .810e+04   .203e+04   .522e+04   .702e+04
970   .649e-01   .389e+05   .801e+04   .202e+04   .516e+04   .694e+04
971   .679e-01   .393e+05   .811e+04   .204e+04   .523e+04   .703e+04
972   .673e-01   .392e+05   .809e+04   .203e+04   .521e+04   .701e+04
973   .676e-01   .393e+05   .810e+04   .203e+04   .522e+04   .702e+04
974   .676e-01   .393e+05   .810e+04   .203e+04   .522e+04   .702e+04
975   .649e-01   .389e+05   .801e+04   .202e+04   .516e+04   .694e+04
976   .679e-01   .393e+05   .811e+04   .204e+04   .523e+04   .703e+04
977   .672e-01   .392e+05   .809e+04   .203e+04   .521e+04   .701e+04
978   .676e-01   .393e+05   .810e+04   .203e+04   .522e+04   .702e+04
979   .676e-01   .393e+05   .810e+04   .203e+04   .522e+04   .702e+04
980   .649e-01   .389e+05   .801e+04   .202e+04   .516e+04   .694e+04
981   .679e-01   .393e+05   .811e+04   .204e+04   .523e+04   .703e+04
982   .673e-01   .392e+05   .809e+04   .203e+04   .521e+04   .701e+04
983   .676e-01   .393e+05   .810e+04   .203e+04   .522e+04   .702e+04
984   .676e-01   .393e+05   .810e+04   .203e+04   .522e+04   .702e+04
985   .649e-01   .389e+05   .801e+04   .202e+04   .516e+04   .694e+04
986   .678e-01   .393e+05   .811e+04   .204e+04   .523e+04   .703e+04
987   .673e-01   .392e+05   .809e+04   .203e+04   .521e+04   .701e+04
988   .676e-01   .393e+05   .810e+04   .203e+04   .522e+04   .702e+04
989   .676e-01   .393e+05   .810e+04   .203e+04   .522e+04   .702e+04
990   .649e-01   .389e+05   .801e+04   .202e+04   .516e+04   .694e+04
991   .678e-01   .393e+05   .811e+04   .204e+04   .523e+04   .703e+04
992   .673e-01   .393e+05   .809e+04   .203e+04   .521e+04   .701e+04
993   .675e-01   .393e+05   .810e+04   .203e+04   .522e+04   .702e+04
994   .676e-01   .393e+05   .810e+04   .203e+04   .522e+04   .702e+04
995   .649e-01   .389e+05   .801e+04   .202e+04   .516e+04   .694e+04
996   .678e-01   .393e+05   .811e+04   .204e+04   .523e+04   .703e+04
997   .673e-0    .393e+05   .809e+04   .203e+04   .521e+04   .701e+04
998   .675e-01   .393e+05   .810e+04   .203e+04   .522e+04   .702e+04
999   .648e-01   .389e+05   .801e+04   .202e+04   .516e+04   .694e+04
1000  .649e-01   .390e+05   .801e+04   .204e+04   .517e+04   .694e+04
```

```
DET        OBS.  SIGNAL          CALC. SIGNAL

 1           .3920e+05              .3896e+05
 2           .5300e+04              .8013e+04
 3           .2051e+04              .2038e+04
 4           .6688e+04              .5169e+04
 5           .6899e+04              .6945e+04


EE(J) (he         SPEC(J) (Watts/h.        TSPEC(J) (Watts)
```

887	.674e-01	.392e+05	.809e+04	.203e+04	.522e+04	.701e+04
888	.676e-01	.393e+05	.810e+04	.203e+04	.522e+04	.702e+04
889	.677e-01	.392e+05	.810e+04	.203e+04	.522e+04	.702e+04
890	.649e-01	.389e+05	.801e+04	.202e+04	.517e+04	.694e+04
891	.679e-01	.393e+05	.811e+04	.204e+04	.523e+04	.703e+04
892	.674e-01	.392e+05	.809e+04	.203e+04	.522e+04	.701e+04
893	.676e-01	.393e+05	.810e+04	.203e+04	.522e+04	.702e+04
894	.677e-01	.392e+05	.810e+04	.203e+04	.522e+04	.702e+04
895	.649e-01	.389e+05	.801e+04	.202e+04	.517e+04	.694e+04
896	.679e-01	.393e+05	.811e+04	.204e+04	.523e+04	.703e+04
897	.674e-01	.392e+05	.809e+04	.203e+04	.522e+04	.701e+04
898	.676e-01	.393e+05	.810e+04	.203e+04	.522e+04	.702e+04
899	.677e-01	.393e+05	.810e+04	.203e+04	.522e+04	.702e+04
900	.649e-01	.389e+05	.801e+04	.202e+04	.516e+04	.694e+04
901	.679e-01	.393e+05	.811e+04	.204e+04	.523e+04	.703e+04
902	.674e-01	.392e+05	.809e+04	.203e+04	.522e+04	.701e+04
903	.676e-01	.393e+05	.810e+04	.203e+04	.522e+04	.702e+04
904	.677e-01	.393e+05	.810e+04	.203e+04	.522e+04	.702e+04
905	.649e-01	.389e+05	.801e+04	.202e+04	.516e+04	.694e+04
906	.679e-01	.393e+05	.811e+04	.204e+04	.523e+04	.703e+04
907	.674e-01	.392e+05	.809e+04	.203e+04	.522e+04	.701e+04
908	.676e-01	.393e+05	.810e+04	.203e+04	.522e+04	.702e+04
909	.677e-01	.393e+05	.810e+04	.203e+04	.522e+04	.702e+04
910	.649e-01	.389e+05	.801e+04	.202e+04	.516e+04	.694e+04
911	.679e-01	.393e+05	.811e+04	.204e+04	.523e+04	.703e+04
912	.674e-01	.392e+05	.809e+04	.203e+04	.522e+04	.701e+04
913	.676e-01	.393e+05	.810e+04	.203e+04	.522e+04	.702e+04
914	.677e-01	.393e+05	.810e+04	.203e+04	.522e+04	.702e+04
915	.649e-01	.389e+05	.801e+04	.202e+04	.516e+04	.694e+04
916	.679e-01	.393e+05	.811e+04	.204e+04	.523e+04	.703e+04
917	.674e-01	.392e+05	.809e+04	.203e+04	.522e+04	.701e+04
918	.676e-01	.393e+05	.810e+04	.203e+04	.522e+04	.702e+04
919	.677e-01	.393e+05	.810e+04	.203e+04	.522e+04	.702e+04
920	.649e-01	.389e+05	.801e+04	.202e+04	.516e+04	.694e+04
921	.679e-01	.392e+05	.811e+04	.204e+04	.523e+04	.703e+04
922	.674e-01	.392e+05	.809e+04	.203e+04	.522e+04	.701e+04
923	.676e-01	.393e+05	.810e+04	.203e+04	.522e+04	.702e+04
924	.677e-01	.393e+05	.810e+04	.203e+04	.522e+04	.702e+04
925	.649e-01	.389e+05	.801e+04	.202e+04	.516e+04	.694e+04
926	.679e-01	.393e+05	.811e+04	.204e+04	.523e+04	.703e+04
927	.674e-01	.392e+05	.809e+04	.203e+04	.522e+04	.701e+04
928	.676e-01	.393e+05	.810e+04	.203e+04	.522e+04	.702e+04
929	.677e-01	.393e+05	.810e+04	.203e+04	.522e+04	.702e+04
930	.649e-01	.389e+05	.801e+04	.202e+04	.516e+04	.694e+04
931	.679e-01	.393e+05	.811e+04	.204e+04	.523e+04	.703e+04
932	.674e-01	.392e+05	.809e+04	.203e+04	.522e+04	.701e+04
933	.676e-01	.393e+05	.810e+04	.203e+04	.522e+04	.702e+04
934	.676e-01	.393e-05	.810e+04	.203e+04	.522e+04	.702e+04
935	.649e-01	.389e+05	.801e+04	.202e+04	.516e+04	.694e+04
936	.679e-01	.393e+05	.811e+04	.204e+04	.523e+04	.703e+04
937	.674e-01	.392e+05	.809e+04	.203e+04	.522e+04	.701e+04
938	.676e-01	.393e+05	.810e+04	.203e+04	.522e+04	.702e+04
939	.676e-01	.393e+05	.810e+04	.203e+04	.522e+04	.702e+04
940	.649e-01	.389e+05	.801e+04	.202e+04	.516e+04	.694e+04
941	.679e-01	.392e+05	.811e+04	.204e+04	.523e+04	.703e+04
942	.674e-01	.392e-05	.809e+04	.203e+04	.521e+04	.701e+04
943	.676e-01	.393e+05	.810e+04	.203e+04	.522e+04	.702e+04
944	.676e-01	.393e-05	.810e+04	.203e+04	.522e+04	.702e+04
945	.649e-01	.389e+05	.801e+04	.202e+04	.516e+04	.694e+04
946	.679e-01	.393e+05	.811e+04	.204e+04	.523e+04	.703e+04
947	.674e-01	.392e-05	.809e+04	.203e+04	.521e+04	.701e+04
948	.676e-01	.393e+05	.810e+04	.203e+04	.522e+04	.702e+04
949	.676e-01	.393e+05	.810e+04	.203e+04	.522e+04	.702e+04
950	.649e-01	.389e+05	.801e+04	.202e+04	.516e+04	.694e+04

35.0000	.2546e+09	.3564e+09
36.4000	.9718e+08	.4978e+09
37.8560	.4885e+08	.5717e+09
39.3702	.2458e+08	.6105e+09
40.9450	.1237e+08	.6307e+09
42.5828	.6221e+07	.6413e+09
44.2862	.3129e+07	.6468e+09
46.0576	.1574e+07	.6497e+09
47.8999	.7916e+06	.6513e+09
49.8159	.3981e+06	.6521e+09
51.8085	.2001e+06	.6525e+09
53.8809	.1007e+06	.6527e+09
56.0361	.5068e+05	.6528e+09
58.2775	.2549e+05	.6529e+09
60.6086	.1290e+05	.6529e+09
63.0330	.6523e+04	.6529e+09
65.5543	.3298e+04	.6529e+09
68.1765	.1698e+04	.6529e+09
70.9035	.8745e+03	.6529e+09
73.7397	.4503e+03	.6529e+09
76.6892	.2400e+03	.6529e+09
79.7568	.1280e+03	.6529e+09
82.9471	.6826e+02	.6529e+09
86.2650	.3818e+02	.6529e+09
89.7156	.2132e+02	.6529e+09
93.3042	.1193e+02	.6529e+09
97.0363	.7075e+01	.6529e+09
100.9178	.4197e+01	.6529e+09
104.9545	.2490e+01	.6529e+09
109.1527	.1572e+01	.6529e+09
113.5188	.9925e+00	.6529e+09
118.0595	.6267e+00	.6529e+09
122.7819	.4214e+00	.6529e+09
127.6932	.2834e+00	.6529e+09
132.8009	.1906e+00	.6529e+09
138.1129	.1364e+00	.6529e+09
143.6374	.9755e-01	.6529e+09
149.3829	.6980e-01	.6529e+09
155.3582	.5296e-01	.6529e+09
161.5725	.4018e-01	.6529e+09
168.0354	.3049e-01	.6529e+09
174.7568	.2443e-01	.6529e+09
181.7471	.1958e-01	.6529e+09
189.0170	.1569e-01	.6529e+09
196.5776	.1321e-01	.6529e+09
204.4407	.1112e-01	.6529e+09
212.6184	.9365e-02	.6529e+09
221.1231	.8239e-02	.6529e+09
229.9680	.7249e-02	.6529e+09
239.1667	.6378e-02	.6529e+09
248.7334	.5833e-02	.6529e+09
258.6827	.5334e-02	.6529e+09
269.0300	.4878e-02	.6529e+09
279.7912	.4613e-02	.6529e+09
290.9828	.4363e-02	.6529e+09
302.6222	.4127e-02	.6529e+09
314.7271	.4018e-02	.6529e+09
327.3161	.3912e-02	.6529e+09
340.4088	.3809e-02	.6529e+09
354.0251	.3803e-02	.6529e+09
368.1861	.3798e-02	.6529e+09
382.9135	.3792e-02	.6529e+09
398.2300	.3866e-02	.6529e+09
414.1592	.3994e-02	.6529e+09

```
430.7256              .4099e-02              .6529e+09
447.9546              .4127e-02              .6529e+09
465.8727              .4149e-02              .6529e+09
484.5076              .4167e-02              .6529e+09
503.8879              .4072e-02              .6529e+09
524.0434              .3977e-02              .6529e+09
545.0051              .3880e-02              .6529e+09
566.8053              .3693e-02              .6529e+09
589.4775              .3496e-02              .6529e+09
613.0566              .3313e-02              .6529e+09
637.5788              .3183e-02              .6529e+09
663.0819              .3058e-02              .6529e+09
689.6052              .2938e-02              .6529e+09
717.1893              .2849e-02              .6529e+09
745.8769              .2763e-02              .6529e+09
775.7119              .2680e-02              .6529e+09
806.7404              .2618e-02              .6529e+09
839.0099              .2558e-02              .6529e+09
872.5703              .2500e-02              .6529e+09
907.4731              .2457e-02              .6529e+09
943.7720              .2415e-02              .6529e+09
981.5228              .2374e-02              .6529e+09
1020.7837             .2345e-02              .6529e+09
1061.6150             .2315e-02              .6529e+09
1104.0796             .2286e-02              .6529e+09
1148.2427             .2265e-02              .6529e+09
1194.1724             .2245e-02              .6529e+09
1241.9392             .2224e-02              .6529e+09
1291.6167             .2209e-02              .6529e+09
1343.2814             .2193e-02              .6529e+09
1397.0126             .2178e-02              .6529e+09
1452.8931             .2166e-02              .6529e+09
1511.0088             .2154e-02              .6529e+09
1571.4491             .2142e-02              .6529e+09
1634.3070             .2130e-02              .6529e+09
1699.6792             .2118e-02              .6529e+09
```

*** COMPLETION OF COMPUTATION ***

********************************* IFLAG=   708

APPENDIX D

APPENDIX D

Listing of Program PIN

This is a listing of the pre-processor program PIN. As listed, this program will execute on a VAX 11/780 under the UNIX operating system. By changing the OPEN statements in subroutines DATIN and DATOUT, the program will run under the VMS operating system.

```fortran
        PROGRAM PIN
C
CCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCC
C                                                                  C
C       This program generates data and outputs the data in a format   C
C       that the program DCON can use as input.                    C
C                                                                  C
CCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCC
C
C
C       Input data...
C
        CALL DATIN
C
C       Generate data for PROGRAM DCON...
C
        CALL DATGEN
C
C       Output data generated in a form suitable for input to
C       PROGRAM DCON...
C
        CAL! DATOUT
        STOP
        END
```

```fortran
      SUBROUTINE DATIN
C
CCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCC
C                                                                    C
C     This subroutine reads in the detector responses and other data C
C     needed for the deconvolution.                                  C
C                                                                    C
CCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCC
C
      COMMON / ALL /  NDET , KITMAX , E1 , RATIO ,
     1 NFILT(15) , MATNUM(15,15) , VO(15,201) , TRES(200) ,
     2 DEL(15,15) , SO , ISHOT ,
     3 NSPEC , T1 , T2 , T3 , T4 , CC , DD ,
     4 AA , BB , ND , NE , NP ,
     5 KITP , ERROR , ERMAX
      COMMON / SERIAL / IDNUM(8) , DETCAL(8) , ISNUM(8) ,
     1                  DSTCAL(8)
      OPEN(UNIT=2,FILE='test.inp',STATUS='OLD')
      REWIND 2
C
C
C     SHOT is shot number for the experiment the data was taken from...
C
      WRITE(6,4)
4     FORMAT(' Enter shot number')
      READ(2,5) ISHOT
5     FORMAT(I4)
C
C     Read in how many detectors were used...
C
      WRITE(6,9)
9     FORMAT(' How many detectors were used?')
      READ(2,10) NDET
10    FORMAT(I1)
C
C     Loop over the number of detectors
C
      DO 100 I = 1 , NDET
C
C     Read in detector serial number...
C
      WRITE(6,12) I
12    FORMAT(' Enter detector serial number for detector ',I2)
      READ(2,14) IDNUM(I)
14    FORMAT(I3)
C
C     Read in voltage from scope trace...
C
      READ(2,150) VO(I,1)
150   FORMAT(E16.8)
C
C     Read number of filters used for detector I...
C
      WRITE(6,11) I
11    FORMAT(' Enter number of filters used for detector ',I2)
      READ(2,10) NFILT(I)
      WRITE(6,*) NFILT(I)
C
C     Loop over the number of filters for detector I...
C
      DO 20 J = 1 , NFILT(I)
C
C     Read in material number and thickness in centimeters for
C     filter J of detector I...
```

```fortran
       WRITE(6,16) J , I
       READ(2,15) MATNUM(I,J)
       WRITE(6,17) J , I
       READ(2,18)  DEL(I,J)
15     FORMAT(I1)
16     FORMAT(' Enter material number of filter ',I2,
      1 ' of detector',I2)
17     FORMAT(' Enter thickness in cm of filter ',I2,
      1 ' of detector',I2)
18     FORMAT(E16.8)
20     CONTINUE
100    CONTINUE
C
C      Input initial spectrum,  iteration/convergence parameters
C
       WRITE(6,200)
200    FORMAT(' Flat/Two-Temperature exponential spectrum? (1/0)',
      1 ' (This is for DCON)')
       READ(2,201) IS
201    FORMAT(I1)
       SO = -1.0
       IF ( IS .EQ. 0 ) GOTO 210
       WRITE(6,202)
202    FORMAT(' Enter initial flat spectrum value:')
       READ(2,203) SO
203    FORMAT(E16.8)
       GOTO 220
210    CONTINUE
C
C      Two-Temp spectrum of the form:
C      Initial spectrum = AA * DEXP( T1 / ENERGY ) +
C                         BB * DEXP( T2 / ENERGY )
C      is to be used as the initial spectrum for DCON
C
       WRITE(6,211)
211    FORMAT(' Enter T1')
       READ(2,203) T1
       WRITE(6,212)
212    FORMAT(' Enter T2')
       READ(2,203) T2
       WRITE(6,213)
213    FORMAT(' Enter AA')
       READ(2,203) AA
       WRITE(6,214)
214    FORMAT(' Enter BB')
       READ(2,203) BB
220    CONTINUE
       WRITE(6,230)
230    FORMAT(' Smoothing? (1/0)')
       READ(2,201) NP
       WRITE(6,240)
240    FORMAT(' Enter max error ERMAX')
       READ(2,203) ERMAX
       WRITE(6,250)
250    FORMAT(' Enter max iterations KITMAX (9999 MAX)')
       READ(2,251) KITMAX
251    FORMAT(I4)
       WRITE(6,260)
260    FORMAT(' New trial spectrum or previous spectrum? (1/0)')
       READ(2,201) NSPEC
C
       WRITE(6,270)
270    FORMAT(' Enter number of energy groups to be used (100 max)')
```

```fortran
      READ(2,251) NE
      WRITE(6,280)
280   FORMAT(' Enter lower limit of the energy groups')
      READ(2,203) E1
      WRITE(6,290)
290   FORMAT(' Enter energy group ratio')
      READ(2,203) RATIO
      CLOSE(UNIT=2)
      RETURN
      END
```

```
      SUBROUTINE DATGEN
      COMMON / ALL /   NDET , KITMAX , E1 , RATIO ,
     1 NFILT(15) , MATNUM(15,15) , VO(15,201) , TRES(200) ,
     2 DEL(15,15) , SO , ISHOT ,
     3 NSPEC , T1 , T2 , T3 , T4 , CC , DD ,
     4 AA , BB , ND , NE , NP ,
     5 KITP , ERROR , ERMAX
C
C     Get detector area in square cm...
C
      PI = 4.0 * ATAN(1.0)
      DETA = PI * ( 2.54 * 1.125 ) ** 2
C
C     Generate detector responses in amps*4PI OMEGA*(1/50.)...
C
      DO 1 I = 1 , NDET
C
C     Vacuum chamber radius in cm...
C
      R = 14.2875
      DO 10 J = 1 , NFILT(I)
      R = R + DEL(I,J)
10    CONTINUE
C
C     Area of sphere for source at center and detector at surface...
C
      A = 4.0 * PI * R ** 2
C
C     Multiply voltage signal by sphere area and divide by detector
C     area to get total from source...
C
      VO(I,1) = VO(I,1) * A
      VO(I,1) = VO(I,1) / DETA
C
C     VO is now per unit steradian...
C
C     Termination resistance was 50 ohms...
C
      VO(I,1) = VO(I,1) / 50.0
1     CONTINUE
      TRES(1) = 0.0
      RETURN
      END
```

```
      SUBROUTINE DATOUT
C
CCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCC
C                                                                    C
C     This subroutine outputs the  detector responses and other data C
C     needed for the deconvolution.                                  C
C                                                                    C
CCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCCC
C
      COMMON / ALL /   NDET , KITMAX , E1 , RATIO ,
     1 NFILT(15) , MATNUM(15,15) , VO(15,201) , TRES(200) ,
     2 DEL(15,15) , SO , ISHOT ,
     3 NSPEC , T1 , T2 , T3 , T4 , CC , DD ,
     4 AA , BB , ND , NE , NP ,
     5 KITP , ERROR , ERMAX
      COMMON / SERIAL / IDNUM(8) , DETCAL(8) , ISNUM(8) ,
     1                  DSTCAL(8)
      CHARACTER*20 FILNAM
C
C     Get name of output file...
C
      WRITE(6,3)
3     FORMAT(' Enter name of output file:')
      READ(5,2) FILNAM
2     FORMAT(A20)
C
C     Open output file...
C
C     Unit 3 is the output file being used...
C
      OPEN(UNIT=3,FILE=FILNAM,STATUS='NEW')
C
C     SHOT is shot number for the experiment the data was taken from...
C
      WRITE(3,5) ISHOT
5     FORMAT(I4)
C
C     Write out how many detectors were used...
C
      WRITE(3,10) NDET
10    FORMAT(I2)
C
C     Loop over the number of detectors
C
      DO 100 I = 1 , NDET
C
C     Write out detector serial number...
C
      WRITE(3,123) IDNUM(I)
123   FORMAT(I3)
C
C     Write number of filters used for detector I...
C
      WRITE(3,10) NFILT(I)
C
C     Loop over the number of filters for detector I...
C
      DO 20 J = 1 , NFILT(I)
C
C     Write out material number and thickness in centimeters for
C     filter J of detector I...
C
      WRITE(3,15) MATNUM(I,J) , DEL(I,J)
15    FORMAT(I1,E16.8)
```

111

```
20      CONTINUE
100     CONTINUE
C
C       Output initial spectrum.  iteration/convergence parameters
C
        WRITE(3,200) SO , NP , ERMAX , KITMAX , NSPEC , T1 , T2
C
C       SO is value for initial flat spectrum.  if < 0 then
C       two temperature DEXPonential is used instead.
C       NP is
C       ERMAX is average error allowed per detector for the
C       completion of the convergence
C       KITMAX is the maximum number of iterations allowed
C       T1 and T2 are the temperatures to be used for the
C       two temperature DEXPonential initial iteration
C
200     FORMAT( E16.8 , I5 , E16.8 , 2I5 , 2E16.8 )
C
C       Output energy limits/number of groups/etc...
C
        WRITE(3,220) NE , E1 , RATIO , AA , BB
220     FORMAT(I5,4E16.8)
C
C       Output initial detector signals and times...
C
        NTIMES = 1
        DO 350 J = 1 , NTIMES
        WRITE(3,300) ( VO(I,J) , I = 1 , NDET ) , TRES(J)
300     FORMAT(9E16.4)
350     CONTINUE
        CLOSE(UNIT=3)
        RETURN
        END
```

APPENDIX E

APPENDIX E

Sample Input for Program PIN

This is a listing of sample input for program PIN. This input to PIN generated the sample input for program DCON in Appendix B. The explanations listed to the right of the input data have no effect upon the program. They should be included though, because they will aid the user in modifying the input.

```
0604              SHOT D6-4
5                 NUMBER OF DETECTORS
125               Serial number of detector 1
27.5              Voltage signal from detector 1
3                 DETECTOR 1 HAS 3 FILTERS
3                 FIRST FILTER IS Fe
0.9525            Fe is 0.9525 cm  (3/8")
5                 SECOND FILTER IS Air
365.552           Air is 365.552 cm = 140.375"
2                 THIRD FILTER IS Al
0.635             Al is 0.635 cm = 1/4"
077               Serial number of detector 2
3.9               Voltage signal from detector 2
3                 DETECTOR 2 HAS 3 FILTERS
3                 FIRST FILTER IS Fe
0.9525            Fe is 0.9525 cm  (3/8")
5                 SECOND FILTER IS Air
356.87            Air is 356.87 cm = 140.5"
4                 THIRD FILTER IS Cu
0.3175            Cu is 0.3175 (1/8")
122               Serial number of detector 3
1.5               Voltage signal from detector 3
3                 DETECTOR 3 HAS 3 FILTERS
3                 FIRST FILTER IS Fe
0.9525            Fe is 0.9525 cm  (3/8")
5                 SECOND FILTER IS Air
358.14            Air is 358.14 cm (141")
1                 THIRD FILTER IS Pb
0.15875           Pb is 0.15875 cm (1/16")
123               Serial number of detector 4
5.0               Voltage signal from detector 4
4                 DETECTOR 4 HAS 4 FILTERS
3                 FIRST FILTER IS Fe
0.9525            Fe is 0.9525 cm  (3/8")
5                 SECOND FILTER IS Air
358.775           Air is 358.775 cm (141.25")
4                 THIRD FILTER IS Cu
0.3175            Cu is 0.3175 (1/8")
2                 FOURTH FILTER IS Al
0.635             Al is 0.635 cm = 1/4"
124               Serial number of detector 5
5.0               Voltage signal from detector 5
3                 DETECTOR 5 HAS 3 FILTERS
3                 FIRST FILTER IS Fe
0.9525            Fe is 0.9525 cm  (3/8")
5                 SECOND FILTER IS Air
359.7275          Air is 359.7275 cm (141.625")
4                 THIRD FILTER IS Cu
0.3175            Cu is 0.3175 (1/8")
1                 Flat Spectrum
1.0               Value is 1.0
1                 Smoothing Chosen (because Pb is a filter)
1.0e-6            ERMAX
1000              FITMAX
1                 New Spectrum
100               100 Energy groups
35.0              E1
1.04              RATIO
```

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Vita

Captain Charles William Beason was born on 18 September 1955 in Dover, Delaware. He graduated from Kecoughtan High School in Hampton, Virginia in 1973 and attended Louisiana Tech University from which he received the degree of Bachelor of Science with a double major in Physics and Mathematics. Upon graduation he received a commission in the USAF through the ROTC program. He attended graduate classes at Louisiana Tech until entering active duty in October 1978. He served as a Computational Physics Research Officer at the Advanced Weapons Concepts Branch at the Air Force Weapons Laboratory until entering the Air Force Institute of Technology in September 1982.

Permanent address: P. O. Box 2015

Alexandria, La. 71301

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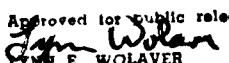
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> This investigation examined the feasibility of using an array of plastic scintillator photomultiplier tube radiation detectors to determine the high energy (h nu greater than 30keV) photon spectrum of SHIVA STAR. A method of determining the spectrum is outlined in which the detectors are filtered with different materials, and the spectrum is deconvoluted by an iterative technique on a computer. Data from SHIVA STAR was not available, so measurements of the r adiation output of a plasma gun were analyzed.

END

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